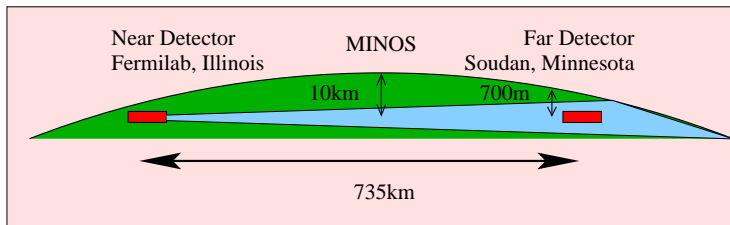


Neutrino oscillations in the MINOS experiment



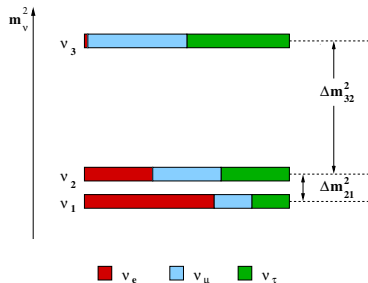
Introduction



- MINOS experiment: two magnetized (1.3 T), underground detectors located **on** NuMI beam axis
- NuMI: ν_μ or $\bar{\nu}_\mu$ enhanced beam
- Near Detector (ND) (1kt) at Fermilab, 1km from target, $\sim 100\text{m}$ underground: **study of neutrino interactions with absence of oscillations**
- Far Detector (FD) (5.4 kt), 735km away, at Soudan (Minnesota), 730m underground: **study/search for neutrino oscillations**
- **Latest results based on data from 2005-2011**

MINOS Oscillation Measurements

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \mathbf{U} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$



MINOS is sensitive to the larger $\Delta m_{ij}^2 (= m_i^2 - m_j^2)$

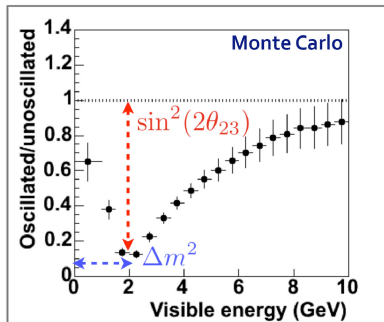
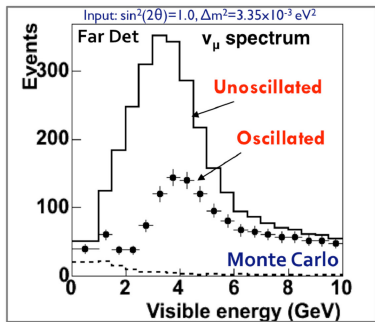
$$\mathbf{U} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\text{atmospheric, accelerator}} \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix}}_{\text{reactor, accelerator}} \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{solar, reactor}}$$

$c_{ij} \equiv \cos \theta_{ij}$, $s_{ij} \equiv \sin \theta_{ij}$, $\delta \equiv CP$ violation phase.

MINOS Oscillation Measurements

Measurement of energy-dependent ν_μ disappearance

$$P(\nu_\mu \rightarrow \nu_\mu) \simeq 1 - \sin^2(2\theta_{23}) \sin^2 \left(1.267 \Delta m^2 \frac{L}{E} \right)$$



In this talk latest (2010) results for statistics: 7.25×10^{20} pot

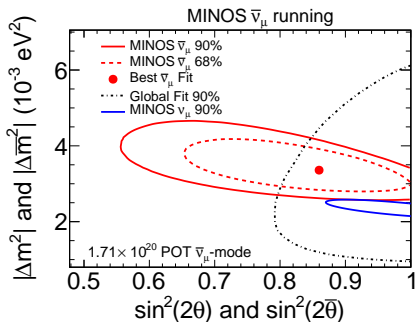
Measurement of energy-dependent $\bar{\nu}_\mu$ disappearance

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu) \simeq 1 - \sin^2 2\bar{\theta}_{23} \sin^2 \frac{1.27 \Delta \bar{m}^2 L}{E_{\bar{\nu}}}$$

- Antineutrino analysis very similar to the neutrino analysis.
- MINOS is unique: magnetic field allows for the event-by-event separation of ν_μ and $\bar{\nu}_\mu$ charged-current interactions.
- **Difference between neutrinos and antineutrinos ?**
- Two possible sources of $\bar{\nu}_\mu$:
 - 7% antineutrino component of neutrino beam
 - By reversing the horn current, NuMI can create $\bar{\nu}_\mu$ -enhanced beam

Measurement of energy-dependent $\bar{\nu}_\mu$ disappearance

MINOS results in 2010



- In 2010, MINOS presented ν_μ and $\bar{\nu}_\mu$ oscillation parameters measurements (for 1.71×10^{20} pot) that are consistent with each other at the 2% level.
- In this talk new 2011 results for increased statistics: 2.95×10^{20} pot

MINOS Oscillation Measurements

Search for ν_e appearance in the ν_μ beam

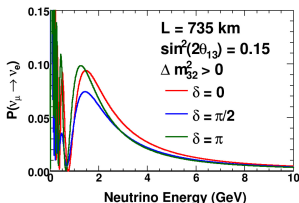
$$P(\nu_\mu \rightarrow \nu_e) \simeq \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \Delta$$

$$\mp \alpha \sin 2\theta_{13} \sin \delta_{CP} \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \sin^3 \Delta$$

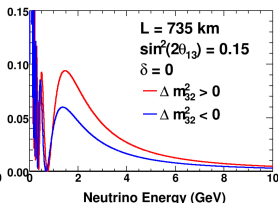
$$+ \alpha \sin 2\theta_{13} \cos \delta_{CP} \cos \theta_{13} \sin 2\theta_{12} \sin 2\theta_{23} \cos \Delta \sin^2 \Delta$$

$$+ \alpha^2 \cos^2 \theta_{23} \sin^2 2\theta_{12} \sin^2 \Delta$$

δ dependence



mass hierarchy dependence



$$\Delta \equiv \frac{\Delta m_{atm}^2 L}{4E_\nu}$$

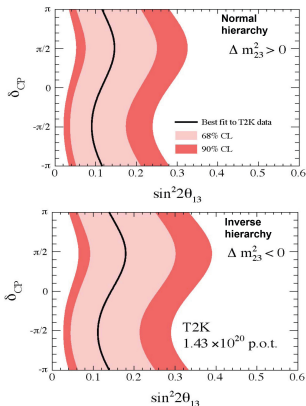
$$\alpha \equiv \frac{\Delta m_{sol}^2}{\Delta m_{atm}^2}$$

For ν_e probabilities are modified by matter effects.

MINOS Oscillation Measurements

Search for ν_e appearance in the ν_μ beam

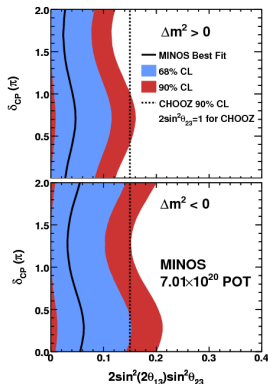
T2K 2011 result



T2K observed 6 events with expected background 1.5 ± 0.3 events for $\theta_{13} = 0$

K.Grzelak (University of Warsaw)

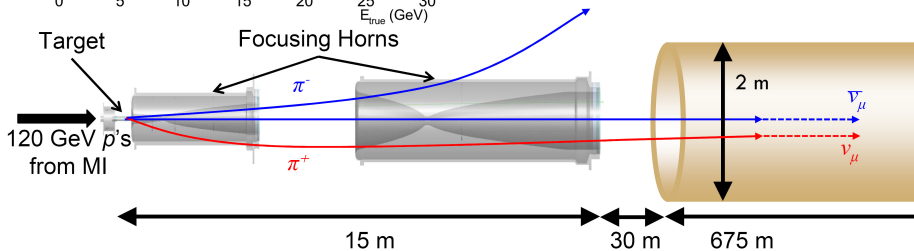
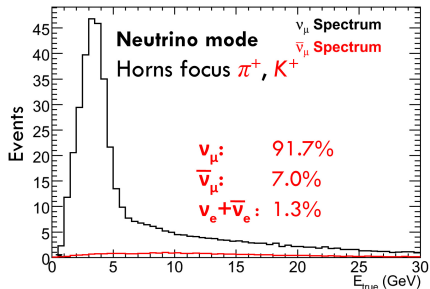
MINOS 2010 result



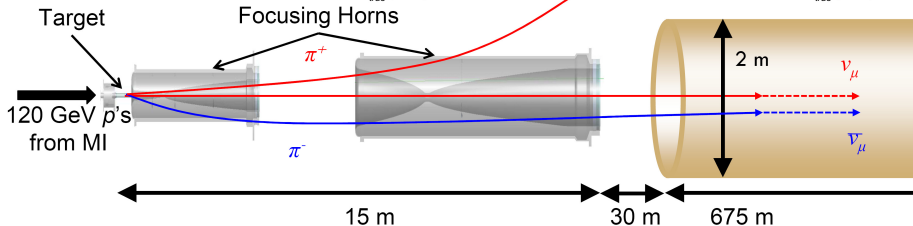
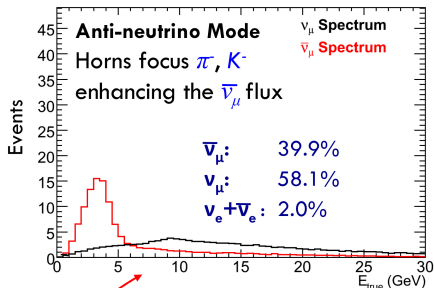
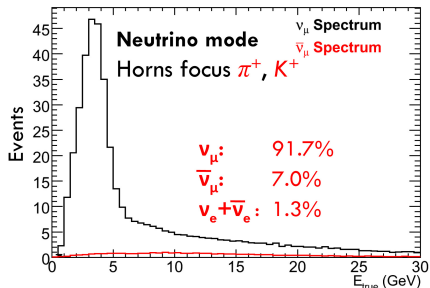
In this talk new 2011 results for increased statistics: 8.2×10^{20} pot

NuMI beam and MINOS detectors

Neutrino Mode



Antineutrino Mode

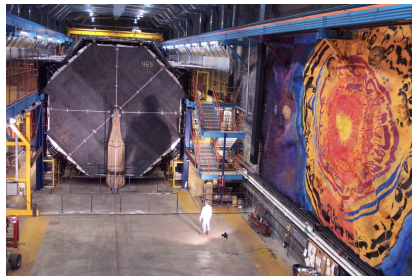


Minos Detectors

Near Detector (ND)



Far Detector (FD)



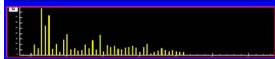
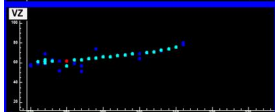
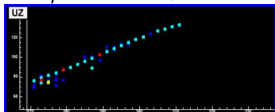
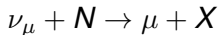
scintillator planes
segmented into 4.1cm
(~ 1.1 Molière radii)
wide strips

- steel-scintillator sandwich
- each plane is 2.54cm of steel and 1cm of solid scintillator strips (~ 1.4 radiation length)
- FD: 484 planes; $8\text{m} \times 8\text{m} \times 30\text{m}$
- ND: 282 planes; $3.8\text{m} \times 4.8\text{m} \times 15\text{m}$
- magnetic field $B = 1.3\text{ T}$

Neutrino interactions in the MINOS detectors

ν_μ ($\bar{\nu}_\mu$) disappearance

charged current (CC)



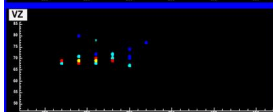
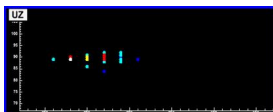
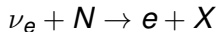
ν_μ -CC, $\bar{\nu}_\mu$ -CC

long, μ track

$$E_\nu = E_{shower} + E_\mu$$

ν_e appearance

charged current (CC)

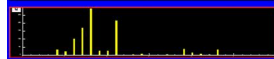
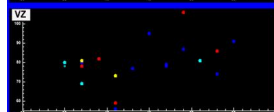
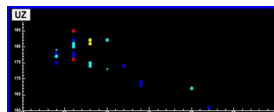
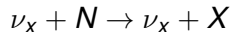


ν_e -CC

compact showers,

EM-profile

neutral current (NC)



NC

diffuse showers

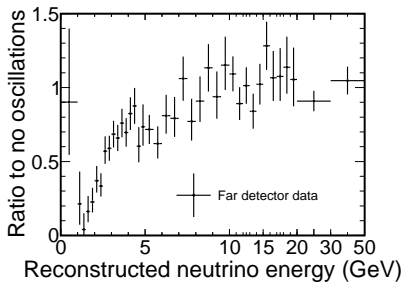
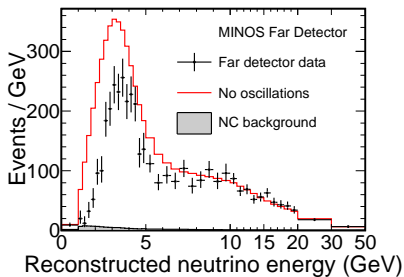
Two detectors (ND and FD) reduce systematics. Methodology:

- measure unoscillated energy spectrum at ND
- use Near Detector spectrum to predict FD spectrum
- compare to measured spectrum at FD

Blind to full FD sample until all analysis procedures are frozen.

ν_μ disappearance (7.25×10^{20} pot)

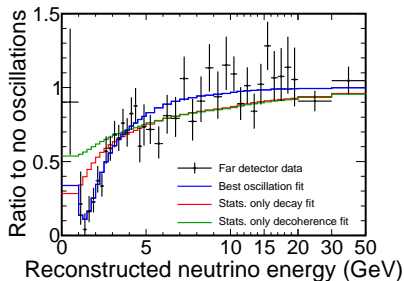
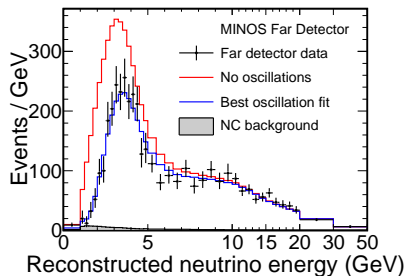
ν_μ disappearance, Far Detector Energy Spectrum



Predicted if no oscillations: **2451** events.

Observed: **1986** events

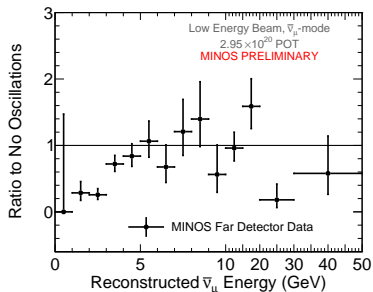
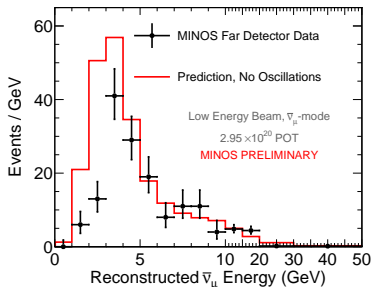
ν_μ disappearance, Far Detector Energy Spectrum



$$\Delta m^2 = (2.32_{-0.08}^{+0.12}) \times 10^{-3} \text{eV}^2$$
$$\sin^2 2\theta_{23} > 0.90 \text{ (90\% C.L.)}$$

$\bar{\nu}_\mu$ disappearance (2.95×10^{20} pot)

$\bar{\nu}_\mu$ disappearance: Far Detector Energy Spectrum

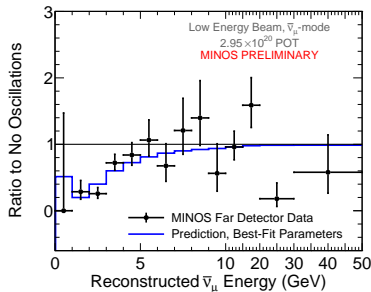
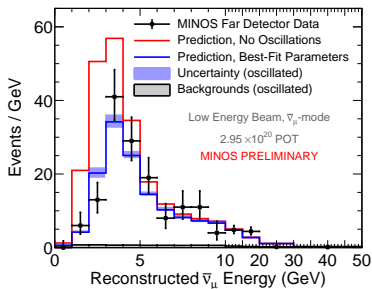


Predicted if no oscillations: **273** events.

Observed: **193** events

Null-oscillations excluded at 7.3σ

$\bar{\nu}_\mu$ disappearance: Far Detector Energy Spectrum



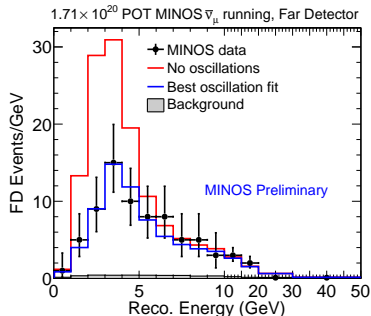
$\bar{\nu}_\mu$ Best Fit Parameters:

$$\Delta \bar{m}^2 = (2.62^{+0.31}_{-0.28}(\text{stat}) \pm 0.09(\text{syst})) \times 10^{-3} \text{eV}^2$$

$$\sin^2 2\bar{\theta}_{23} = 0.95^{+0.10}_{-0.11}(\text{stat}) \pm 0.01(\text{syst})$$

Comparison with previous result

2010 result



Predicted if no oscillations: **155** events.

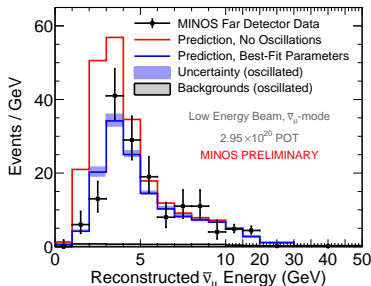
Observed: **97** events

$\Delta\bar{m}^2 =$

$(3.36^{+0.46}_{-0.40}(\text{stat}) \pm 0.06(\text{syst})) \times 10^{-3} \text{eV}^2$

$\sin^2 2\bar{\theta}_{23} = 0.86^{+0.11}_{-0.12}(\text{stat}) \pm 0.01(\text{syst})$

2011 result



Predicted if no oscillations: **273** events.

Observed: **193** events

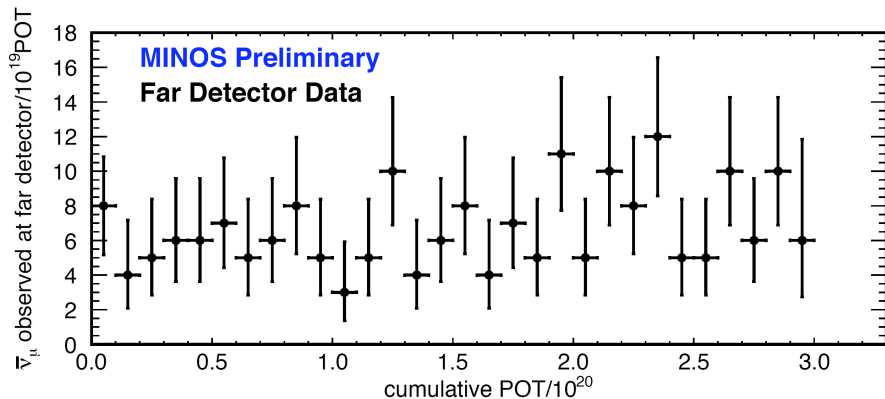
$\Delta\bar{m}^2 =$

$(2.62^{+0.31}_{-0.28}(\text{stat}) \pm 0.09(\text{syst})) \times 10^{-3} \text{eV}^2$

$\sin^2 2\bar{\theta}_{23} = 0.95^{+0.10}_{-0.11}(\text{stat}) \pm 0.01(\text{syst})$

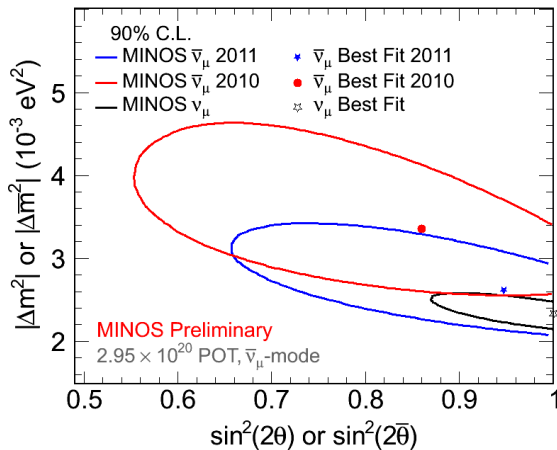
Comparison with previous result

Far Detector Data Rates



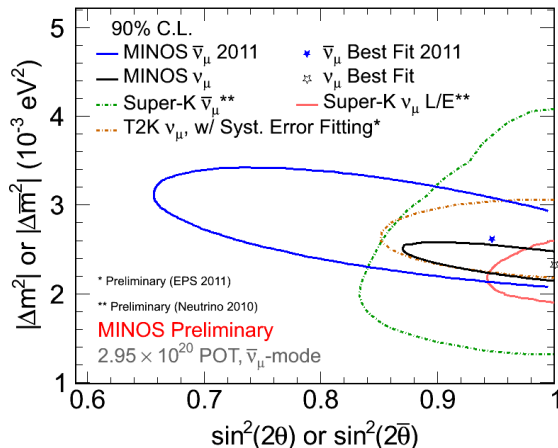
Comparison with previous results

$\bar{\nu}_\mu$ oscillation parameters



ν_μ and $\bar{\nu}_\mu$ oscillation parameters measurements are consistent with each other at the 42% C.L. (was 2% level).

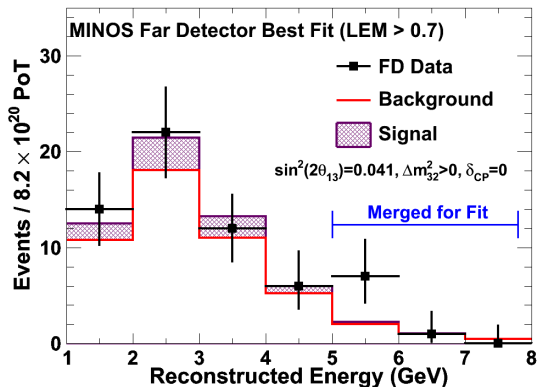
Contour Comparisons



Comparisons with Super-K and T2K results.

ν_e appearance (8.2×10^{20} pot)

ν_e appearance: Far Detector Energy Spectrum



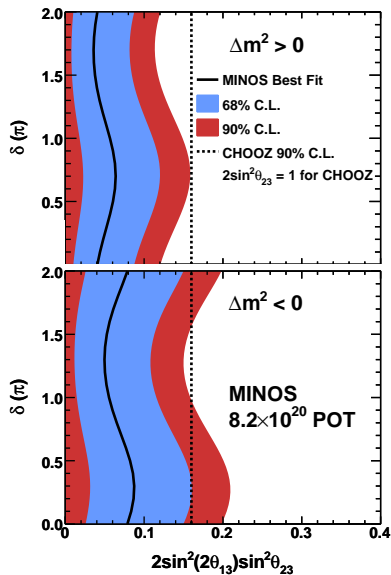
Predicted if $\theta_{13} = 0$:

$49.6 \pm 7.0(\text{stat.}) \pm 2.7(\text{syst.})$

events

In the signal region
observed: **62** events.

ν_e oscillation parameters

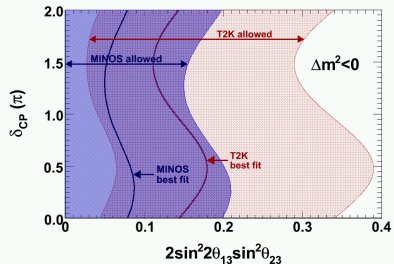
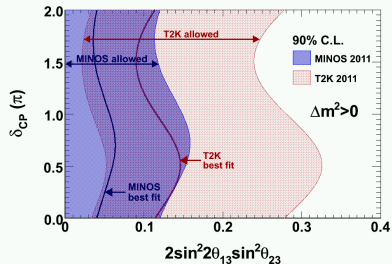


$2 \sin^2 \theta_{23} \sin^2 2\theta_{13} < 0.12$ at
90% C.L. for normal hierarchy

$2 \sin^2 \theta_{23} \sin^2 2\theta_{13} < 0.20$ at
90% C.L. for inverted hierarchy

$\theta_{13} = 0$ is disfavoured at 89% C.L.

Comparison to T2K result



MINOS result significantly constrains θ_{13} region allowed by T2K data.

- ν_e appearance analysis:
MINOS result significantly constrains θ_{13} region allowed by T2K data.
- ν_μ disappearance analysis:
The most precise measurement of larger $|\Delta m^2|$ to date
- $\bar{\nu}_\mu$ disappearance analysis:
**The most precise, direct measurement of larger $|\Delta \bar{m}^2|$.
 ν_μ and $\bar{\nu}_\mu$ oscillation parameters measurements are consistent with each other at the 42% C.L.**
- **Future:** MINOS+ = MINOS at NOvA beam from 2013