

SOUND WAVES

Before: water waves

• ideal fluid

• potential flow $\underline{v} = \nabla \phi$

implies $\nabla \cdot \underline{v} = 0$

Now we take into account compressibility

$$\nabla \cdot \underline{v} \neq 0$$

Euler's equation still holds:

$$\rho \left(\frac{\partial \underline{u}}{\partial t} + (\underline{u} \cdot \nabla) \underline{u} \right) = -\nabla p \quad (1)$$

Density $\rho = \rho(\underline{x}, t)$ is a variable.

Mass conservation

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \underline{u}) = 0 \quad (2)$$

$$\hookrightarrow \frac{\partial \rho}{\partial t} + \rho(\nabla \cdot \underline{u}) + \underline{u} \cdot \nabla \rho = 0$$

$$\frac{D\rho}{Dt} = -\rho(\nabla \cdot \underline{u})$$

We need the ρ vs p relationship
(thermodynamics)

Assume :
• ideal gas
• slow heat conduction in the fluid

Then $p \rho^{-\gamma} = \text{const}$ $\gamma = c_p / c_v = 1.4$
for a fluid element at usual T, p

$$\frac{D}{Dt} (p \rho^{-\gamma}) = 0 \quad (3)$$

Eqs (1) - (3) are the basis.

Small-amplitude sound waves

Undisturbed fluid (p_0, ρ_0) at rest $\underline{u} = 0$.

Trivial solution.

Consider a disturbance:

$$\underline{u} = \underline{u}_1, \quad p = p_0 + p_1, \quad \rho = \rho_0 + \rho_1$$

Perturbations are small, we aim to linearise.

$$p \rho^{-\gamma} = \text{const.} = p_0 \rho_0^{-\gamma}$$

↳ in the beginning

$$(p_0 + p_1)(\rho_0 + \rho_1)^{-\gamma} = p_0 \rho_0^{-\gamma}$$

$$\hookrightarrow \left(1 + \frac{p_1}{p_0}\right) \left(1 + \frac{\rho_1}{\rho_0}\right)^{-\gamma} = 1$$

$$\left(1 + \frac{p_1}{p_0}\right) \left(1 - \gamma \frac{\rho_1}{\rho_0} + \dots\right) = 1$$

\curvearrowright neglect.

we find:

$$\frac{p_1}{p_0} = \gamma \frac{\rho_1}{\rho_0}$$

really

$$c^2 = \left(\frac{\partial p}{\partial \rho}\right)_s$$

We define $c^2 = \frac{\gamma p_0}{\rho_0}$

$$\boxed{p_1 = c^2 \rho_1}$$

Euler's equation

$$(\rho_0 + \cancel{\rho_1}) \left(\frac{\partial \underline{u}_1}{\partial t} + (\underline{u}_1 \cdot \nabla) \underline{u}_1 \right) = -\nabla (\cancel{p_0} + \cancel{p_1})$$

\downarrow
 $\sim u_1^2$

$$\rho_0 \frac{\partial \underline{u}_1}{\partial t} = -\nabla p_1 \quad (\text{linearised Euler eq.})$$

Mass conservation

$$\frac{\partial \rho_1}{\partial t} + \rho_0 \nabla \cdot \underline{u}_1 = 0 \quad (\text{linearised})$$

Take $\nabla \cdot$ (Euler eq.)

$$\nabla \cdot \left(\rho_0 \frac{\partial \underline{u}_1}{\partial t} \right) = \nabla \cdot (-\nabla p_1)$$

↳

$$\rho_0 \frac{\partial}{\partial t} (\nabla \cdot \underline{u}_1) = -\nabla^2 p_1$$

$$\left(\nabla \cdot \underline{u}_1 = \frac{1}{\rho_0} \frac{\partial \rho_1}{\partial t} \quad (\text{M.C.}) \right)$$

$$\left(p_1 = c^2 \rho_1 \right)$$

$$\boxed{\frac{\partial^2 p_1}{\partial t^2} = c^2 \nabla^2 p_1}$$

3D Wave equation
for the pressure

For 1D waves

$$\frac{\partial^2 p_1}{\partial t^2} = c^2 \frac{\partial^2 p_1}{\partial x^2}$$

general solution:

$$p_1 = f(x-ct) + g(x+ct)$$

waves propagating without a change of shape.

• Sound waves are non-dispersive. $c = \text{const.}$

$$c = \sqrt{\frac{\gamma P_0}{\rho_0}} \approx 340 \frac{\text{m}}{\text{s}} \quad \text{at } p = p_0, T = 20^\circ\text{C}$$

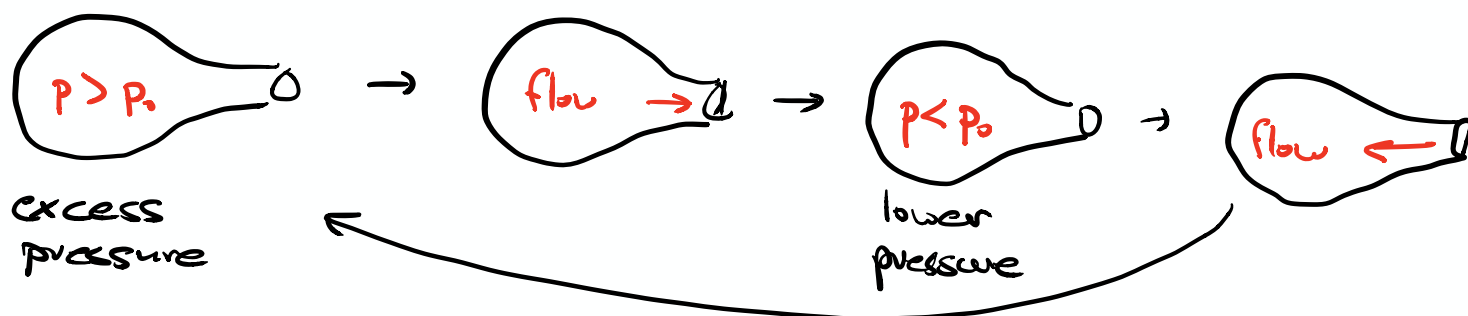
• In spherical geometry

$$p_1 = \frac{1}{r} F(r-ct)$$

Resonance and cavities

(Helmholtz resonator)

Consider a bottle or a cavity with a thin neck



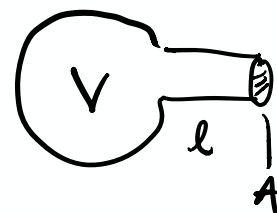
Fully reversible process (lack of dissipation & inertia)
with a period T .

→ Viscous attenuation of sound waves.

How long does it take?

Parameters:

- * Δp excess pressure
- * v peak air velocity in the tube
- * p_0 atmospheric pressure
- * ρ air density
- * A area of the neck
- * l length of the neck
- * V volume of the cavity



Let's estimate the period.

① How long does it take for the air to speed up?

The force on the air in the neck

$$F = A \Delta p = A(p - p_0)$$

so the acceleration

$$a = \frac{F}{m} = \frac{\Delta p A}{\rho l A} = \frac{\Delta p}{\rho l}$$

so the velocity after a time t

$$v = \frac{\Delta p}{\rho l} t$$

so the volume change

$$\Delta V = A v t = \frac{A \Delta p}{\rho l} t^2$$

How much time does the air take for the pressure to fall to atmospheric p ?

$$\frac{\Delta p}{p_0} = \frac{\Delta V}{V}$$

$$t^2 = \frac{\rho l \Delta V}{A \Delta p} = \frac{\rho l \Delta V}{A p_0 \Delta V} V = \frac{\rho l}{A p_0} V$$

$$\text{but } c^2 = \gamma \frac{p_0}{\rho} ; \quad c^2 \approx \frac{p_0}{\rho}$$

$$t^2 = \frac{\rho l V}{A c^2} \quad \text{is } \frac{1}{4} \text{ of the resonance time}$$

Therefore

$$T = 4 \sqrt{\frac{2V}{Ac^2}}$$

is under-estimated
(we assumed $\Delta p = \text{const.}$
 $\alpha = \text{const.}$)

The actual answer is

$$T = 2\pi \sqrt{\frac{2V}{Ac^2}}$$

A: wider neck \rightarrow faster flow \rightarrow lower T

l: longer neck \rightarrow longer to get the air moving

V: larger volume: more air has to leave before p falls

c: faster speed \rightarrow quicker response

The general procedure for finding resonance frequencies in a cavity:

• assume sinusoidal changes of pressure

Basic linearised eqs.

$$\frac{\partial^2 p}{\partial t^2} = c^2 \nabla^2 p$$

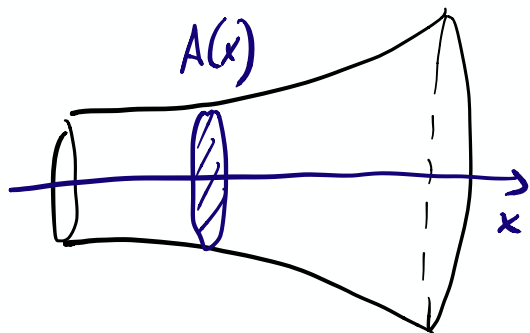
$$\left\{ \begin{array}{l} \frac{D \underline{u}}{Dt} = \frac{1}{\rho} \nabla p \\ \frac{D p}{Dt} = \gamma p_0 (\nabla \cdot \underline{u}) \end{array} \right.$$

look for solutions that satisfy

$$\nabla^2 p = -\frac{\omega^2}{c^2} p$$

with a b.c. of $p=0$ at the openings.

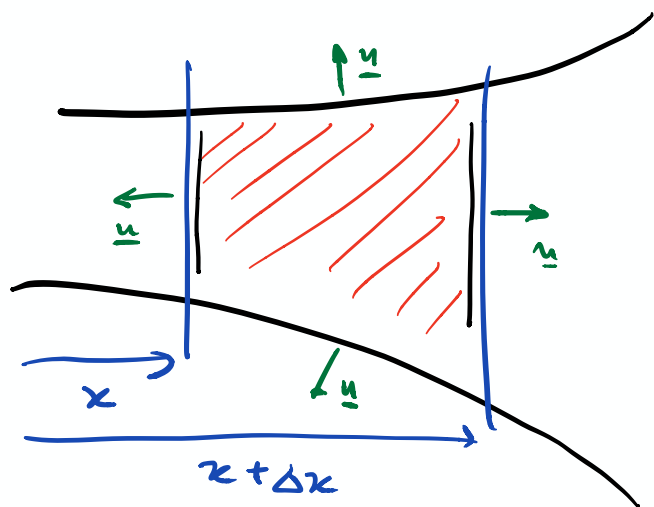
→ discrete set of resonance frequencies ω .




Webster Horn Equation

Quasi-1D model of sound propagation in a rigid-walled duct of variable cross-sectional area $A(x)$

Quasi 1D $p = p(x, t)$



wave equation for the acoustic pressure over the volume 

$$\int \frac{\partial^2}{\partial t^2} p \, dV = \int c^2 \frac{\partial^2}{\partial x^2} p \, dV$$

Gauss' theorem $\int dV (\nabla \cdot \underline{f}) = \int dS (\underline{f} \cdot \underline{n})$

$$\frac{\partial^2}{\partial t^2} \int p \, dV - c^2 \int \frac{\partial}{\partial x} \cdot \frac{\partial p}{\partial x} \, dV = \frac{\partial^2}{\partial t^2} \int dV p - c^2 \int dS (\nabla p \cdot \underline{n})$$

on the walls: $\nabla p \cdot \underline{n} = 0$

$$c^2 \left(\int ds \frac{\partial p}{\partial x} \Big|_{x+\Delta x} - \int ds \frac{\partial p}{\partial x} \Big|_x \right) = \frac{\partial^2}{\partial t^2} \int ds \int dx p$$

$$c^2 \Delta x \frac{\partial}{\partial x} \int ds \frac{\partial p}{\partial x} = \frac{\partial^2}{\partial t^2} \int ds p \Delta x$$

In the limit of $\Delta x \rightarrow 0$, we get

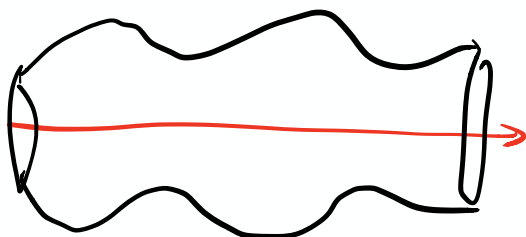
$$\frac{\partial}{\partial x} \int ds \frac{\partial p}{\partial x} - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \int ds p = 0$$

If p is uniform over a cross-section, we get the Webster horn equation:

$$\frac{1}{A(x)} \frac{\partial}{\partial x} \left(A(x) \frac{\partial p}{\partial x} \right) - \frac{1}{c^2} \frac{\partial^2 p}{\partial t^2} = 0$$

↳

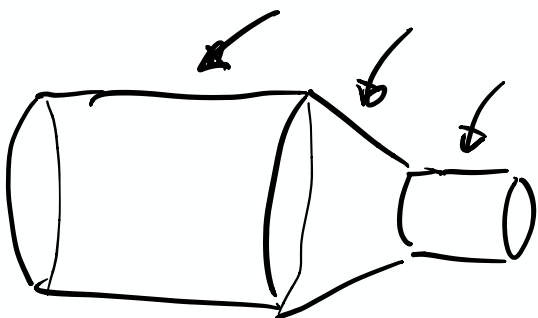
$$\left\{ \frac{\partial^2}{\partial x^2} + \frac{1}{A} \frac{\partial A}{\partial x} \frac{\partial}{\partial x} - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \right\} p = 0$$



↳ p

If the velocity is needed, we use the Euler's eq.

$$\rho \frac{\partial v_x}{\partial t} = - \frac{\partial p}{\partial x}$$



→ ω ?

$$\frac{1}{A} \frac{\partial A}{\partial x} = \alpha = \text{const.}$$

$$\frac{\partial A}{\partial x} = \alpha A$$

$$\left\{ \partial_x^2 + \partial_x - \frac{1}{2} \partial_t^2 \right\} \psi = 0$$

$$A = B e^{\alpha x}$$

$e^{\alpha x}$

(Exponential growth)