

# UNEXPECTED PHYSICS AT THE LHC

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Warsaw, 18 May 2009

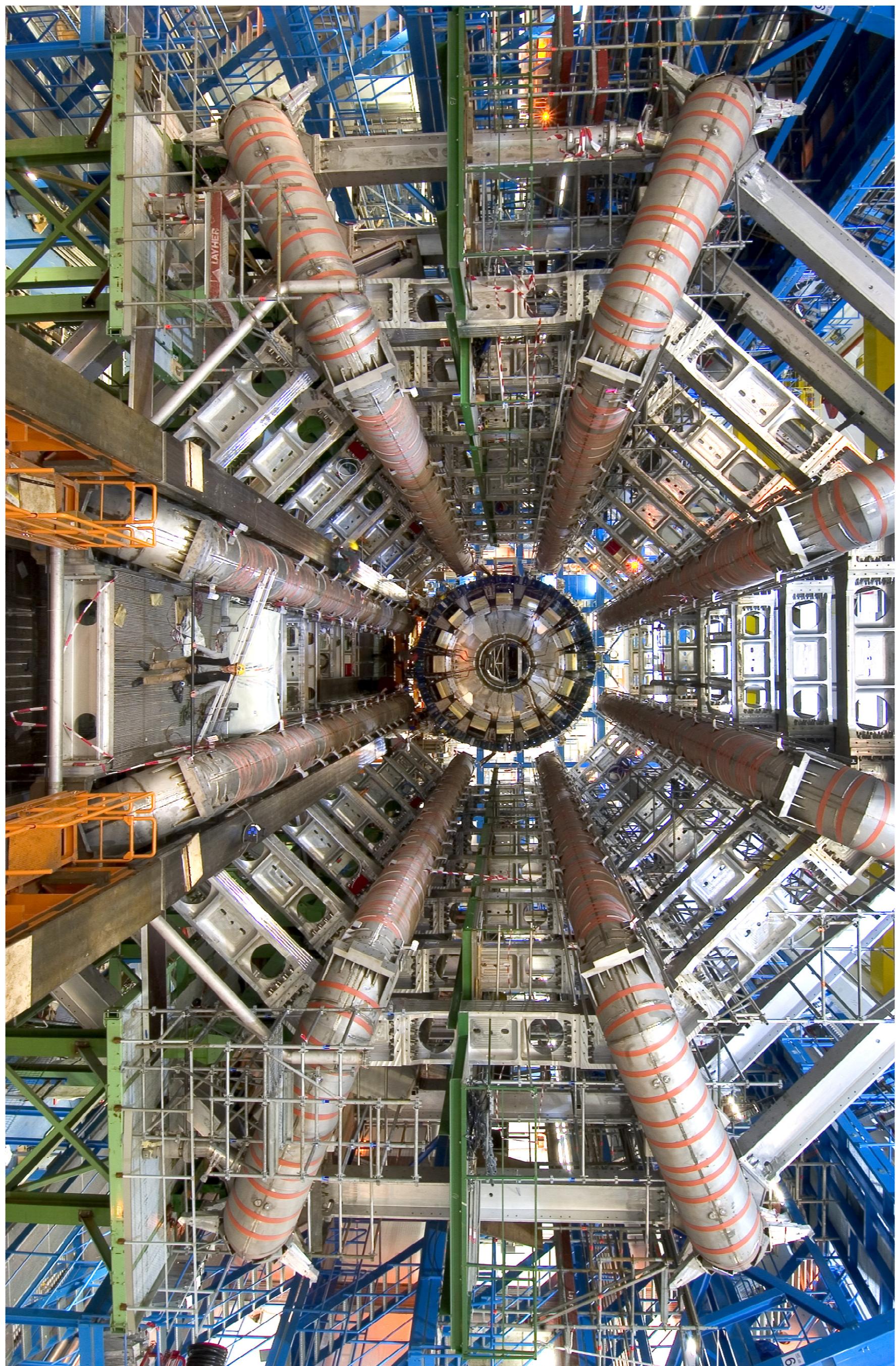
J.R. Espinosa  
ICREA @ IFAE, Barcelona  
& CERN

★ Most expected : SUSY

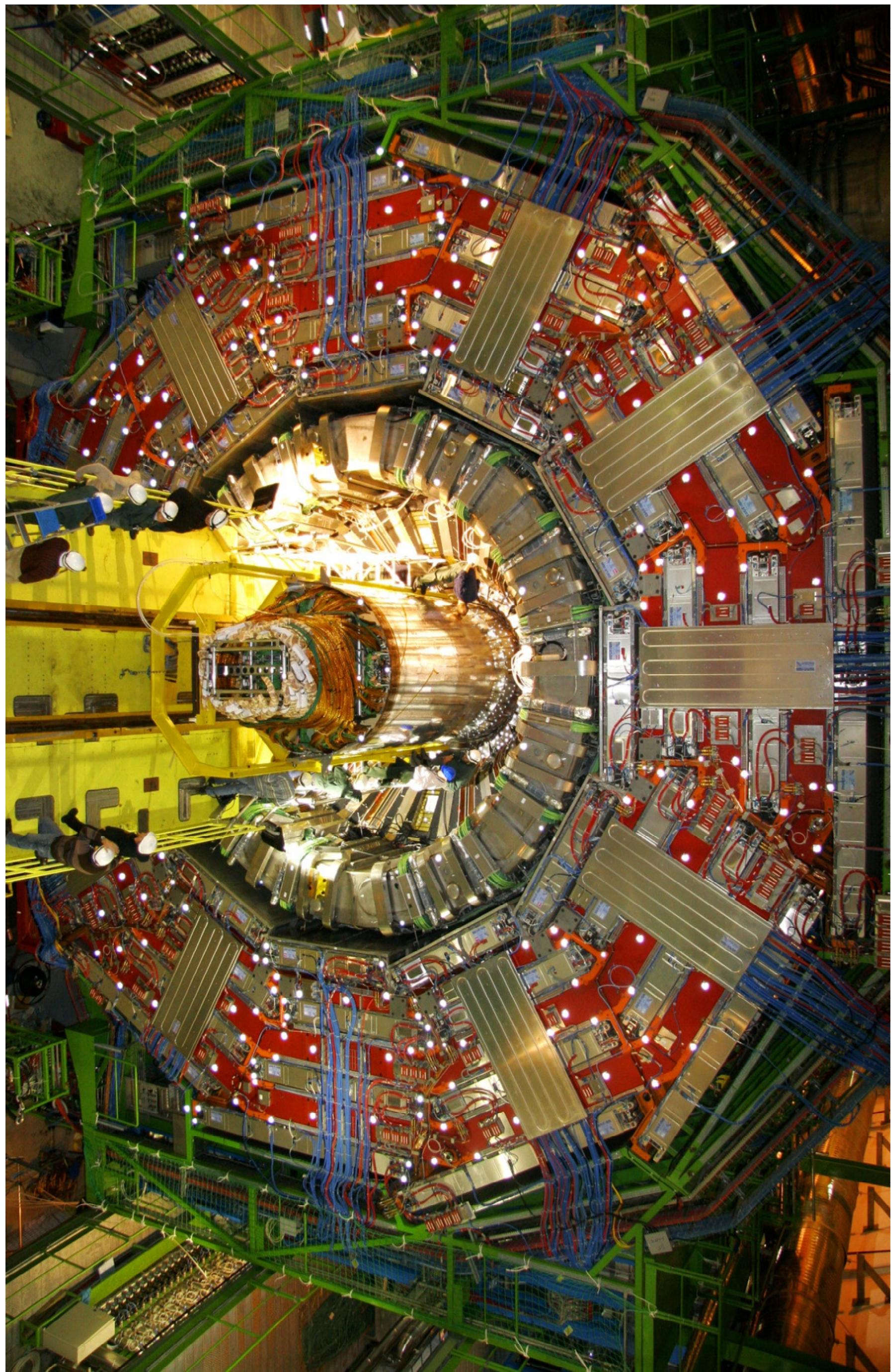
★ Quite unexpected : Unparticles

MUST BE READY FOR THE UNEXPECTED





We'll see here stops flying around

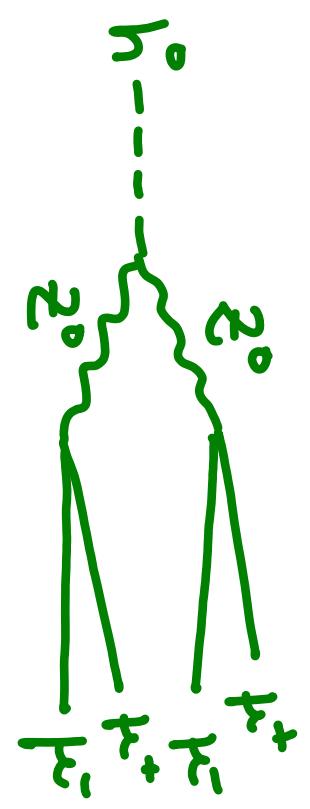
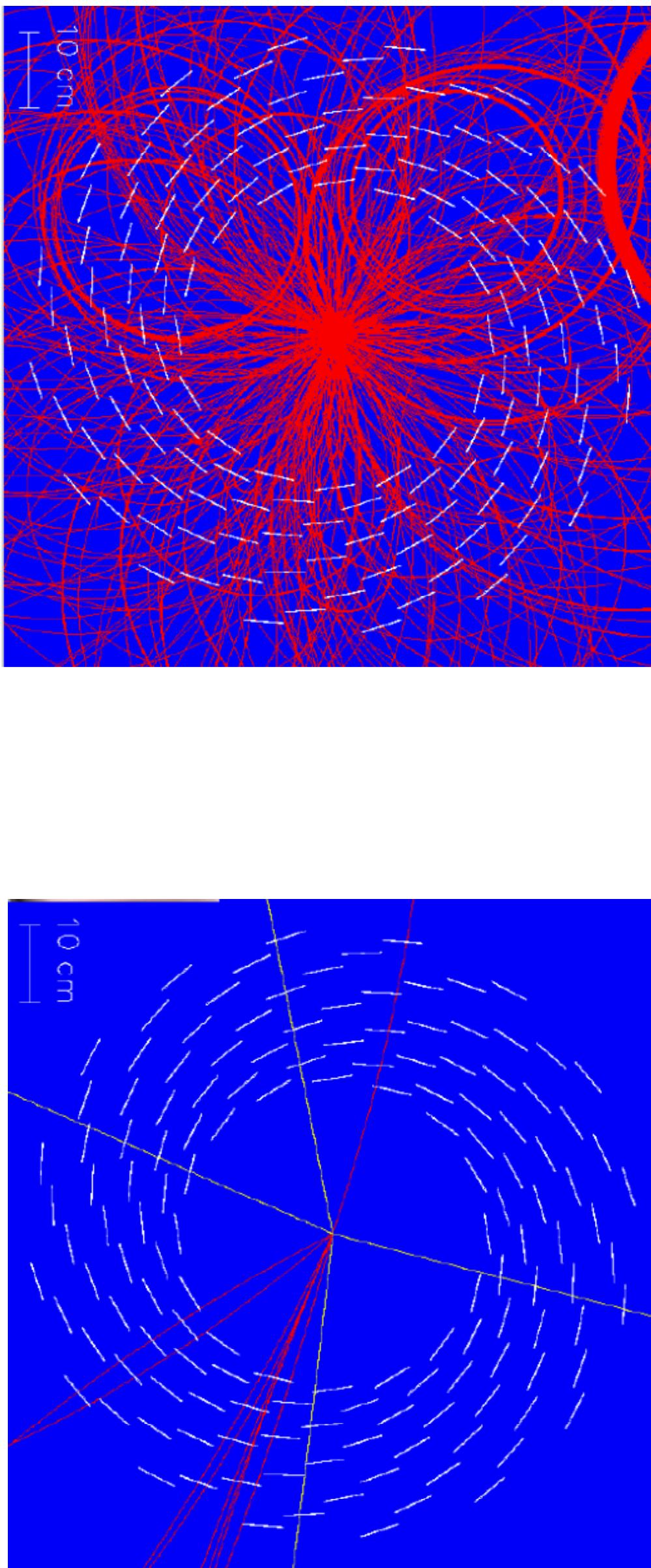


We'll see here gluinos flying around

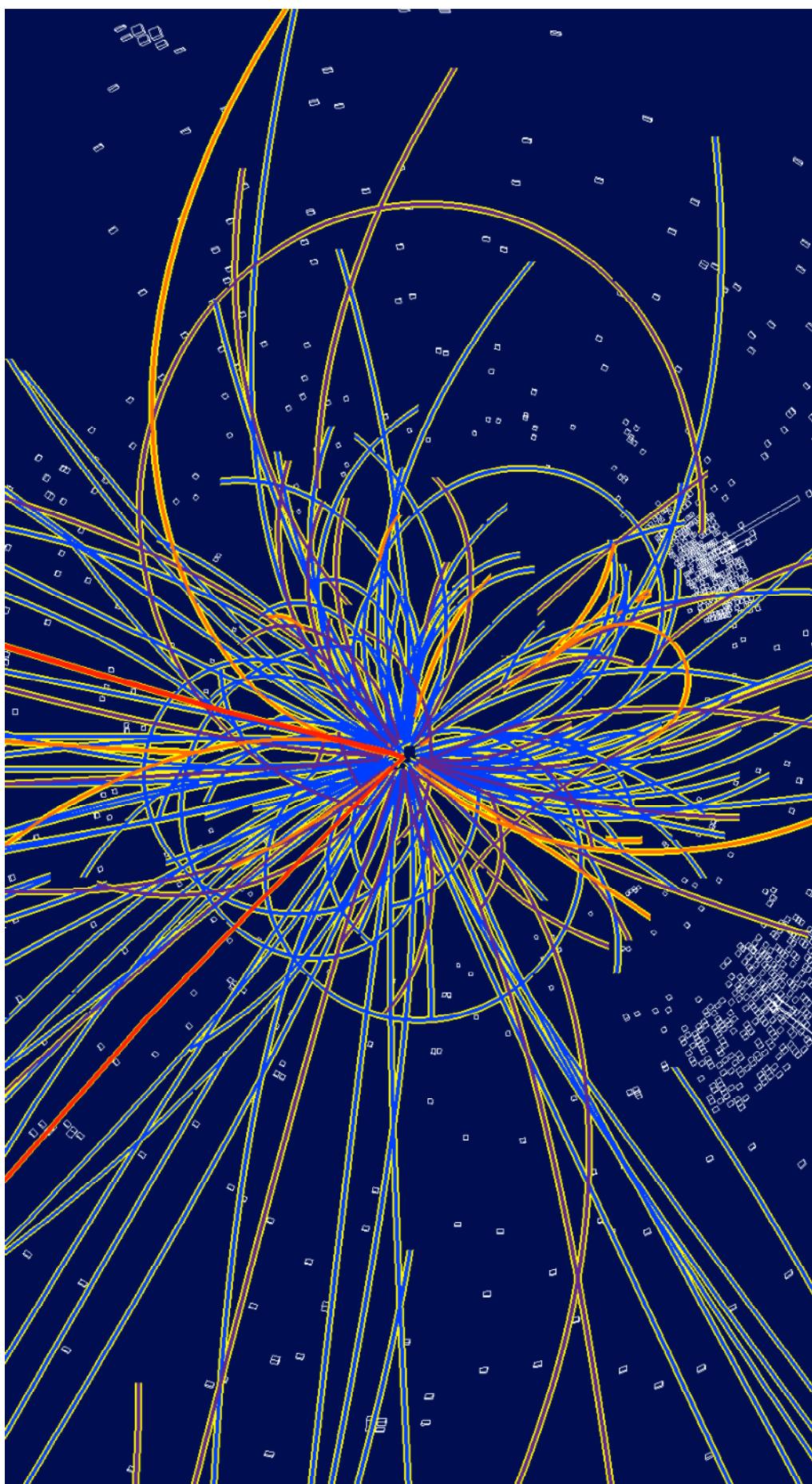
# FINDING A NEEDLE IN A MILLION HAYSTACKS

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Find 4 energetic muons  
(straight tracks) central tracker can do it!



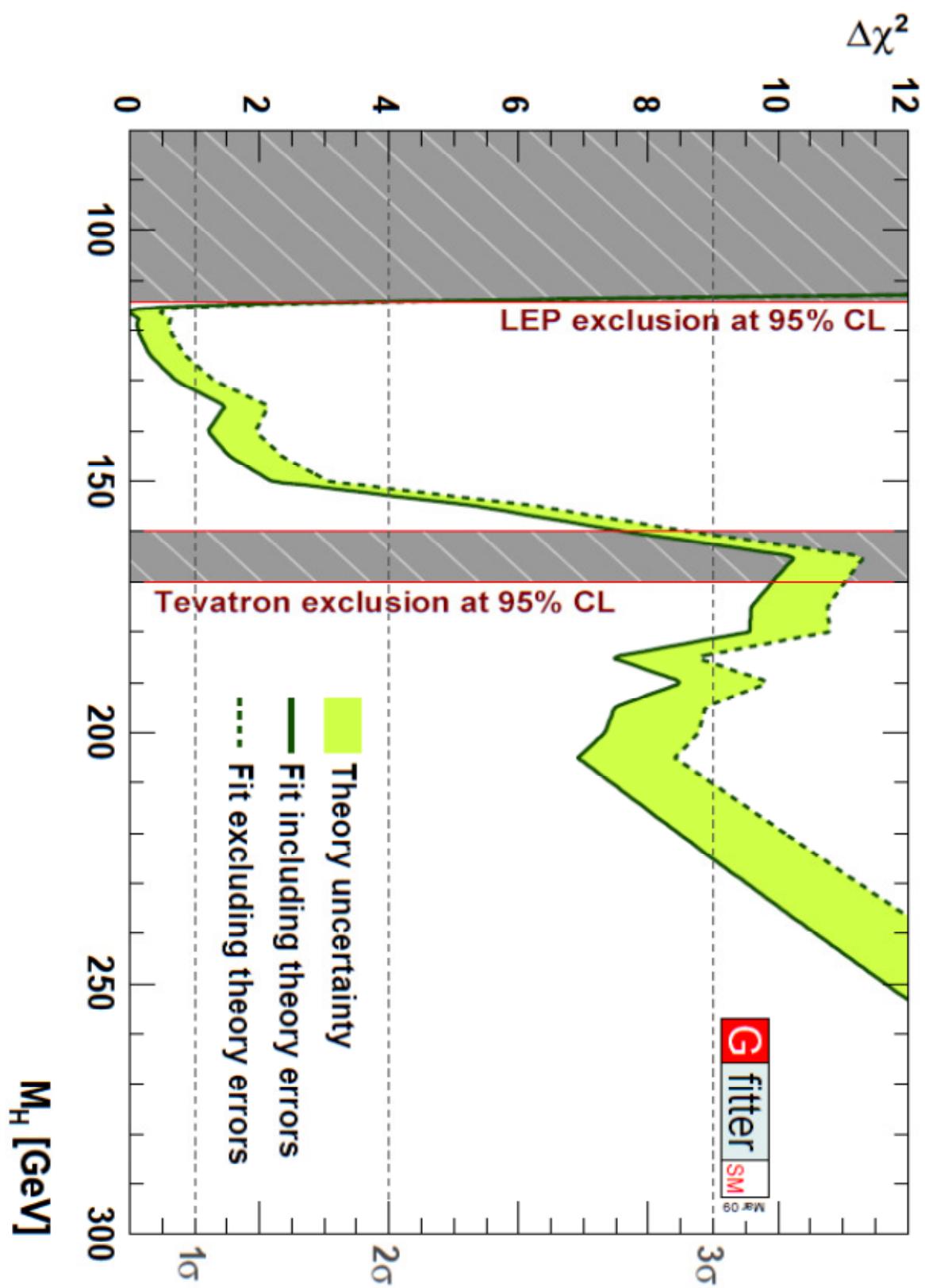
# FINDING A NEEDLE IN A HILLION HAYSTACKS



Maybe it's not a needle ...

# SM Higgs is PROBABLY light

$m_h < 153 \text{ GeV}$   
at 95% CL



# HIERARCHY PROBLEM

It is not natural to have  $M_{EW} \ll M_{Pl}$   
fixed by  $\delta_V = \frac{1}{2} m^2 h^2$

and  $m^2$  is very much UV sensitive. Quantum corrections

give

$$\delta m^2 \sim \Lambda^2$$

$m^2$  natural requires new physics beyond the SM at

$$\Lambda \sim \text{TeV} \rightarrow \text{LHC!}$$

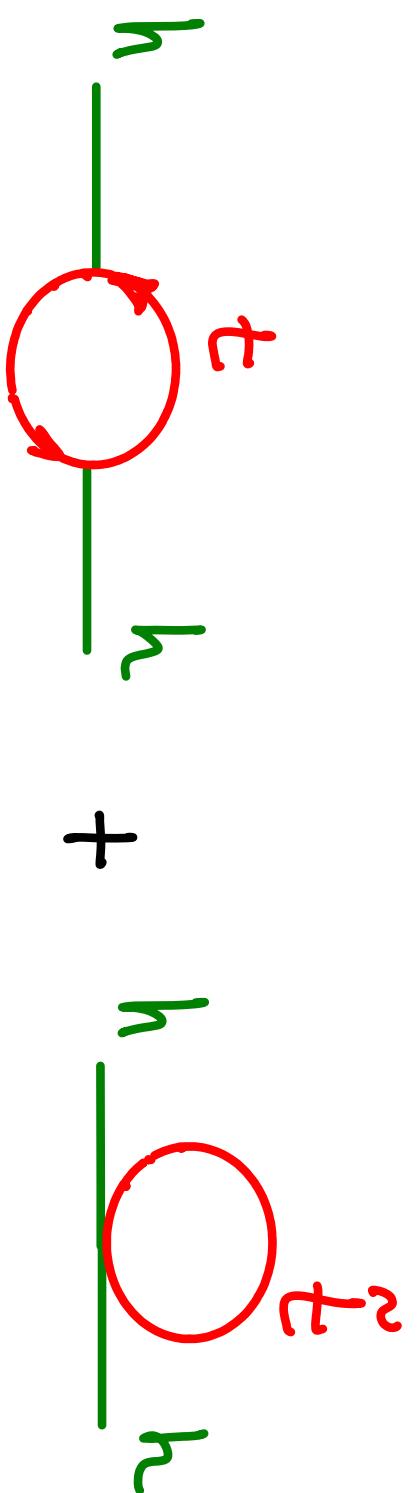
Many alternatives for such physics have been proposed

# SUPERSYMMETRY

Postulates a very powerful boson  $\leftrightarrow$  fermion symmetry with very profound implications.

Predicts a doubling of the particle spectrum.

New particles cancel out the  $\delta m^2 \sim \Lambda^2$  behaviour.  
But requires "soft" breaking of susy symmetry with mass splittings  $m_{\text{soft}} \lesssim \text{TeV}$  to maintain naturalness.



$$\delta m^2 \sim m_{\text{soft}}^2 \log \frac{\Lambda^2}{m_{\text{soft}}^2}$$

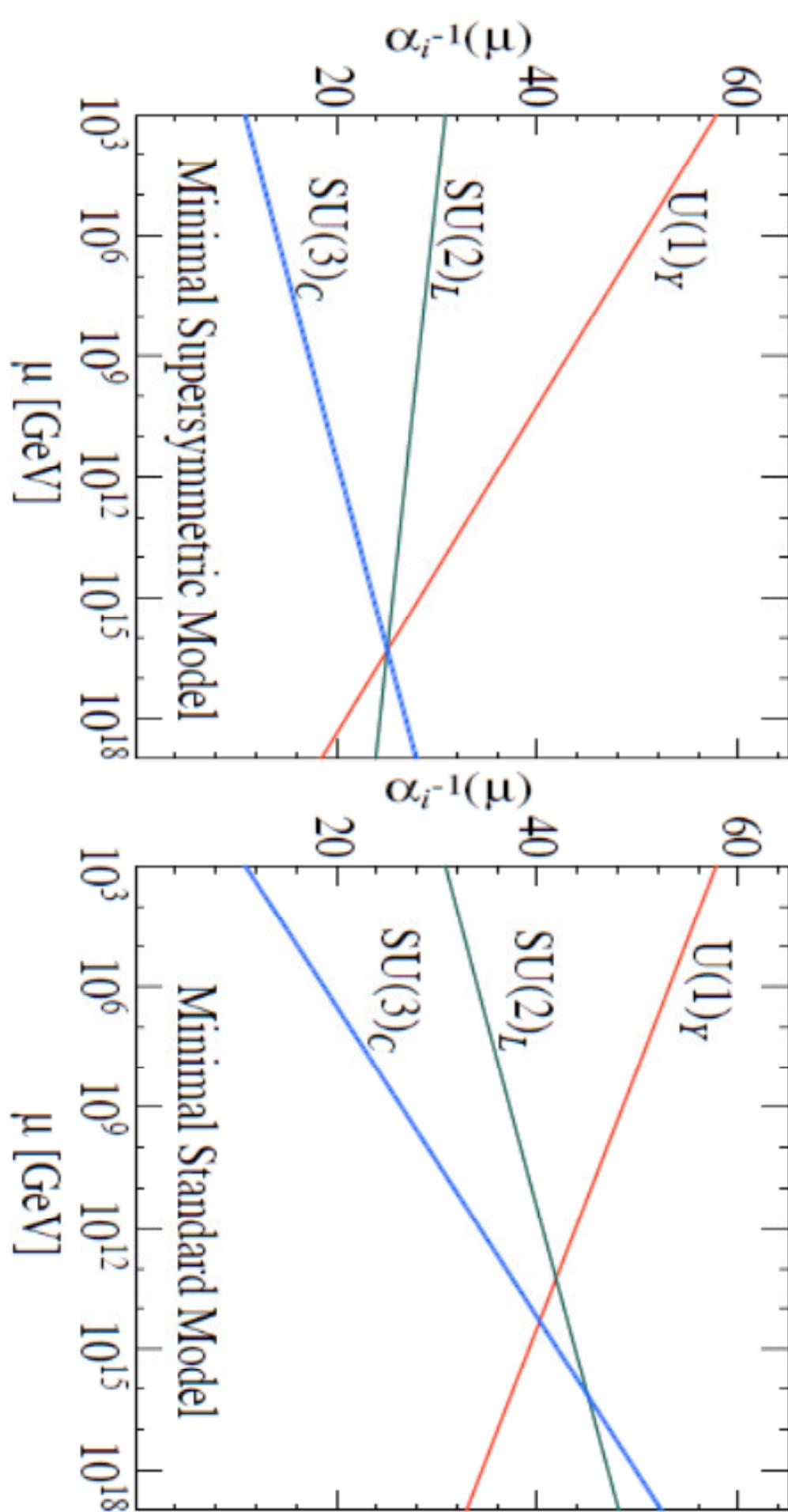
# VIRTUES OF SUPERSYMMETRY

- Based on a beautiful and serious symmetry principle  
(most general sym. of a local relativistic QFT)

$$P_\mu = \frac{1}{4i} \bar{Q} \gamma_\mu Q$$

- local version fits nicely in string theory
- extends natural range of the theory up to  $M_{Pl}$
- Perturbative  $\rightarrow$  calculable
- Simple to get fermion masses
- With R-parity implemented passes easily EW precision tests.
- BONUS: DM candidate
- BONUS: Gauge coupling unification

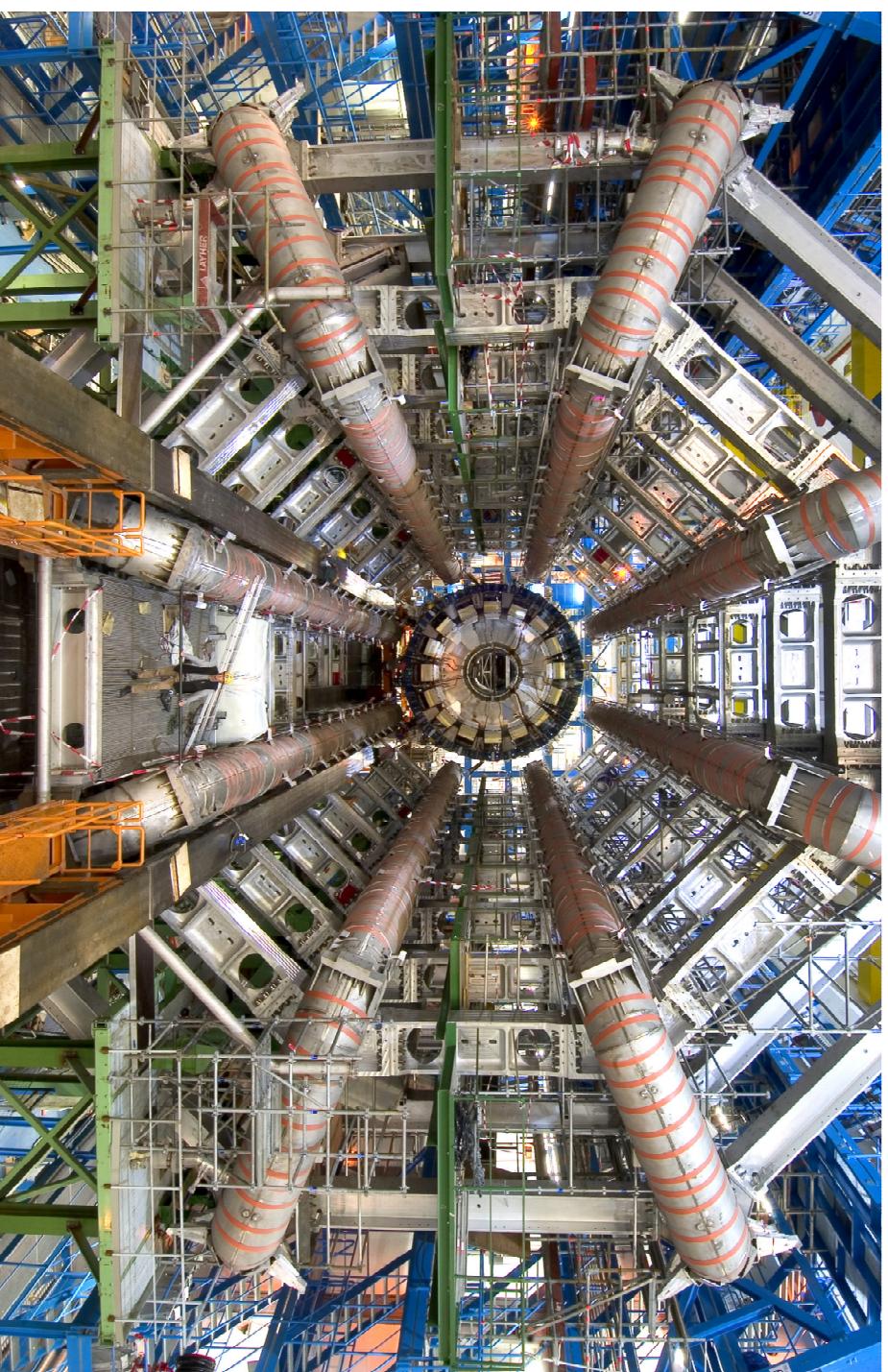
# GAUGE COUPLING UNIFICATION



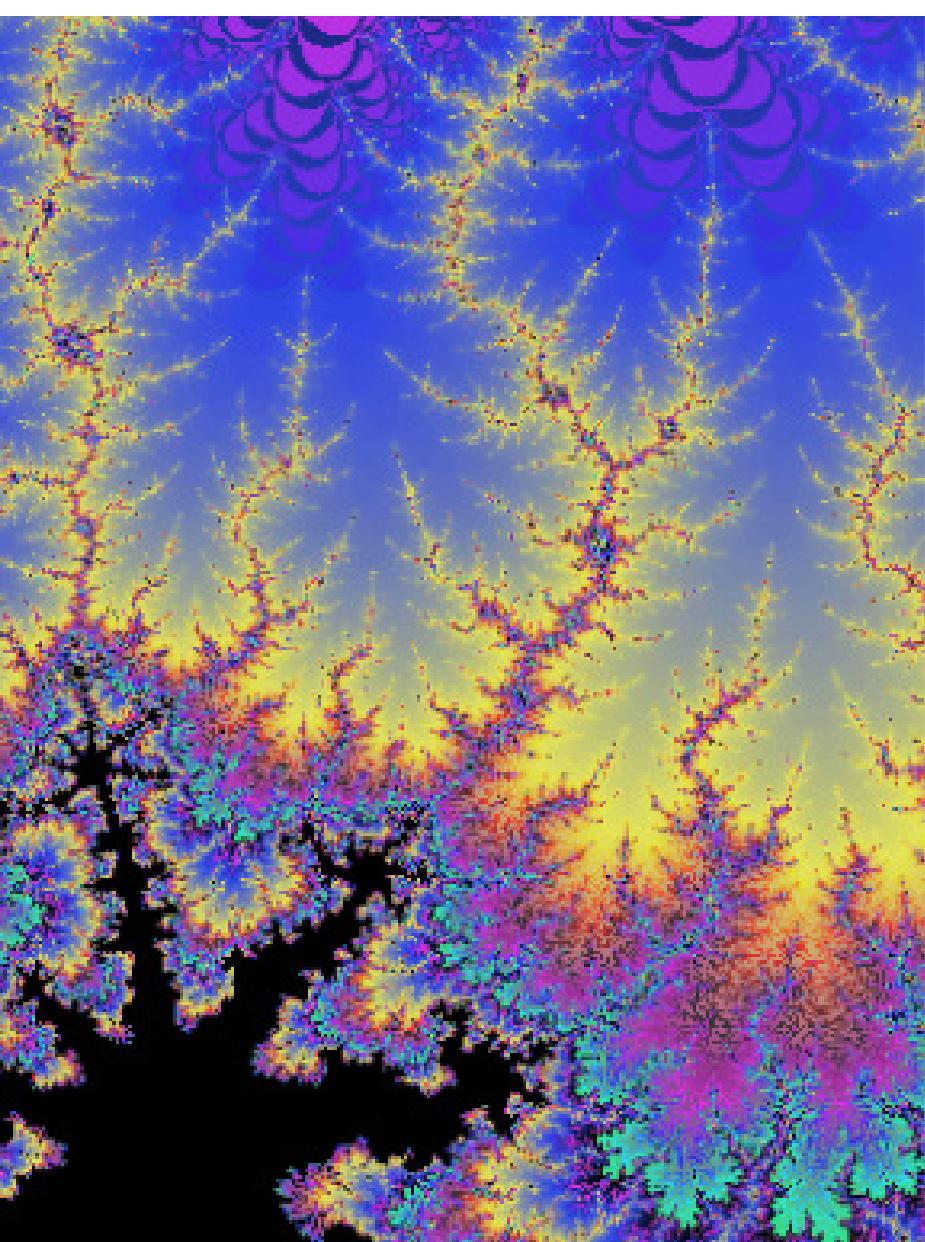
# UNPARTICLES

George '07

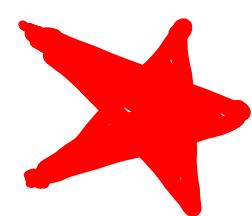
Assume a scale-invariant sector of the theory  
→ Very different from our observable sector



Our world



Scale-invariant world

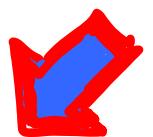


SM contains an explicit mass term

$$V(H) = -m^2|H|^2 + \lambda|H|^4$$

sets the  $\tau_W$  scale

Scale invariant sector



Massless fields \*

★ Quantum corrections can generate a scale  
(like in QCD)

Scale invariant sector



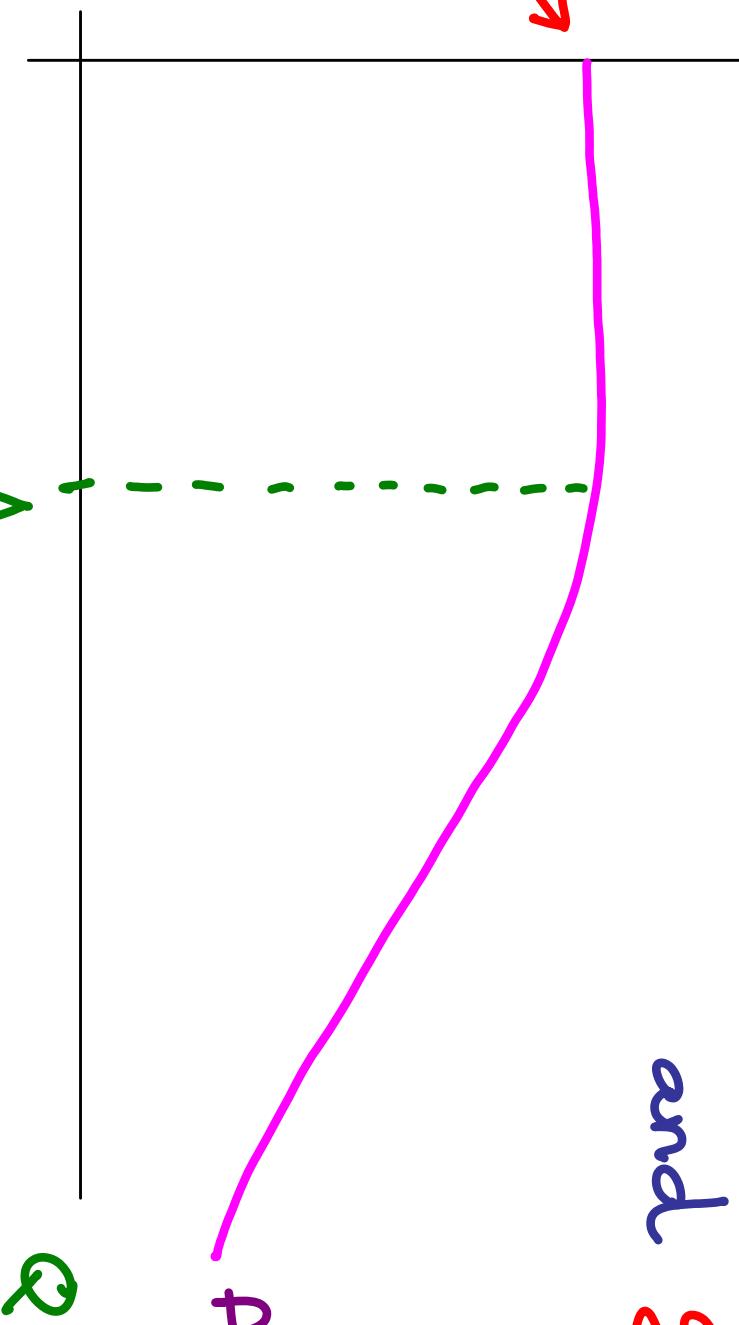
$$\beta(g) = 0$$

required

and  $g \neq 0$



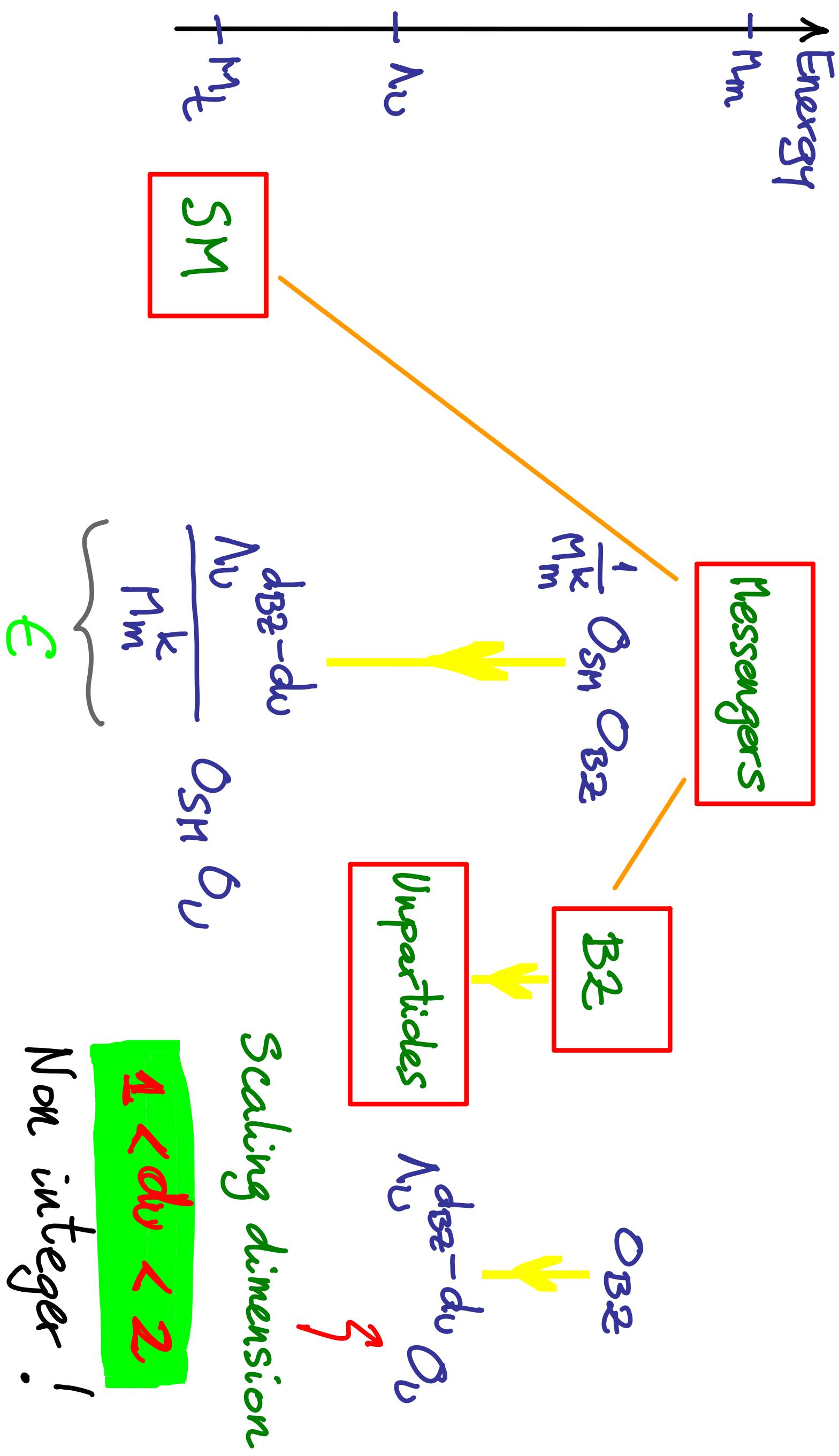
Nontrivial  
IR fixed  
point



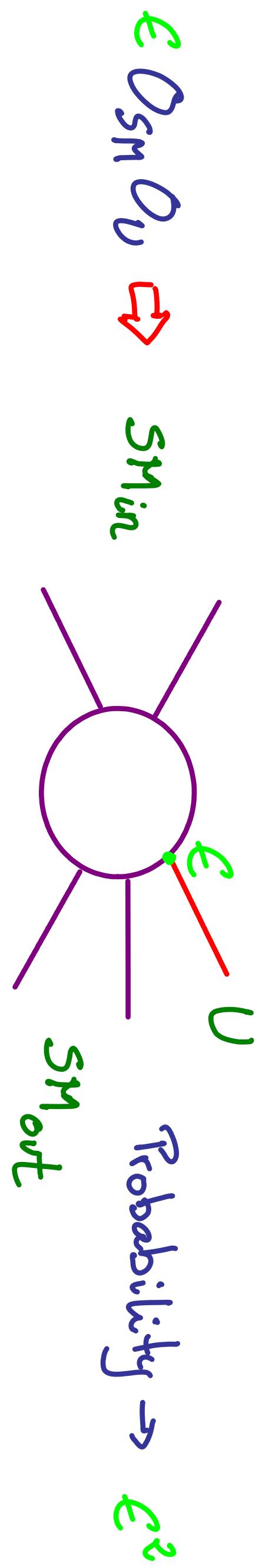
DIMENSIONAL  
ANTI-TRANSFORMATION

e.g. Banks-Zaks  $SU(N_c) N_f$

# COUPLING SM TO SCALE INV. SECTOR



# UNPARTICLE PRODUCTION



$\epsilon \propto s_{\text{min}} \rightarrow s^2$   
Probability  $\rightarrow e^{-s^2}$

Phase space for unparticles  $\rightarrow$  determined by scale invariance

$$\langle 0 | \phi_v(x) \phi_v^\dagger(0) | 0 \rangle \sim x^{-2d_v}$$

$$P_0(p^2) \sim A_d v^{(p^2)} d_v^{-2}$$

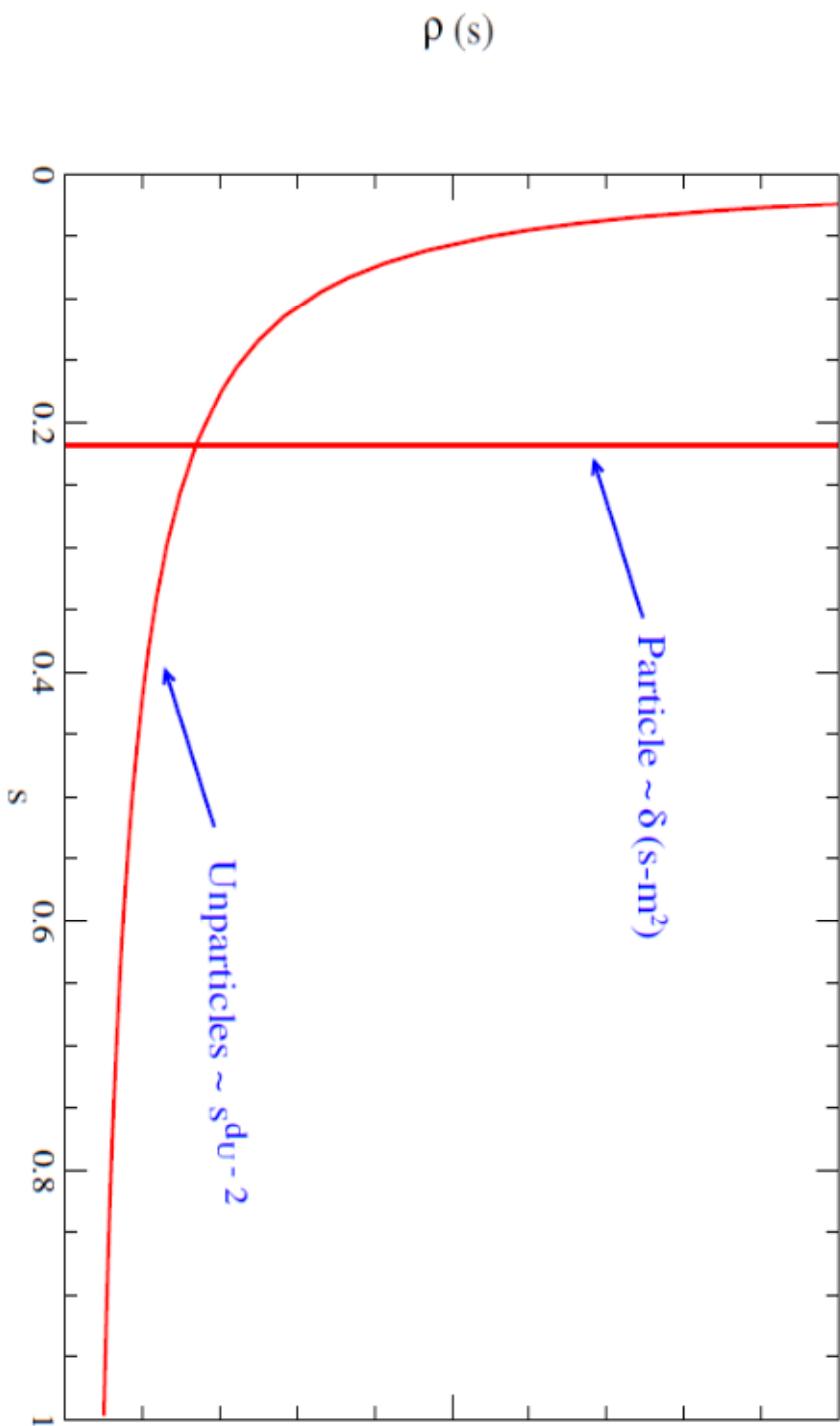
Phase space for  $n$  massless particles  $d\Omega P_S n = A_n S^{n-2}$

Unparticles  $\sim$  non-integer number of massless particles ( $!?$ )

# UNPARTICLE SPECTRAL FUNCTION

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From  $P(p^2) \sim (p^2)^{d_U - 2} \rightarrow \rho(s) \sim s^{d_U - 2}$



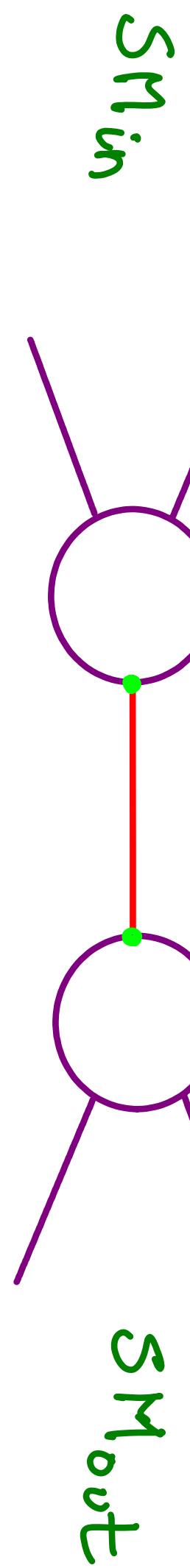
O<sub>V</sub> does not create particles out of the vacuum but  
a non-localized wave over the full range of  $p^2$

↳ Unparticles



lots of phenomenological implications  
have been explored

Example :



$$\text{Interference} \sim e^2 P_u (P^2)$$

Strange phase inside

# DECONSTRUCTING UNPARTICLES

Stephanov '07

Unparticles as infinite tower of massive scalars  $\varphi_n$  ( $n=1, \dots, \infty$ )

$$m_n^2 = \Delta^2 n$$

In the limit of zero mass splitting  $m_{n+1}^2 - m_n^2 = \Delta^2 \rightarrow 0$

→ Scale invariant continuous mass spectrum!

$$S = \int d^4x \sum_{n=1}^{\infty} \left[ \frac{1}{2} (\partial^\mu \varphi_n)^2 + \frac{1}{2} m_n^2 \varphi_n^2 \right]$$

$$\varphi_n(x) \rightarrow \lambda \varphi_n(x\lambda)$$

$$\delta^2 \rightarrow d\lambda^2$$

$$\varphi_n \rightarrow \Delta \cdot u(\lambda^2)$$



$$S = \int d^4x \int_0^\infty dm^2 \left[ \frac{1}{2} (\partial^\mu u)^2 + \frac{1}{2} m^2 u^2 \right] u(m^2, x) \rightarrow u(m^2/\lambda, \lambda x)$$

Deconstructed  $\phi_U$ :

$$\phi = \sum_{n=1}^{\infty} f_n \varphi_n \rightarrow \phi_U$$

$$f_n^2 = \frac{Ad\omega}{2\pi} \Delta^2 (M_n^2)^{dw-2}$$

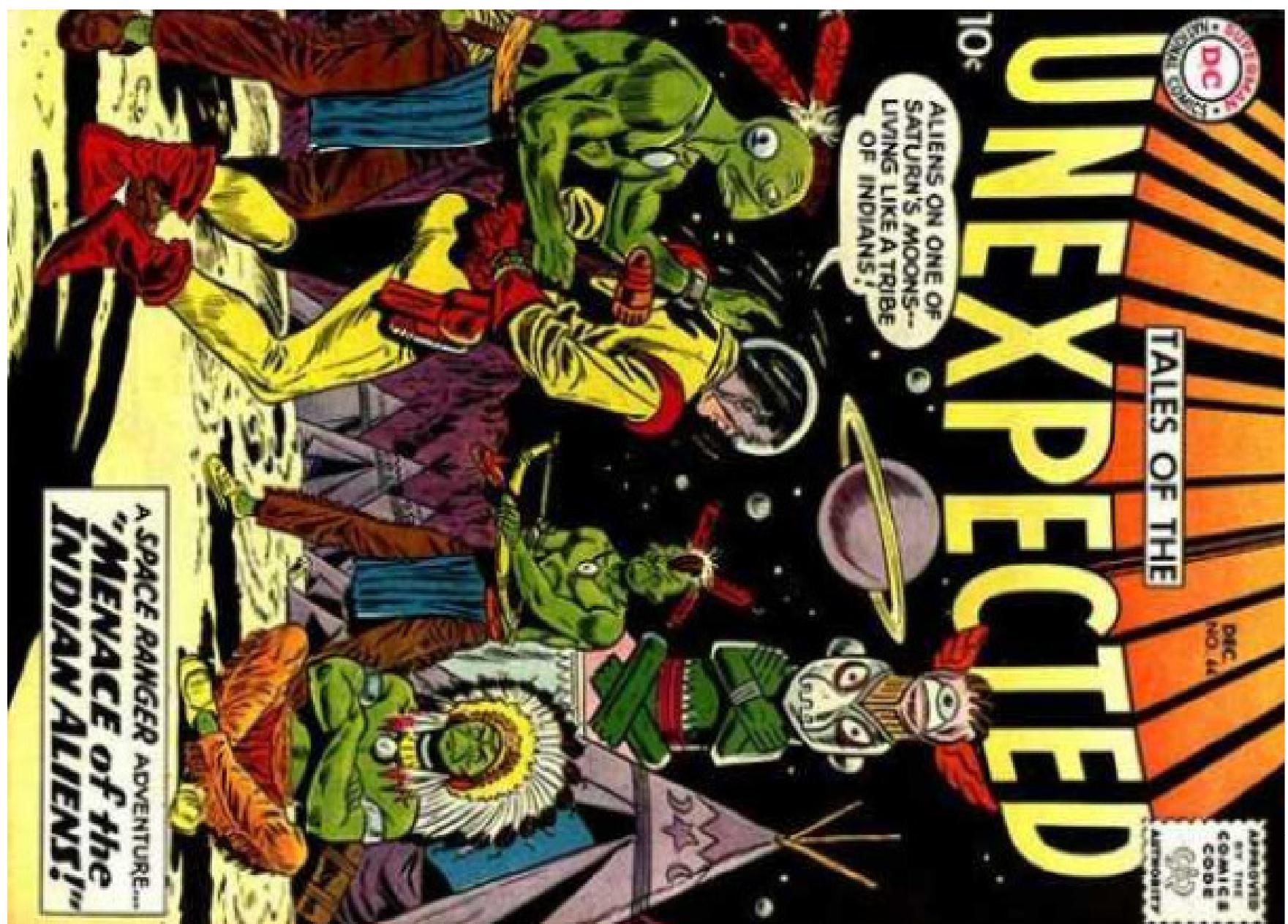
→  $\langle \phi \phi^\dagger \rangle \rightarrow \langle \phi_U \phi_U^\dagger \rangle$



Useful calculational tool. Related to extra-d

scenarios: RS 2, soft-wall...

It's difficult to have truly original ideas!



# HIGGS AS A PORTAL TO UNPARTICLE WORLD

Fox, Rajeraman, Shirman '07

Delgado, E, Quiros '07

There is room for a relevant operator to  $H$ :

$$K \nu O_{SU} \rightarrow K \nu |H|^2 O_U$$

$\hookrightarrow$  dimension  $2-d_U > 0$

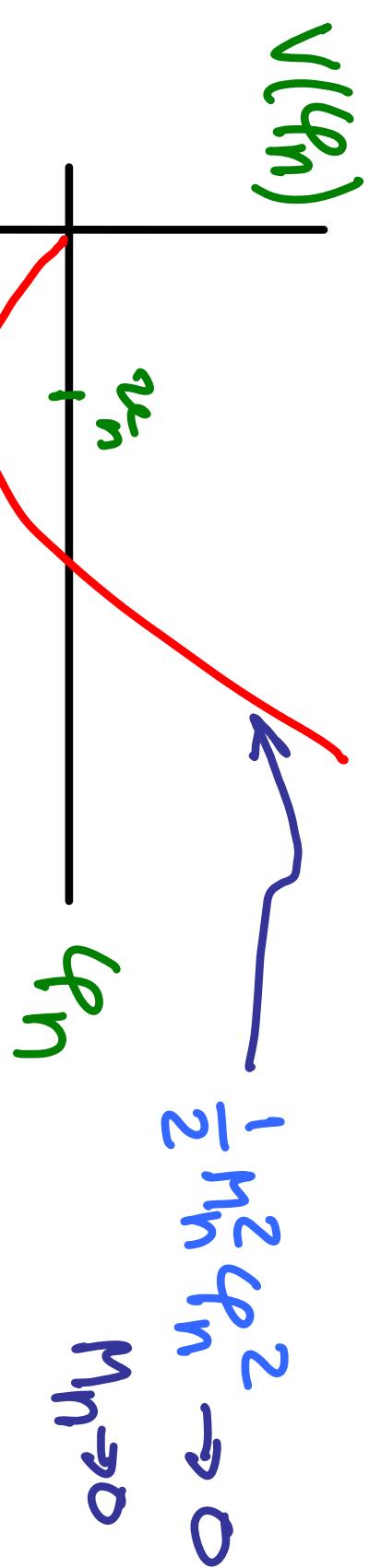
Main source of conformal breaking after EWSB

Diffractivity :  $\langle H \rangle = \frac{q^2}{2}$  induces a tadpole for  $O_U = \sum_n f_n \phi_n$

and  $\langle O_U \rangle$  has an IR divergence

# ORIGIN OF DIVERGENCE

Calculate  $\langle \partial_U \rangle$  through  $\langle \varphi_n \rangle = v_n$



$$\frac{1}{2} m_n^2 \varphi_n^2 \rightarrow 0$$

$$m_n \rightarrow 0$$

$$\left( \frac{1}{2} k_B T_n \right) \varphi_n \rightarrow \infty$$

$$m_n \rightarrow 0$$

$$\langle \partial_U \rangle \propto v^2 \sum_n v_n \sim$$

$$\int_0^\infty (m^2) dv - 3 dM^2$$

**IR div for**  
 $d_U < 2$

# SIMPLE CURES FOR IR PROBLEM

I. Add coupling

$$\underbrace{\frac{1}{4} H^2 \sum_n \phi_n^2}$$

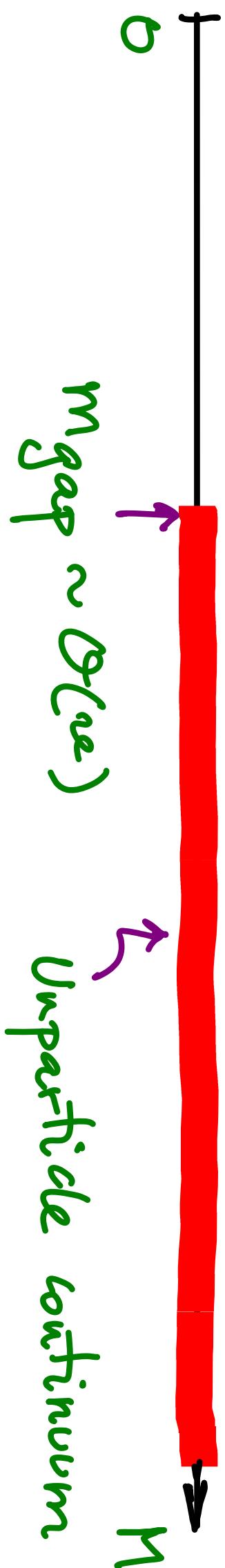
acts as IR cutoff mass

II. Add coupling

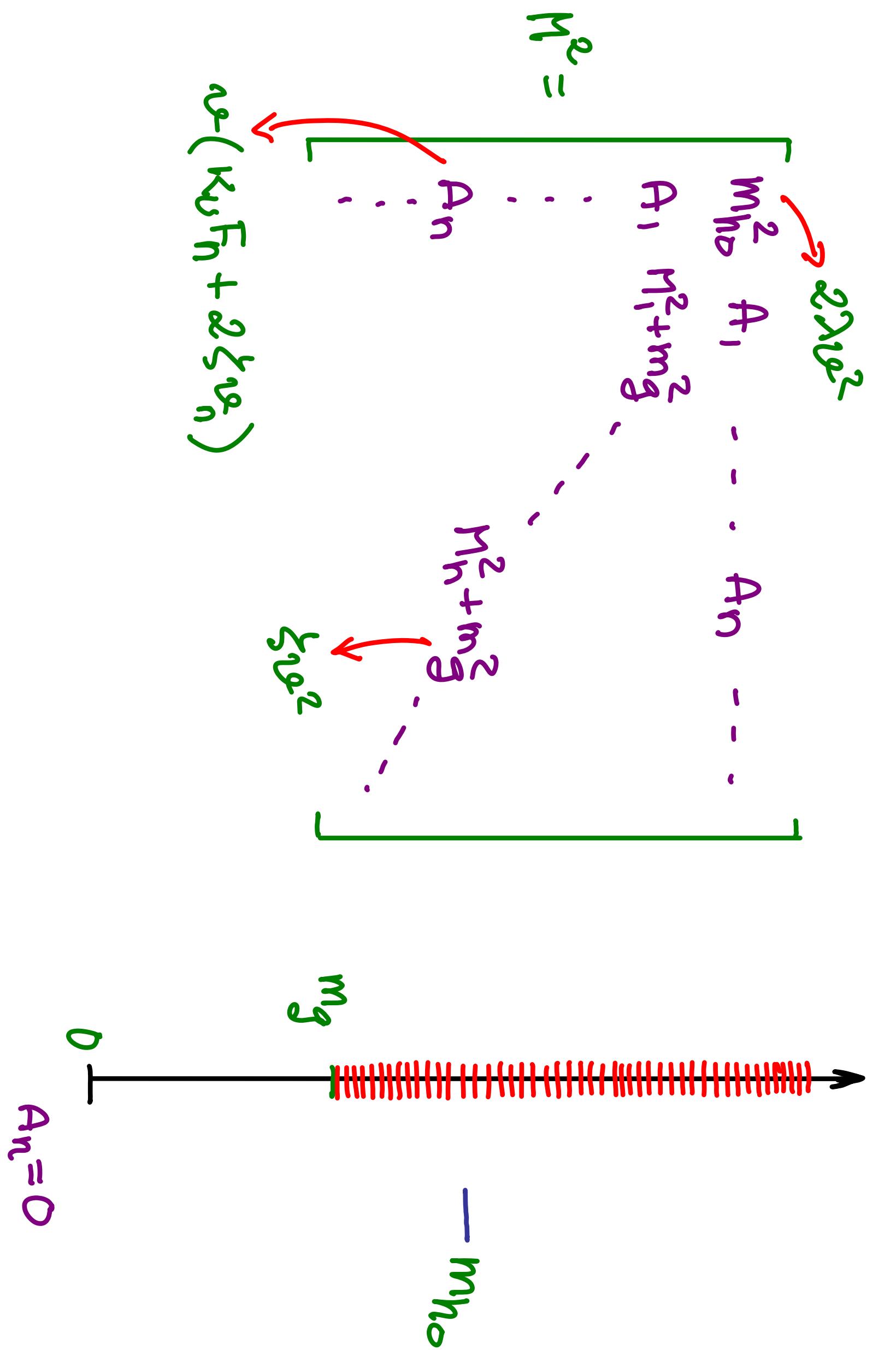
$$\frac{1}{4} \lambda_0 \left( \sum_n \phi_n \right)^2$$

generates a mass  $\sim \lambda_0 \left( \sum_n \phi_n \right)^2$

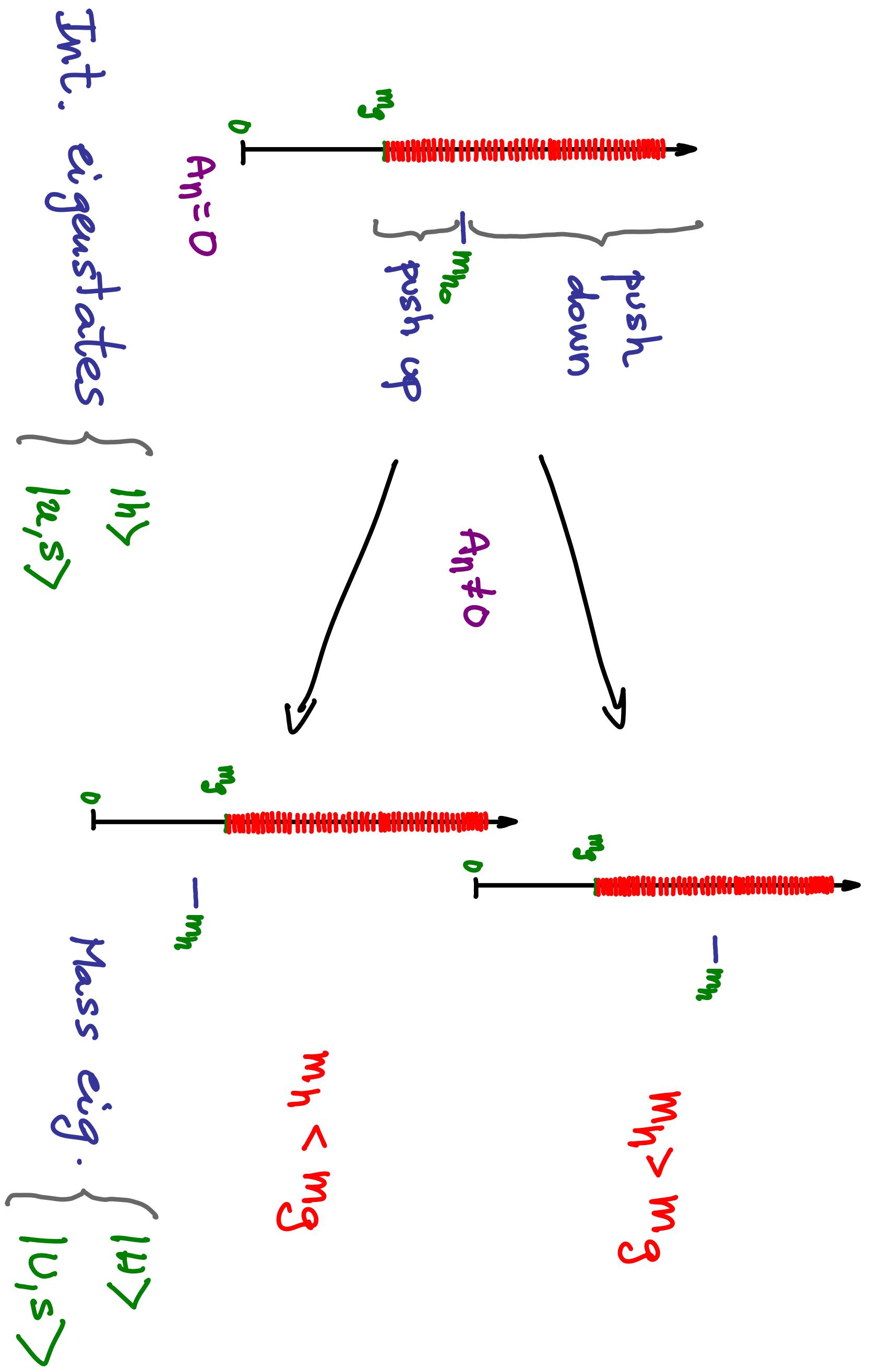
→ A mass gap is generated



# Higgs-unparticle Interplay



# Two Possible Outcomes:



# MIXED PROPAGATOR

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or

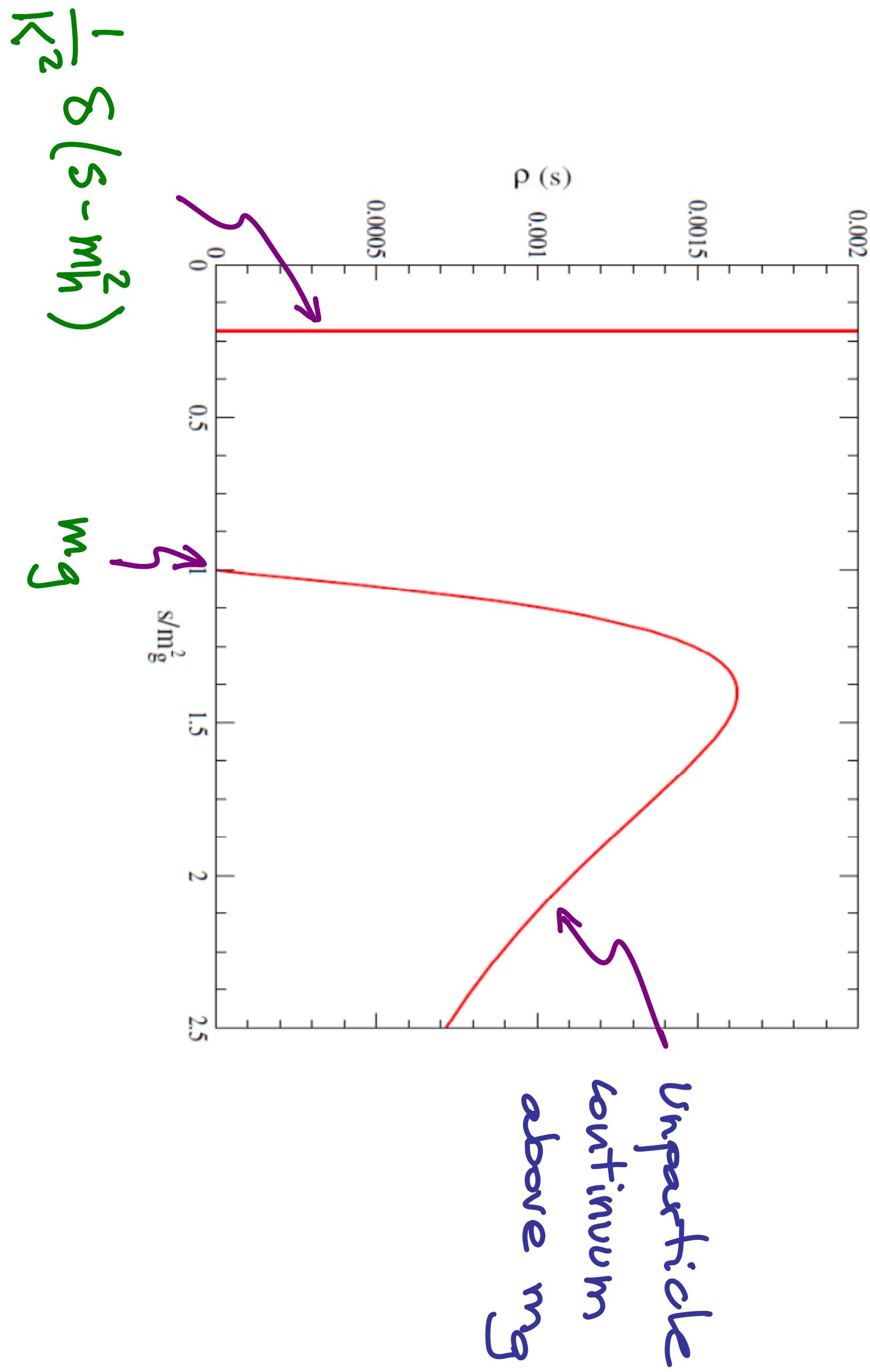
$$h = \text{---} + \text{---} \times \frac{U}{X} + \dots$$

$$U - \text{---} = \text{---} + \text{---} \times h + \dots$$

$$\Rightarrow i\Gamma_{hh}(p^2) = p^2 - m_h^2 + \omega^2(\mu_0^2)^{2-d} \int_0^\infty \frac{(M^2)dw^{-2}}{(M^2 + m_g^2 - p^2)} \left[ \frac{M^2}{M^2 + m_g^2} \right]^2 dm^2$$

We can study its poles or better still obtain the spectral function  $\rho_{hh}(s)$

# SPECTRAL FUNCTION, $m_h < m_g$



# INTERPRETATION OF $\rho(s)$

$$\rho_{hh}(s) \equiv \langle h|s\rangle\langle s|h\rangle = \underbrace{|\langle h|H\rangle|^2}_{R_h} \delta(s-m_h^2) + \underbrace{\Theta(s-m_g^2)|\langle h|U_s s\rangle|^2}_{R_U(s)}$$

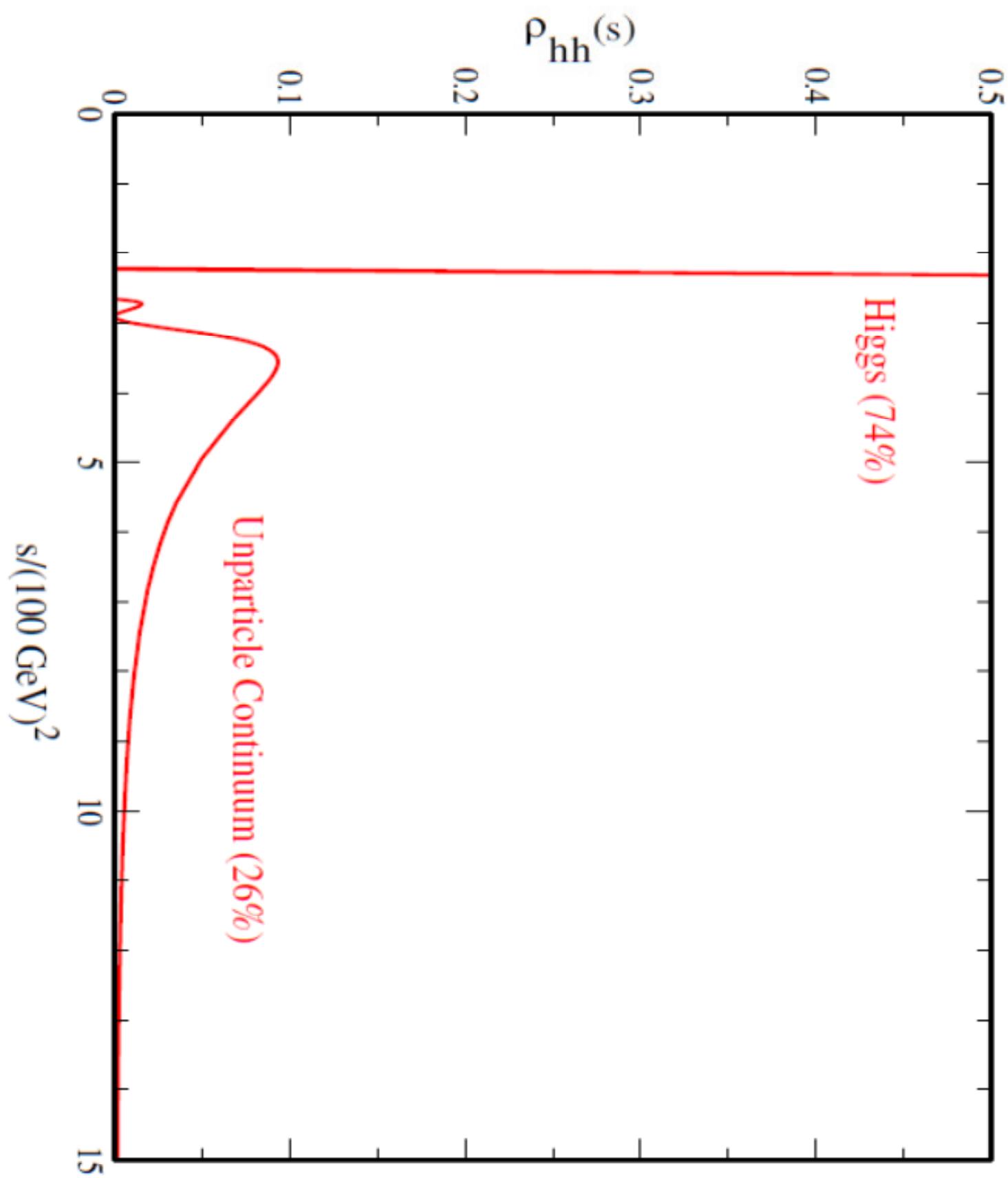
★  $H^2\bar{2}$  coupling diluted by unparticle contamination of isolated pole. Pure Higgs composition  $R_h \equiv \langle h|H\rangle = \frac{1}{K}$

★  $U^2\bar{2}$  coupling induced by Higgs contamination ~ 'Higgsness' diffuses in unparticle continuum

Sum Rule

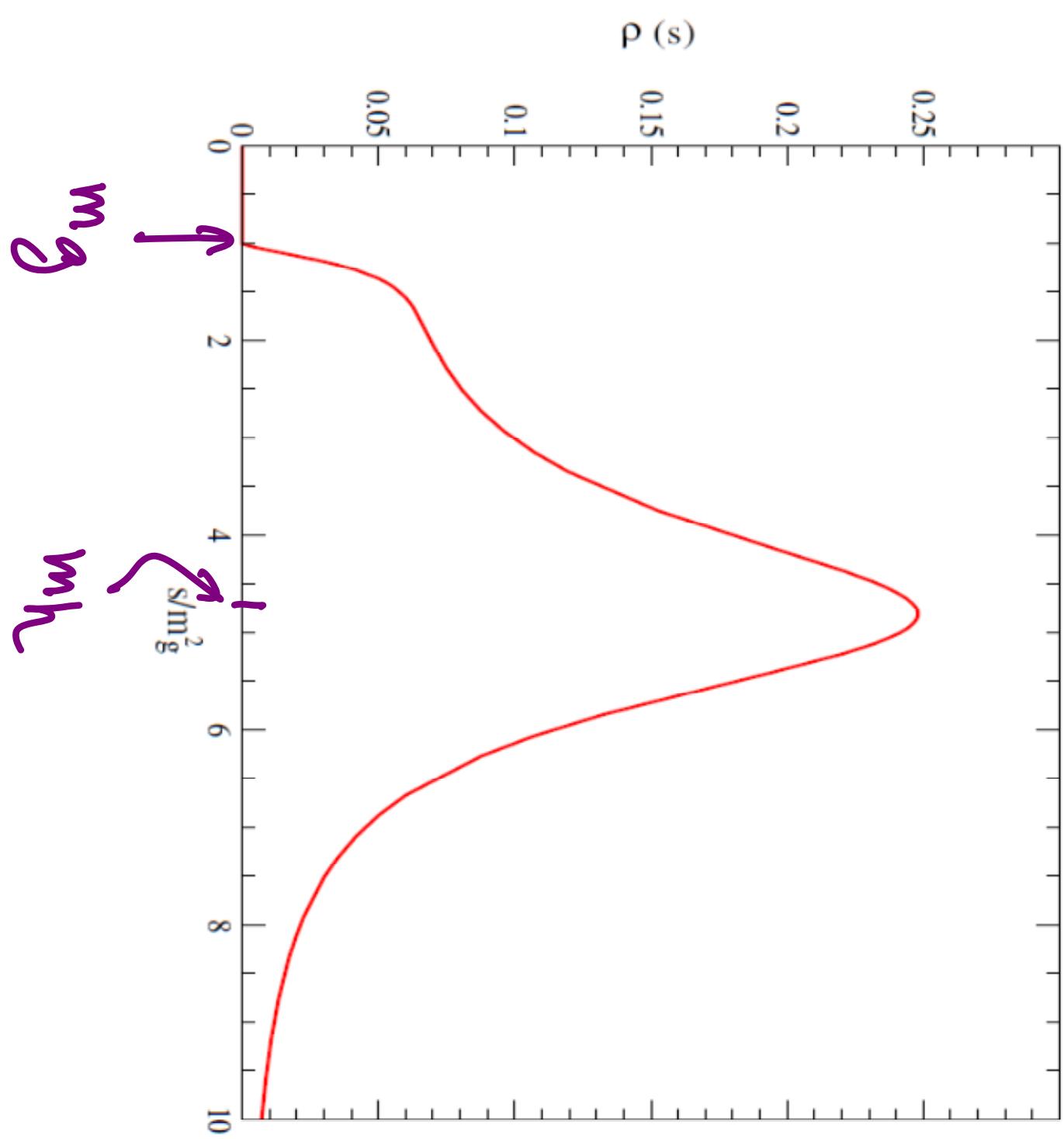
$$R_h^2 + \int_{m_g^2}^{\infty} R_U^2(M^2) dM^2 = 1 = \langle h|h\rangle$$

# EXAMPLE



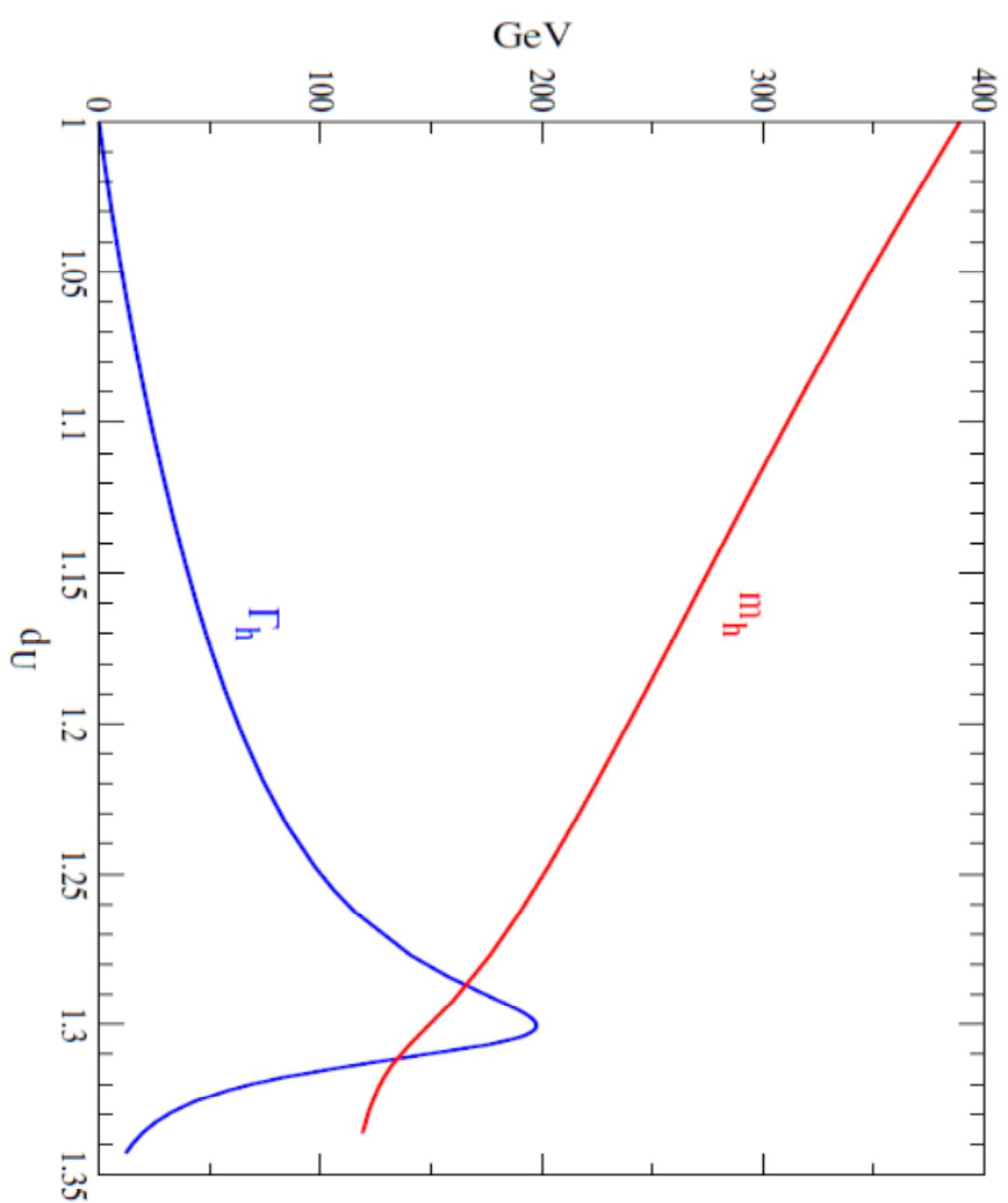
# SPECTRAL FUNCTION, $m_h > m_g$

Higgs pole subsumed in unparticle continuum



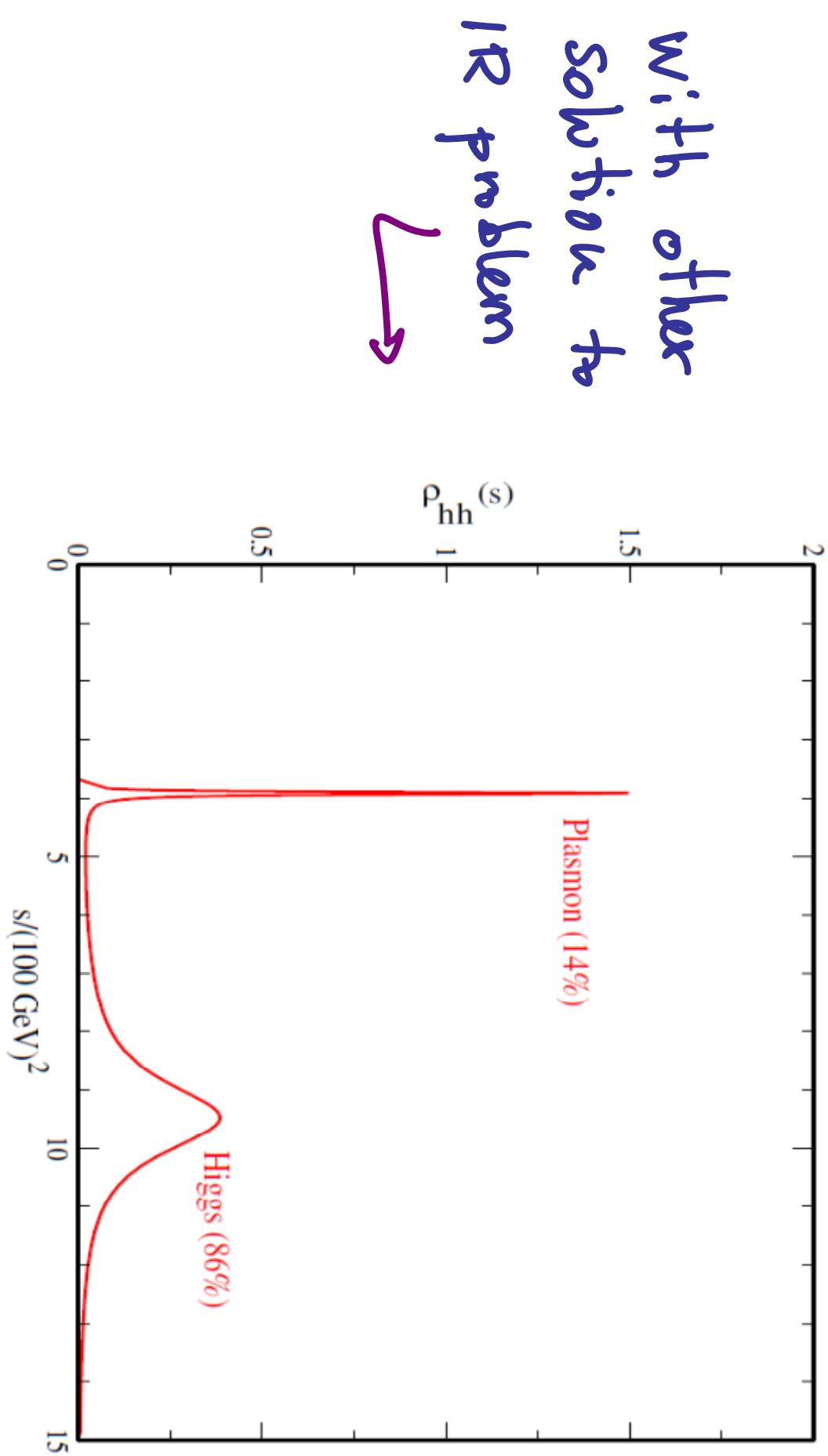
large  
mixing  
width

Higgs width,  $m_h > m_g$



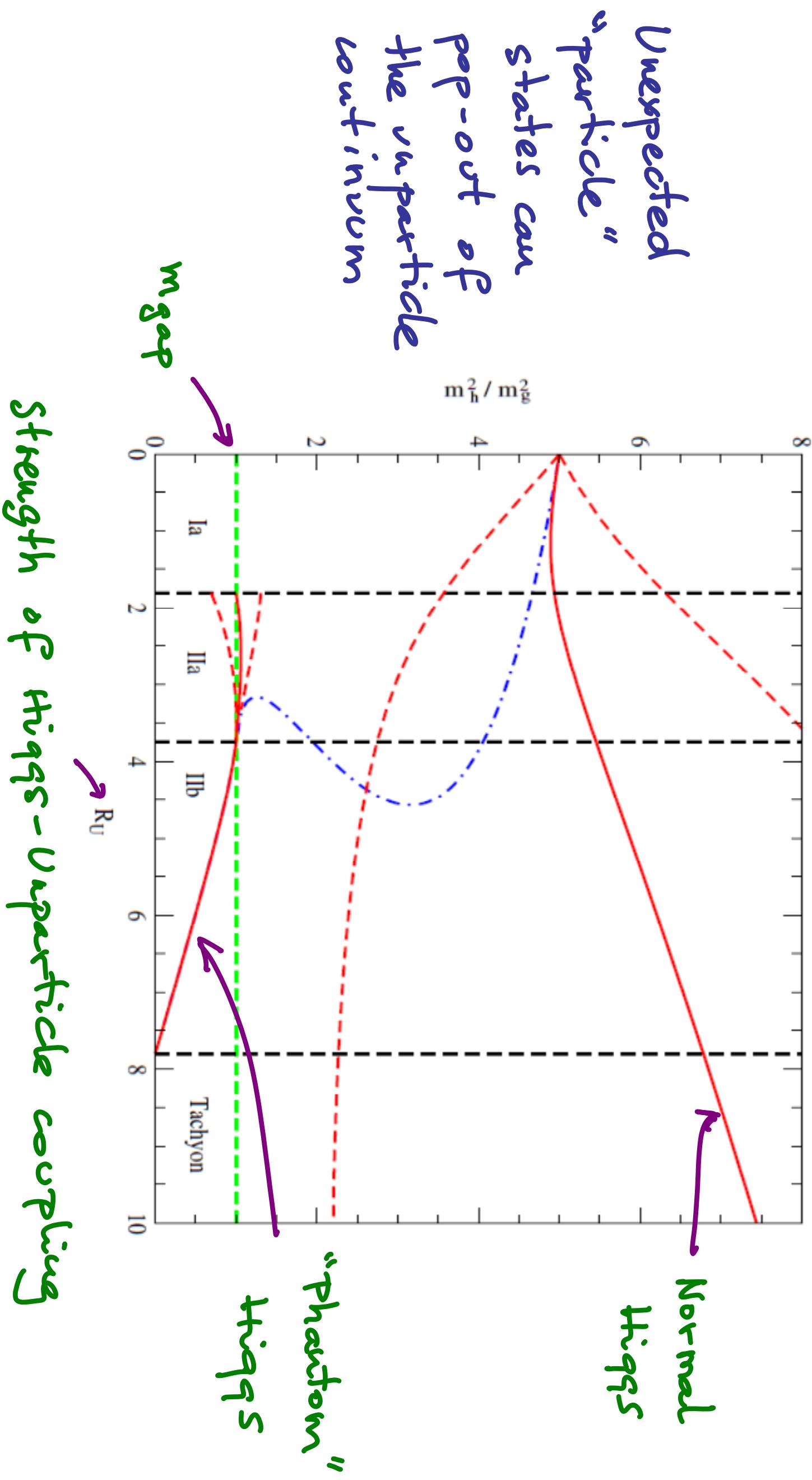
# OTHER EFFECTS

Many similarities to effects studied in condensed matter  
(e.g. Anderson-Fano)



Higgs can excite "plasmonic" resonances in the unperturbed continuum

# OTHER EFFECTS, CONT'D



# Conclusions

- ★ We don't really know what LHC will find  
*that's why it's so exciting !*
- ★ We do have good reasons to expect physics beyond the SM at the reach of LHC
- ★ To find **SUSY** would be a triumph of theoretical imagination and of experimental performance
- ★ To get the most of LHC we have to be prepared for all reasonable alternatives we can imagine
- ★ In any case, we have **exciting times** ahead !

# IMPLICATIONS: EFFECTS ON ENSE

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Deconstructed potential:

$$V = m^2 |H|^2 + \lambda |H|^4 + \frac{1}{2} \sum_n m_n^2 \varphi_n^2 + \kappa |H|^2 \sum_n f_n \varphi_n + \zeta |H|^2 \sum_n \varphi_n^2$$

Minimization

$$m^2 + \lambda v^2 + \kappa \sum_n f_n v_n + \zeta \sum_n v_n^2 = 0$$

$$\hookrightarrow v_n = -\frac{\kappa v^2 f_n}{2(m_n^2 + \zeta v^2)}$$

$$\text{Continuum} \rightarrow m^2 + \lambda v^2 - \lambda \nu (\mu_\nu^2)^2 - \nu^2 2(\nu^{-1}) = 0$$

like coming from  $\delta V \sim h^2 d\nu$ !

+ It's possible to have  $v \neq 0$  even for  $m^2 > 0$