

How many bodies are required before we have a problem? G.E. Brown points out that this case can be answered by a look at history. In XVIII century Newtonian mechanics, the three-body problem was insoluble. With the birth of relativity around 1910 and quantum electrodynamics in 1930, the two- and one-body problems become insoluble. And within modern quantum field theory, the problem of zero bodies (vacuum) is insoluble. So, if we are out after exact solutions, no bodies at all is already too many.

R.D. Mattuck, A guide to Feynmann diagrams
McGraw Hill , NY 1976

Impresje niskoenergetycznego obserwatora

- narzędzia: SPS \rightarrow RHIC \rightarrow LHC
- elementarz: y , p_t , detektory
- motywacje
- wyniki „globalne”: krotności, temperatura, potencjał bariochemiczny
- „jet quenching”
- mezony w materii
- jak mierzyć fotony w polu magnetycznym?

Some Pre-Conference Advice

Advice to Experimentalists:

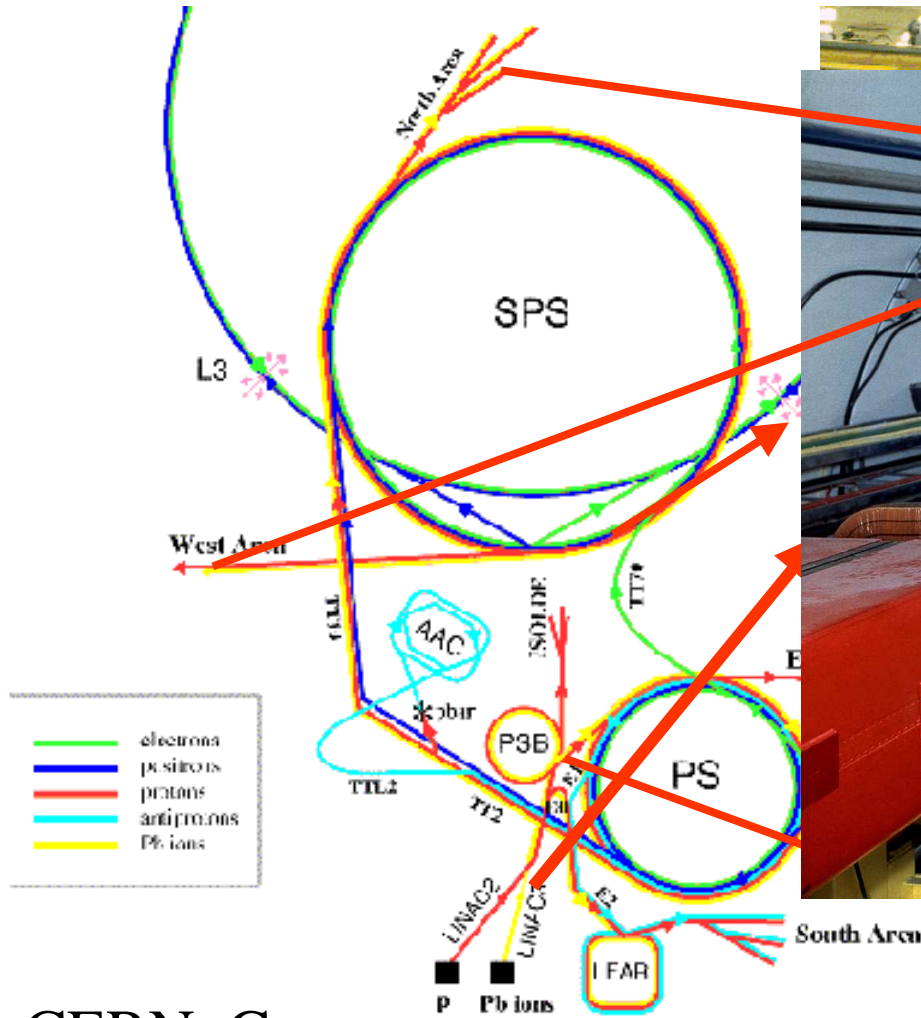
Never let a theorist tell you something is too complicated to explain.

Advice to Theorists:

Never let an experimentalist tell you something is too complicated to explain.

Most all things are really simple once we understand.

Super Proton Synchrotron SPS



CERN, Geneva



160A GeV Pb Beam
for physics
 $\sqrt{s} \sim 17A$ GeV

- West Area
WAXX experiment (Switzerland)
- North Area
NAXX experiment (France)

SPS Tunnel

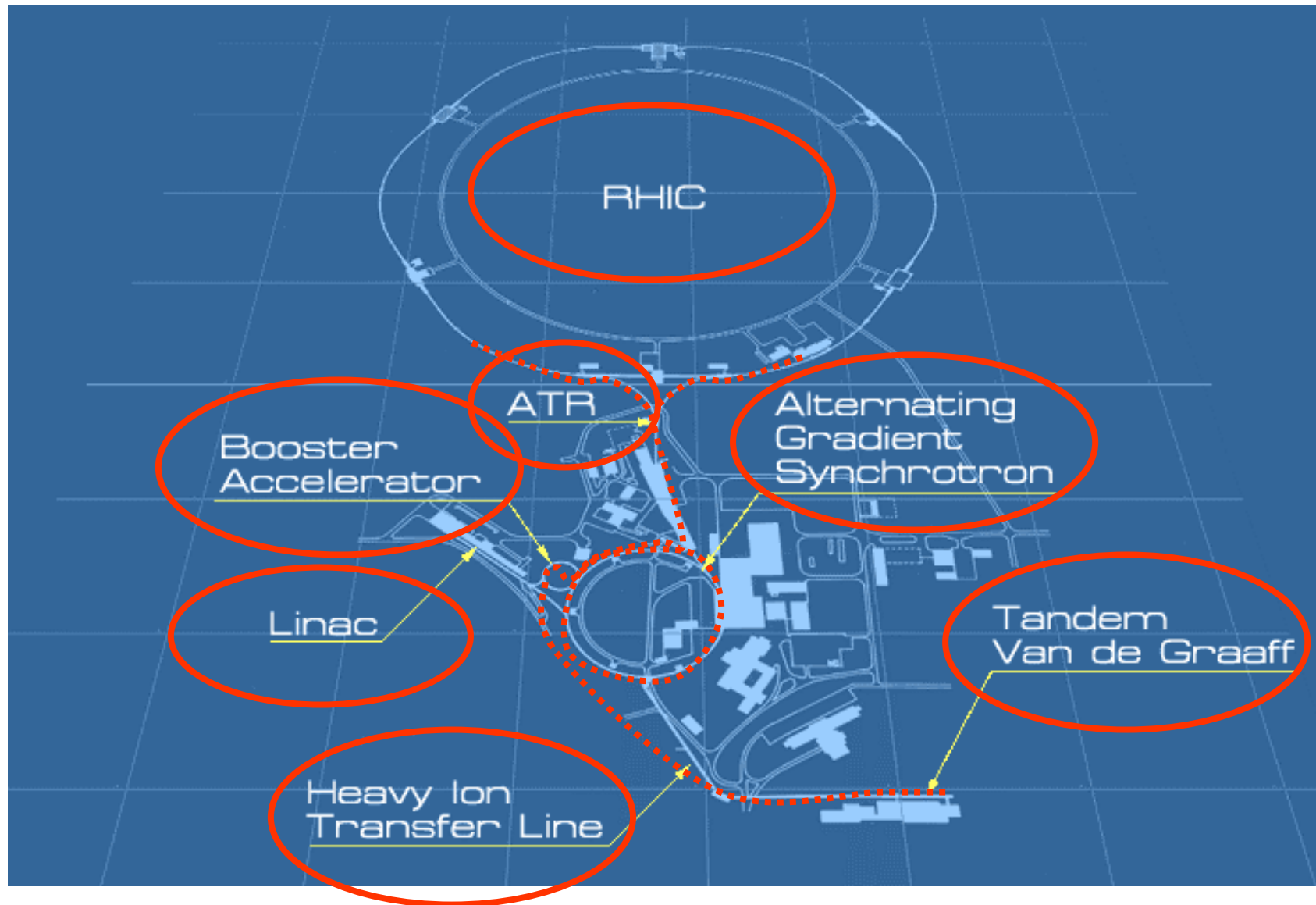
Relativistic Heavy Ion Collider, RHIC

- 3.83 km circumference
- Two *separated* rings
 - 120 bunches/ring
 - 106 ns bunch crossing time
- A+A, p+A, p+p
- Maximum Beam Energy :
 - 500 GeV for p+p
 - 200A GeV for Au+Au
- Luminosity
 - Au+Au: $2 \times 10^{26} \text{ cm}^{-2} \text{ s}^{-1}$
 - $\vec{p}+\vec{p}$: $2 \times 10^{32} \text{ cm}^{-2} \text{ s}^{-1}$
- **Mid-rapidity at 90°**
- **Interaction Point**



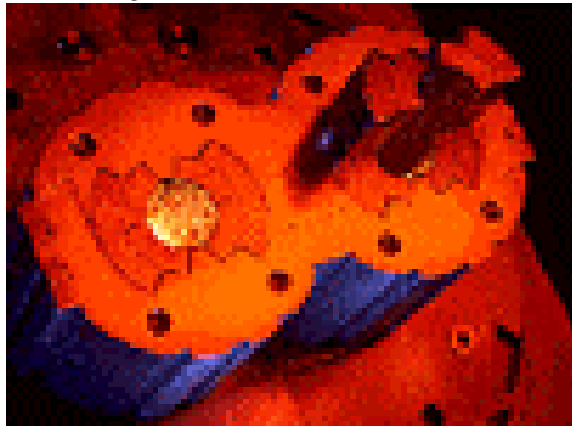
Upton, Long Island, New York

Beam @ RHIC Complex



Large Hadron Collider, LHC

- Pb^{+82} @ 2.76A TeV
- Initial Luminosity $10^{27} \text{ cm}^{-2}\text{s}^{-1}$
- Luminosity half-life 4.2 h
 - 430b of e.m. processes in Pb+Pb collisions
 - Quench of quadrupoles
- Of course
 - p @ 7 TeV
 - $L_0 = 10^{34} \text{ cm}^{-2}\text{s}^{-1}$



p+p @ 14A TeV in ~2007
Pb+Pb @ 5.5A TeV in ~2008

Why Rapidity?

Kinematical reason:

- The shape of the rapidity distribution, dn/dy , is invariant

$$y = 0.5 \times \ln \left(\frac{E + p_z}{E - p_z} \right) \longrightarrow y^* = y + y_0$$

Dynamical reason:

- The invariant cross-section can be factorized

$$\frac{d^2\sigma}{2\pi p_T dy dp_T} = \frac{d\sigma}{2\pi p_T dp_T} \times \frac{dn}{dy}$$

Pseudo rapidity

$$\eta = -\ln \tan\left(\frac{\theta}{2}\right)$$

$$y \approx \eta, \quad p \gg m, \quad \theta \gg 1/\gamma$$

$$y \approx \eta \approx \pi/2 - \theta$$

$$\gamma_{\text{SPS}} = 9, \quad \theta \gg 6^\circ \quad // \quad \gamma_{\text{RHIC}} = 100, \quad \theta \gg 1.6^\circ \quad // \quad \gamma_{\text{LHC}} = 2750, \quad \theta \gg 0.02^\circ$$

Maximum Rapidity

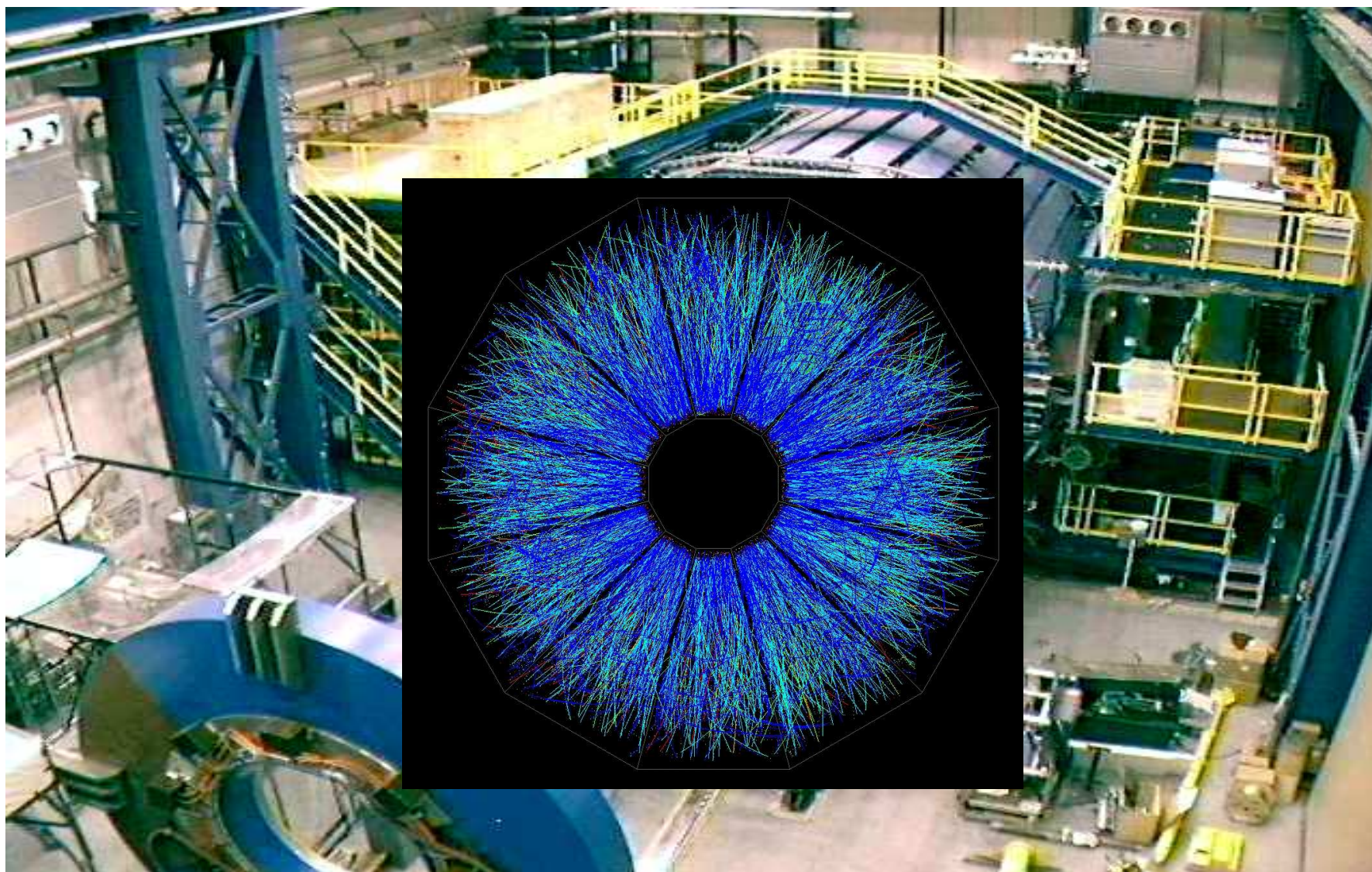
$$y_{\text{max}} = \ln\left(\frac{\sqrt{s}}{m}\right)$$

$$y_{\text{max}}^{\text{SPS}} = \pm 2.8$$

$$y_{\text{max}}^{\text{RHIC}} = \pm 5.3$$

$$y_{\text{max}}^{\text{LHC}} = \pm 8.6$$

STAR (working) TPC



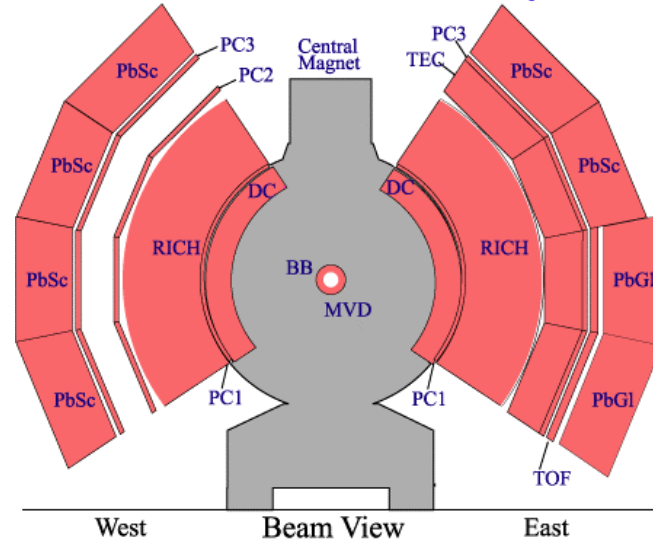
Charged Particle Identification at PHENIX

HBT analysis

Time-of-Flight
by Calorimeter

- large acceptance ($\Delta\phi = \pi$)

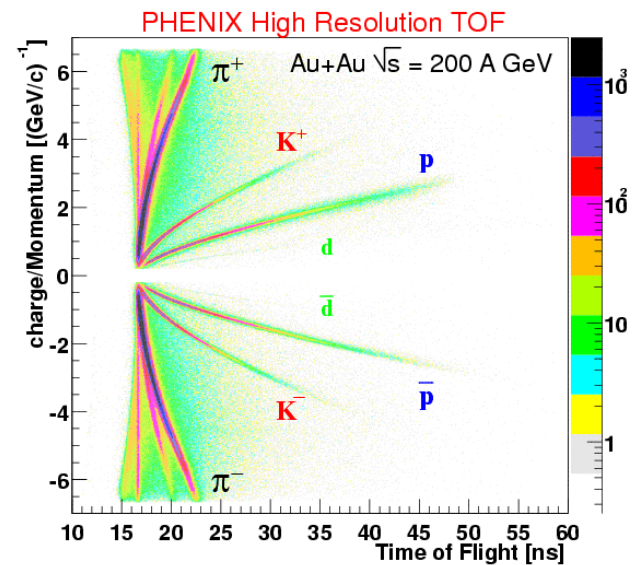
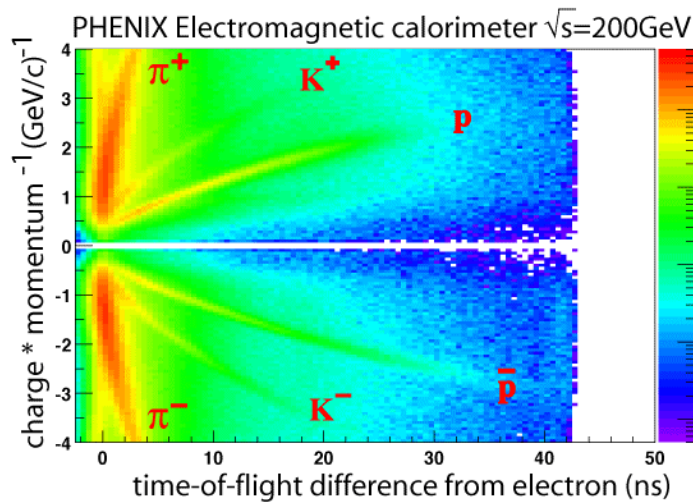
PHENIX Detector - Second Year Physics Run



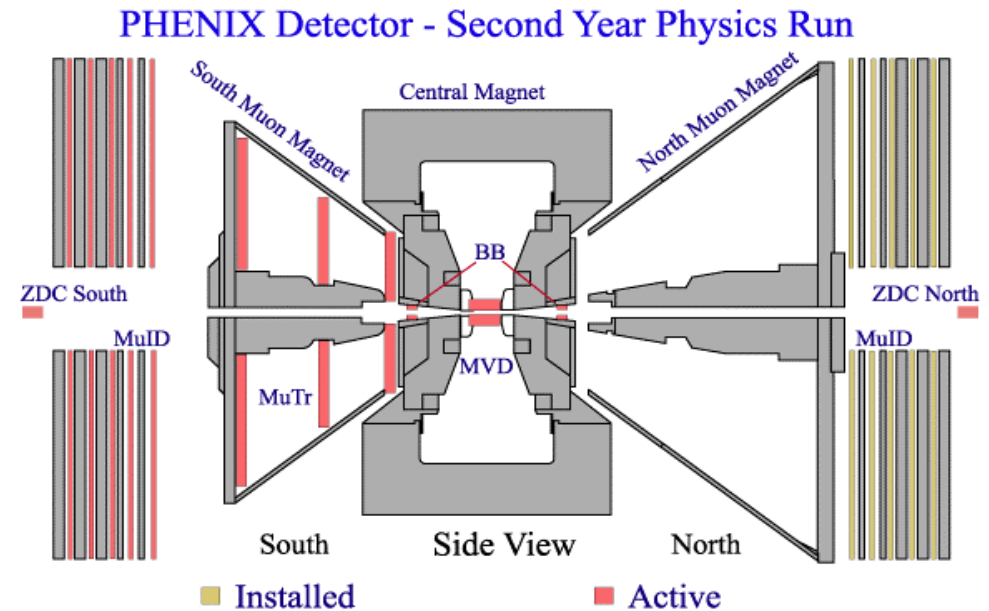
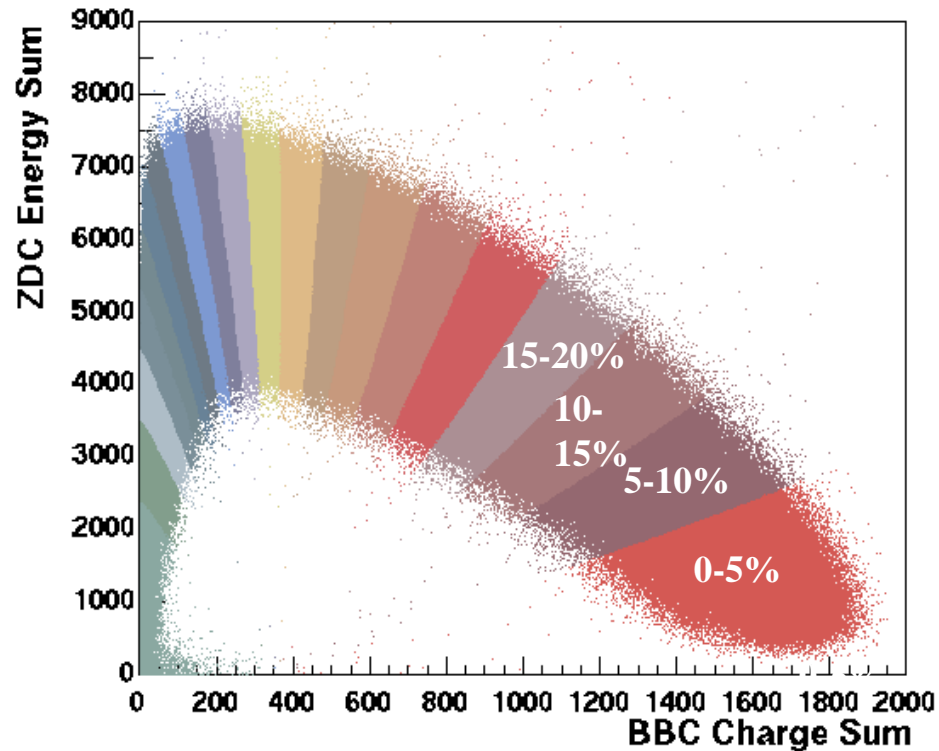
Single particle spectra and elliptic flow w.r.t reaction plane analysis

PID by high resolution TOF

- broad p_T range
 - $\pi, K < 2$ GeV
 - proton, anti-proton < 4 GeV
- $\Delta\phi = \pi/4$



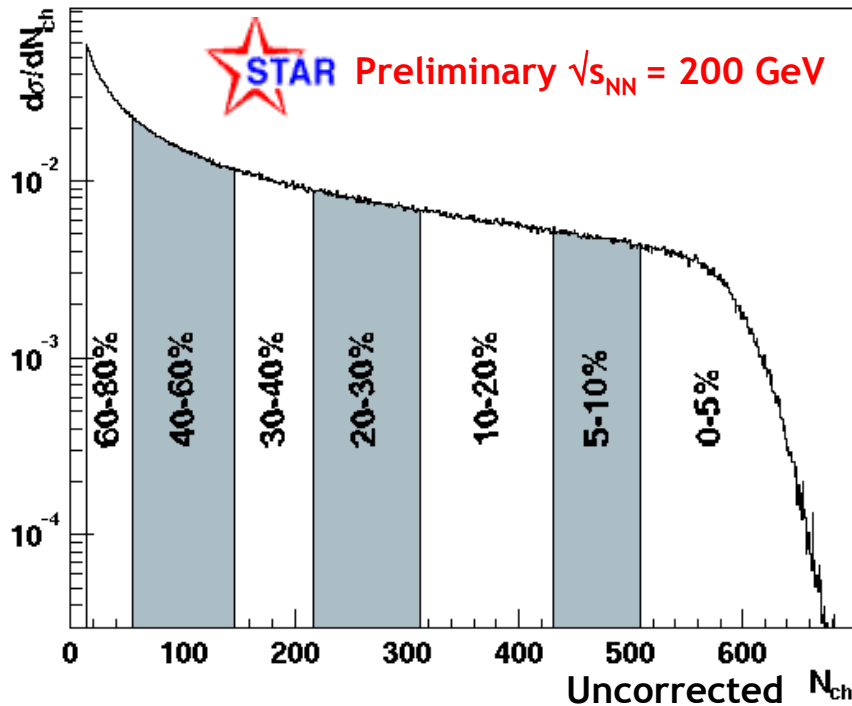
Event Selection



- Centrality selection : Used charge sum of Beam-Beam Counter (BBC, $|\eta|=3\sim 4$) and energy of Zero-degree calorimeter (ZDC) in minimum bias events.
- Extracted N_{part} based on Glauber model.

Au+Au Analysis

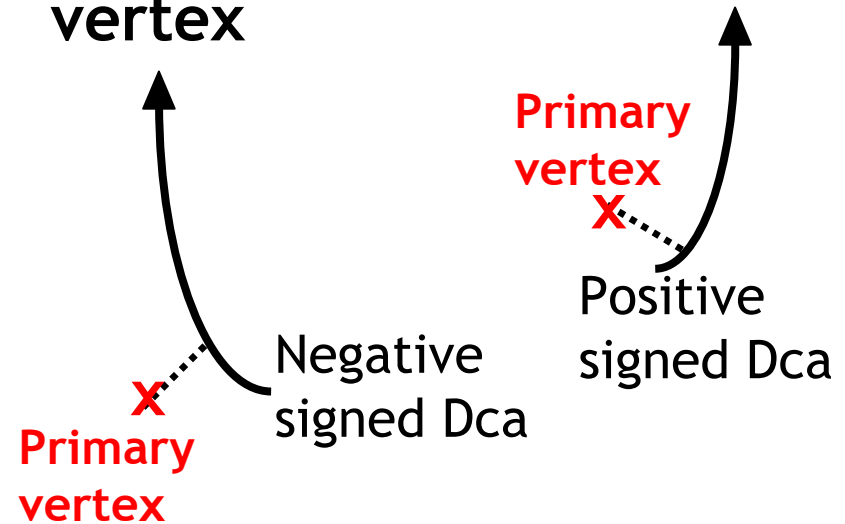
Event Selection:



Centrality classes based on mid-rapidity multiplicity

High p_T Track Selection:

dca \equiv distance of closest approach to the primary vertex



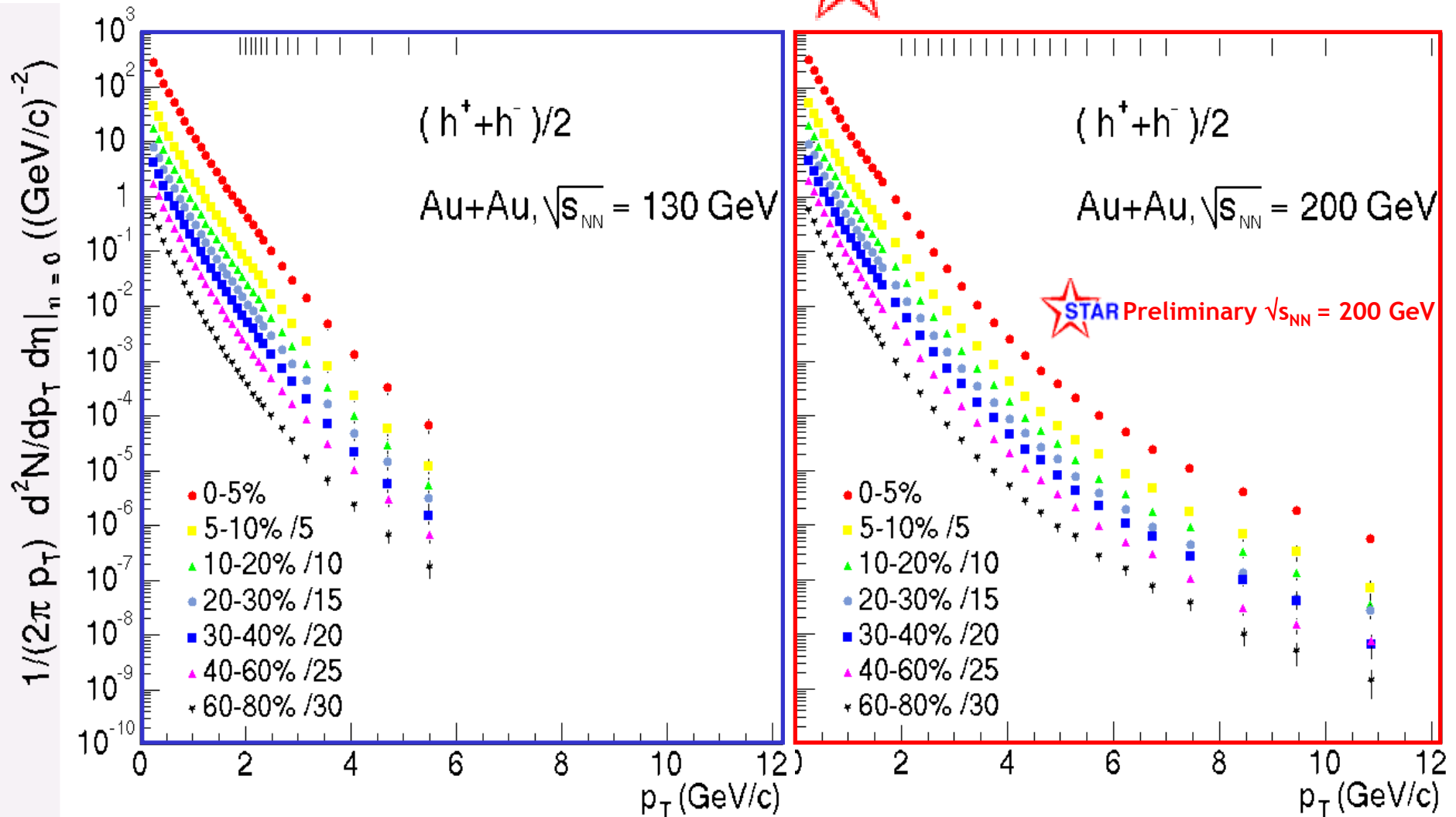
$$|\text{signed Dca}| < 1 \text{ cm}$$

$$|\eta| < 0.5$$

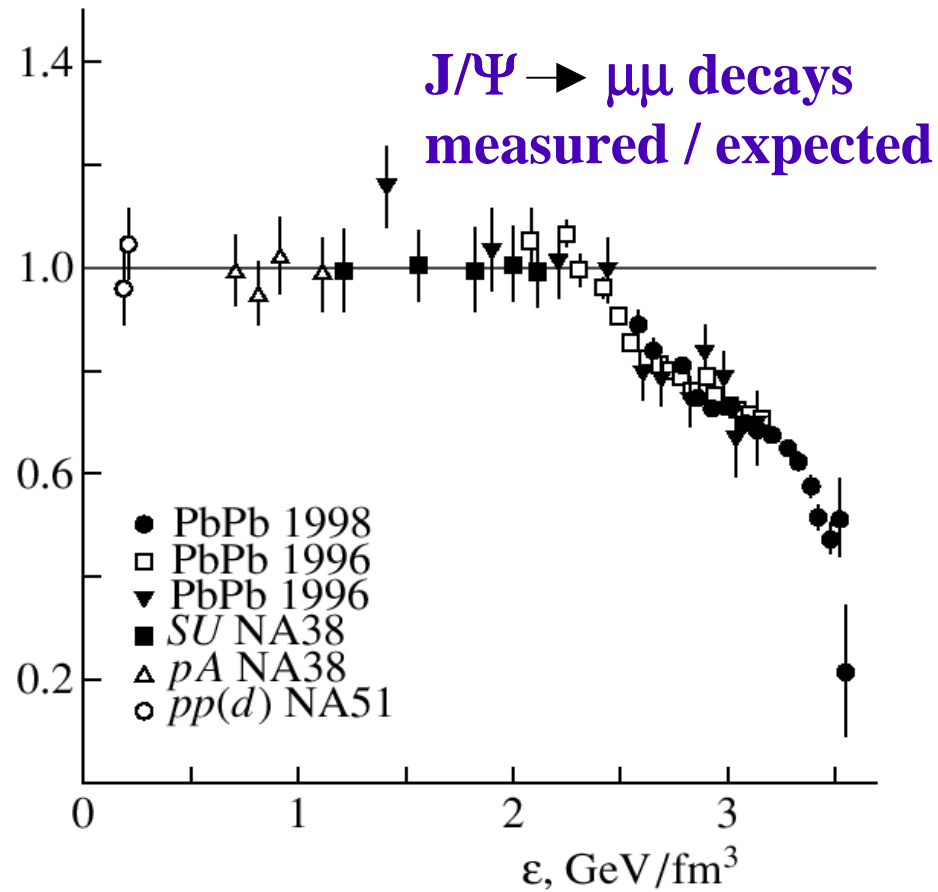
Charged hadron p_t spectra

130 GeV nucl-ex/0206011

 STAR Preliminary $\sqrt{s_{NN}} = 200$ GeV



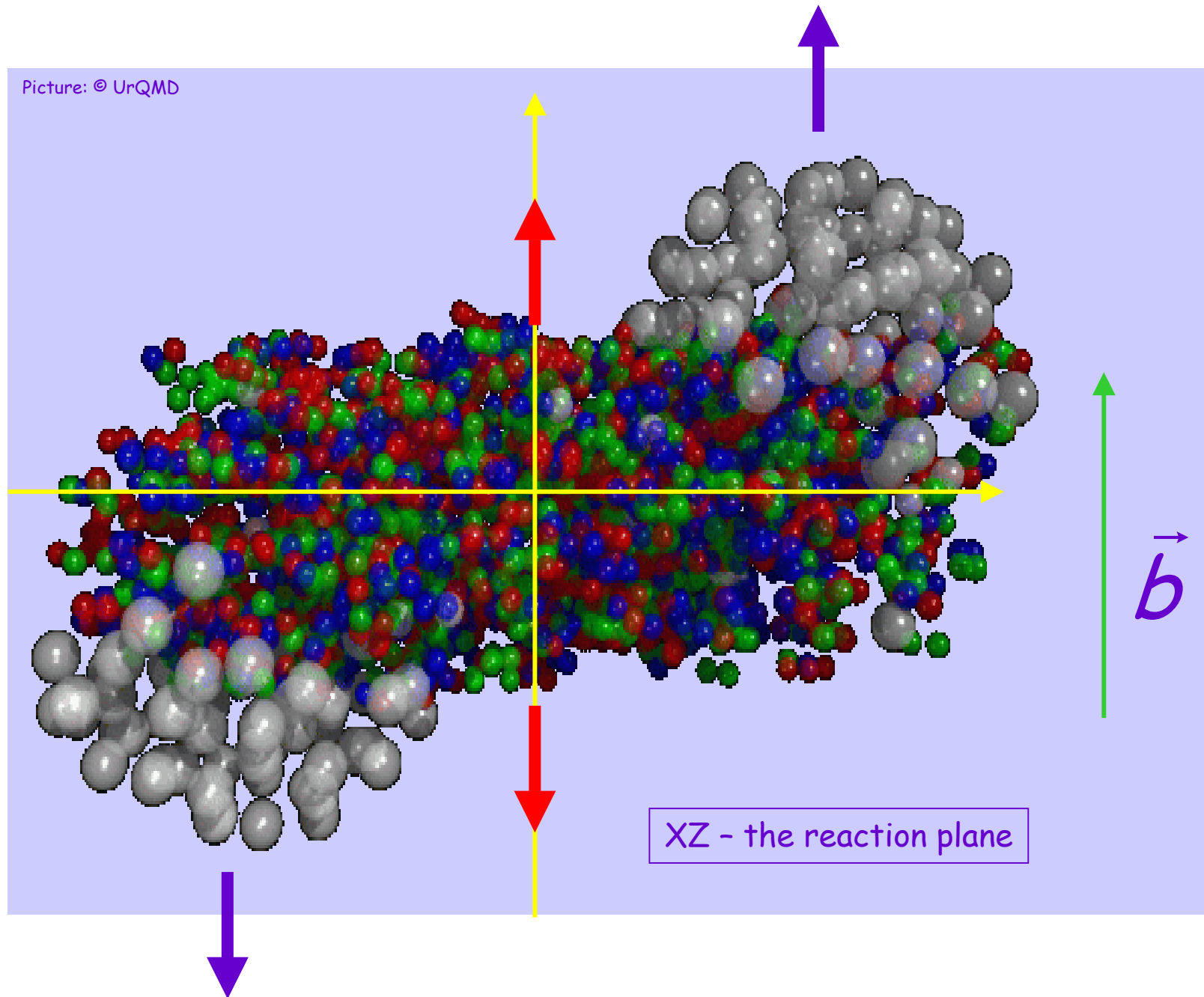
CERN SPS results (NA50, NA38, NA51)
(eg. Physics Of Atomic Nuclei 65, 325 (2002))

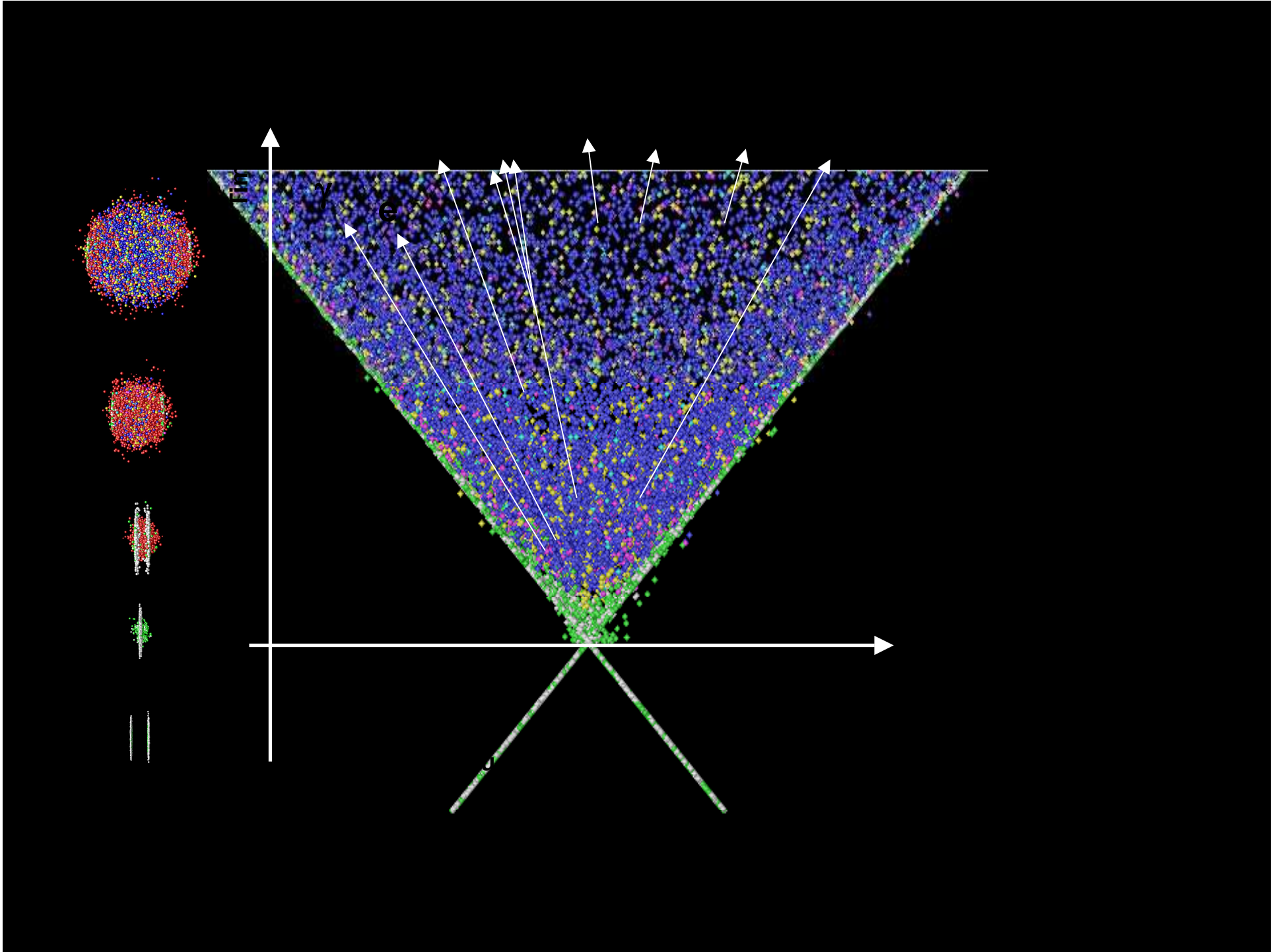


Expected J/ψ yields are corrected for "ordinary" nuclear absorption assuming **6.4 mb** absorption cross section

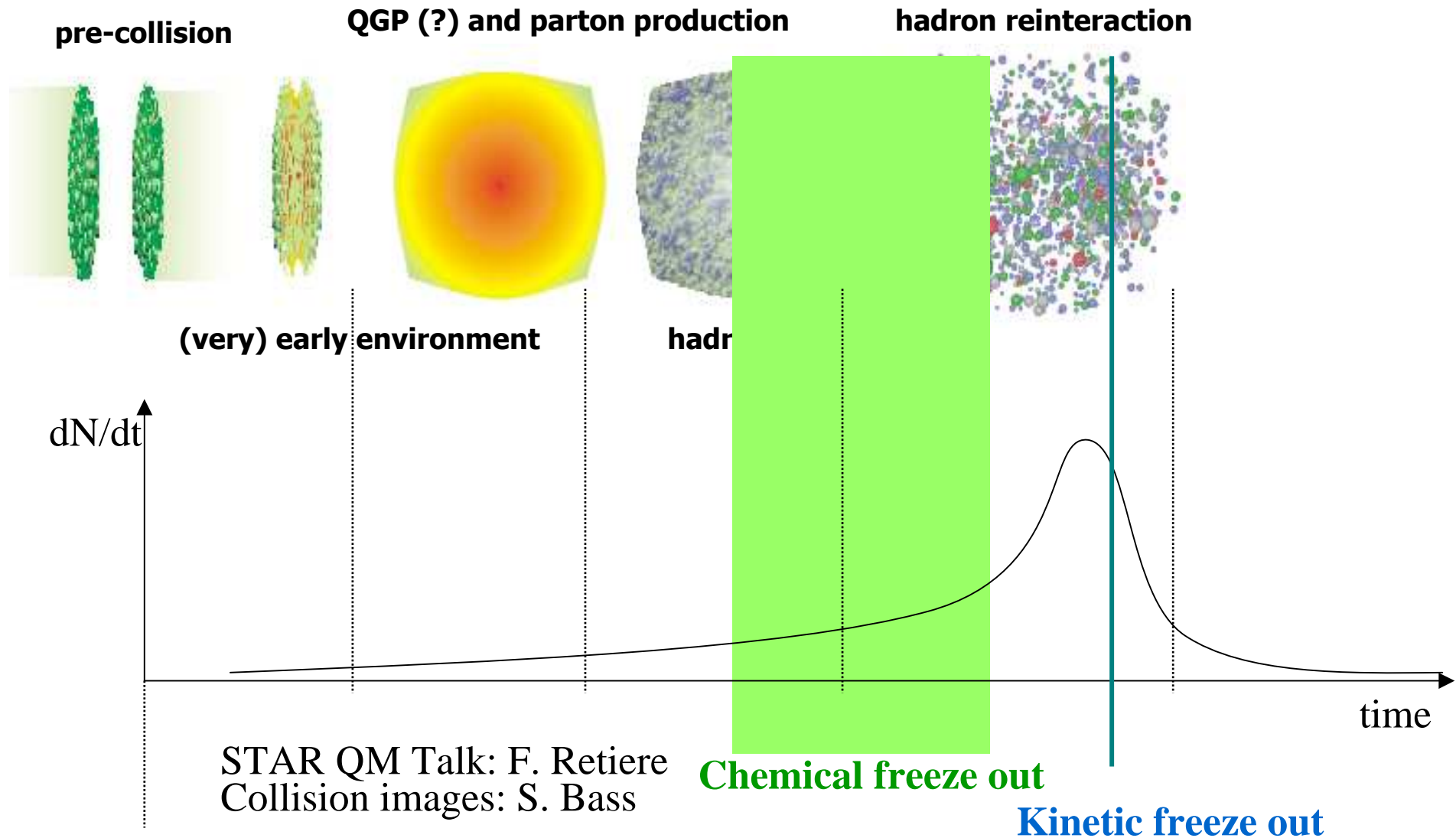
Many arguments about whether this is strong evidence of QGP at SPS energies, but it is clearly VERY interesting!

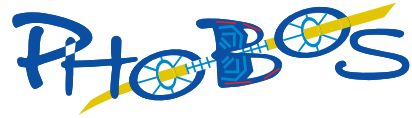
Picture: © UrQMD





Outline of a heavy ion collision





Collaboration



[ARGONNE NATIONAL LABORATORY](#)
[BROOKHAVEN NATIONAL LABORATORY](#)

[INSTITUTE OF NUCLEAR PHYSICS, KRAKOW](#)

[MASSACHUSETTS INSTITUTE OF TECHNOLOGY](#)

[NATIONAL CENTRAL UNIVERSITY, TAIWAN](#)

[UNIVERSITY OF ILLINOIS AT CHICAGO](#)

[UNIVERSITY OF MARYLAND](#)

[UNIVERSITY OF ROCHESTER](#)



Birger Back, Alan Wuosmaa

Mark Baker, Donald Barton, Alan Carroll, Nigel George, Stephen Gushue, George Heintzelman, Burt Holzman, Robert Pak, Louis Remsberg, Peter Steinberg, Andrei Sukhanov

Andrzej Budzanowski, Roman Hołyński, Jerzy Michałowski, Andrzej Olszewski, Pawel Sawicki, Marek Stodulski, Adam Trzupek, Barbara Wosiek, Krzysztof Woźniak

Maartin Ballintijn, Wit Busza (Spokesperson), Patrick Decowski, **Kristjan Gulbrandsen**, **Conor Henderson**, **Jay Kane**, Judith Katzy, Piotr Kulinich, **Jang Woo Lee**, Heinz Pernegger, **Corey Reed**, Christof Roland, Gunther Roland, Leslie Rosenberg, **Pradeep Sarin**, Stephen Steadman, George Stephans, **Carla Vale**, Gerrit van Nieuwenhuizen, Gábor Veres, Robin Verdier, Bernard Wadsworth, Bolek Wysłouch

Chia Ming Kuo, Willis Lin, Jaw-Luen Tang

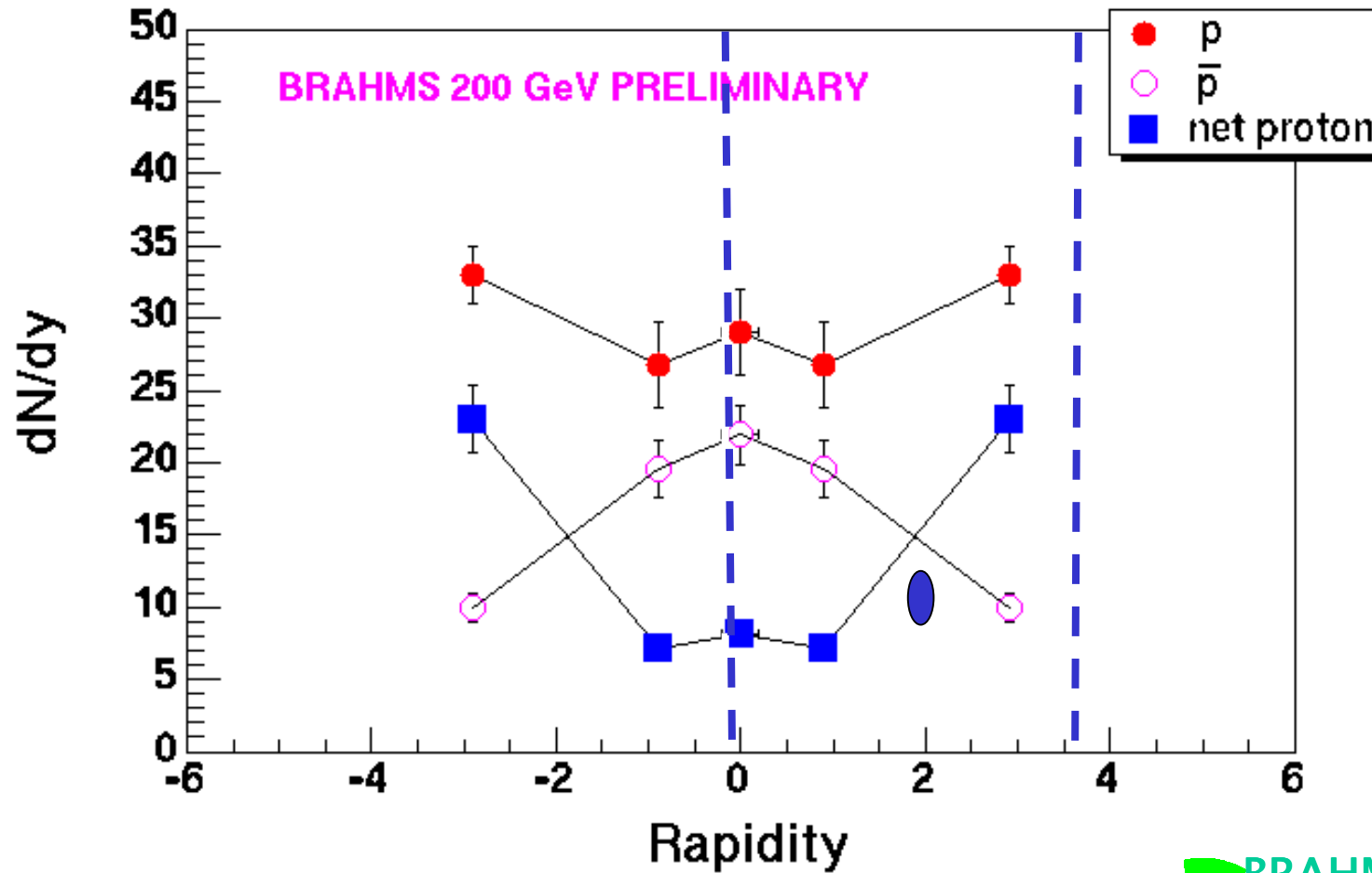
Russell Betts, Edmundo Garcia, Clive Halliwell, David Hofman, **Richard Hollis**, **Aneta Iordanova**, Wojtek Kucewicz, Don McLeod, Rachid Nouicer, Michael Reuter, **Joe Sagerer**

Abigail Bickley, **Richard Bindel**, Alice Mignerey, Marguerite Belt Tonjes

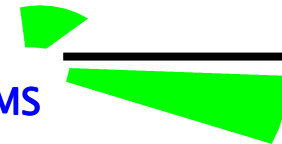
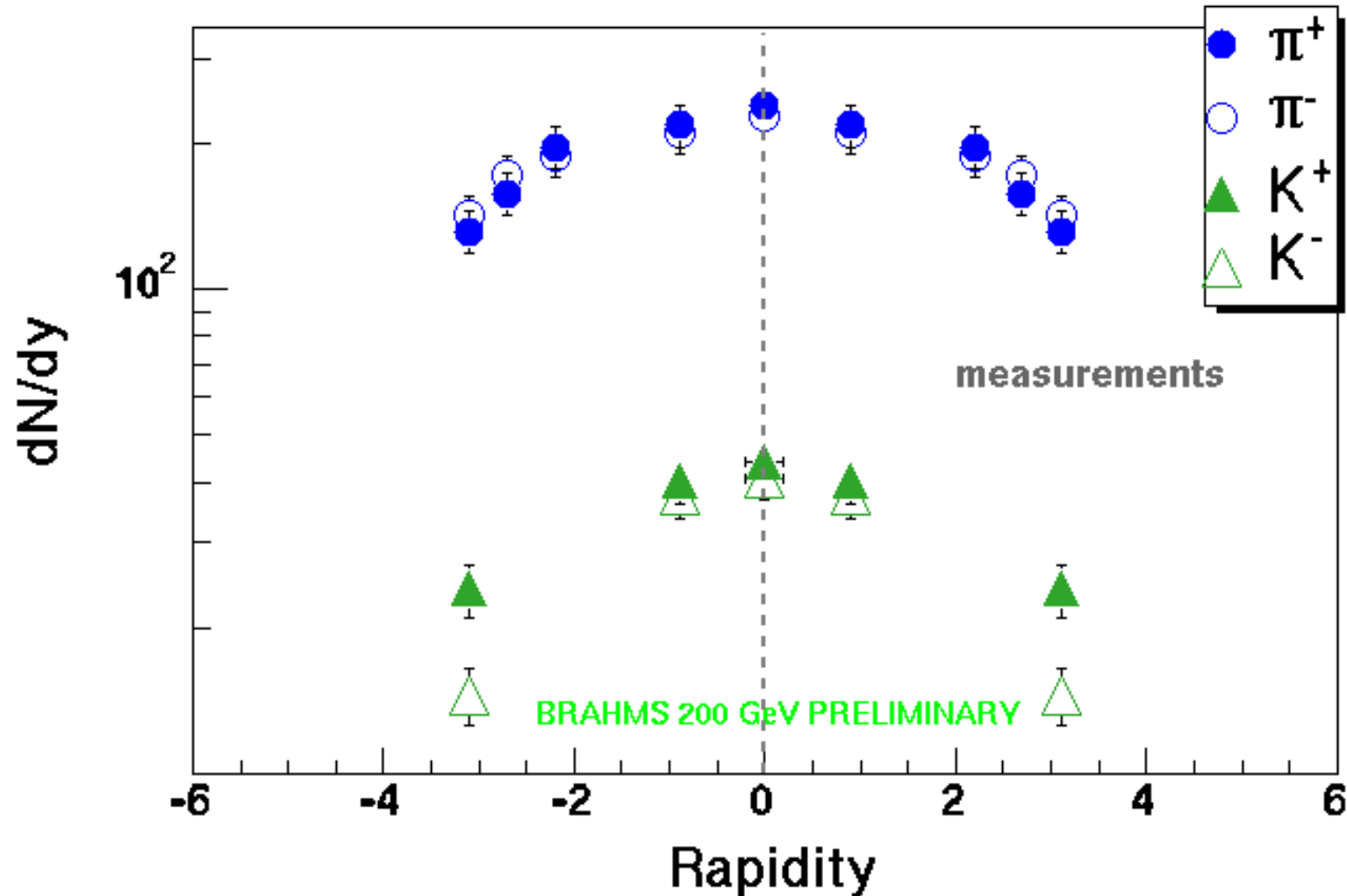
Joshua Hamblen, **Erik Johnson**, **Nazim Khan**, Steven Manly, Inkyu Park, Wojtek Skulski, Ray Teng, Frank Wolfs



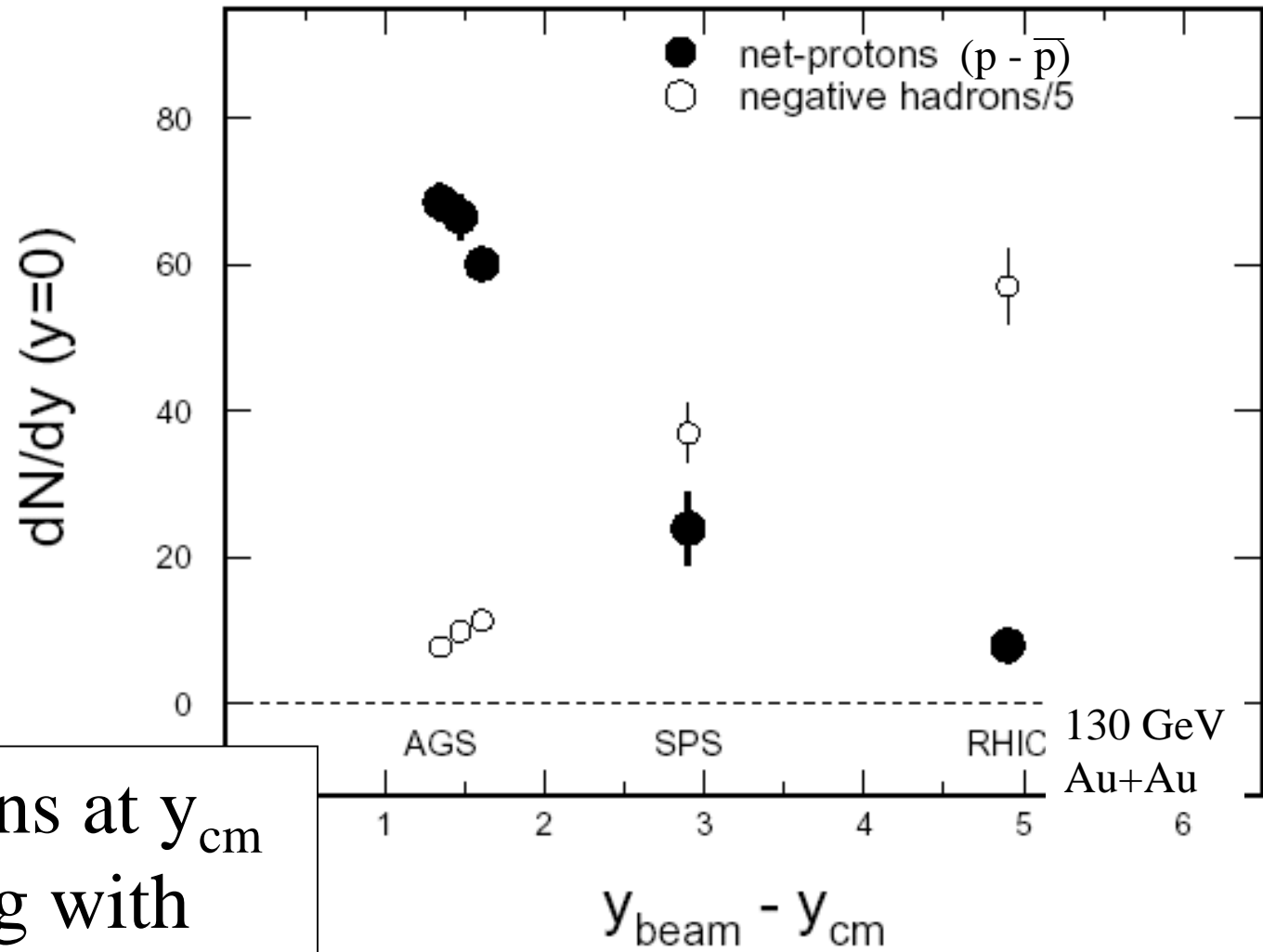
Net protons vs rapidity at RHIC



Rapidity density at RHIC (0-10% central)



Net protons



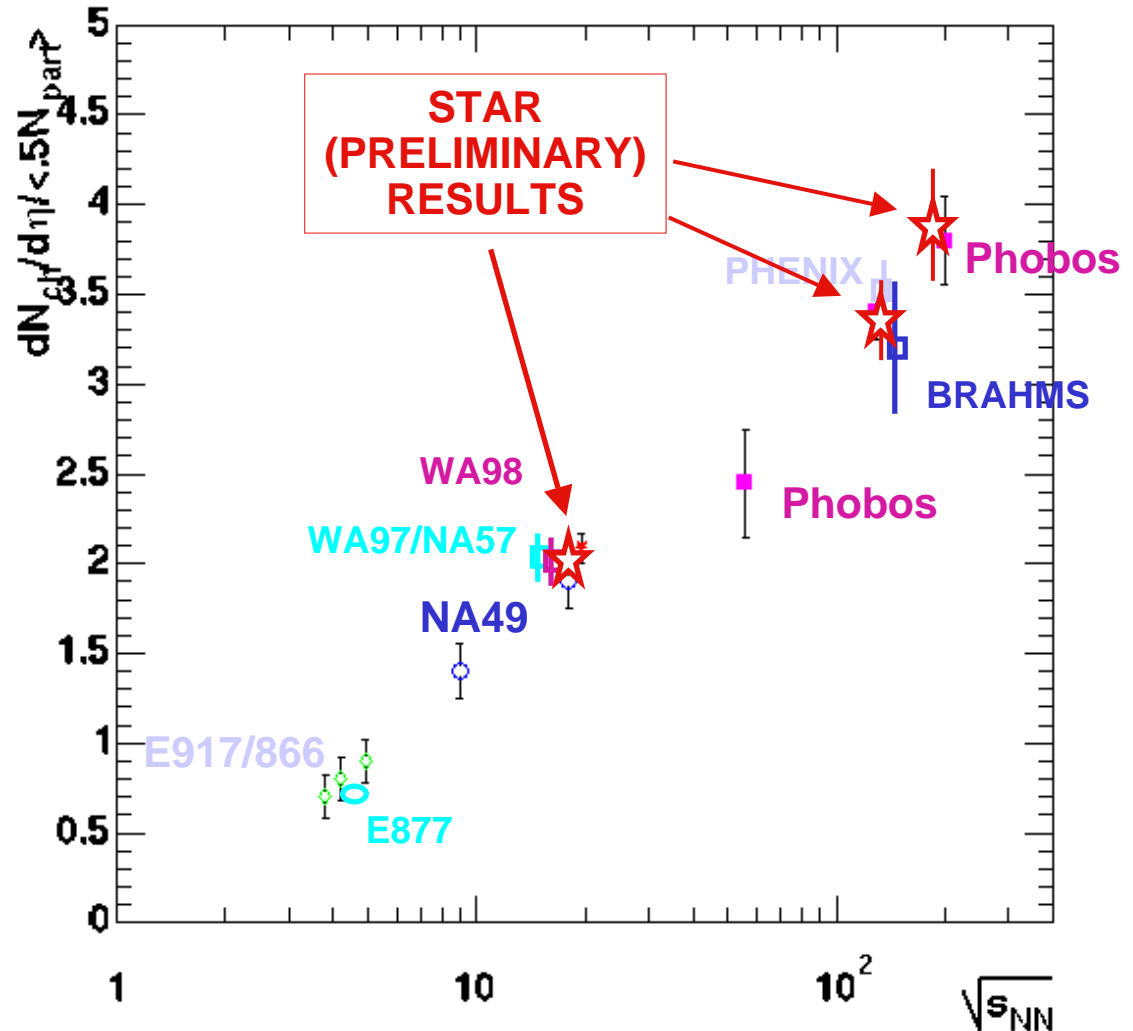
Net protons at y_{cm}
decreasing with
increasing energy

$dN/d\eta$ vs. collision energy

~5% central Au+Au (Pb+Pb) collisions

Phobos PRL 85 (2000) 3100
Phobos nucl-ex 018009 (2002)
PHENIX PRL 86 (2001) 3500
STAR(130) nucl-ex 106004 (2001)
BRAHMS QM2001
NA49
WA98 nucl-ex 0008004 (2000)
WA97/NA57 CERN-EP-2000-002
E866/917 PRC59 (1999) 2173
E877 PRC51 (1995) 3309

STAR QM Posters on
19.6 GeV data: D. Cebra
200 GeV data: M. Calderon



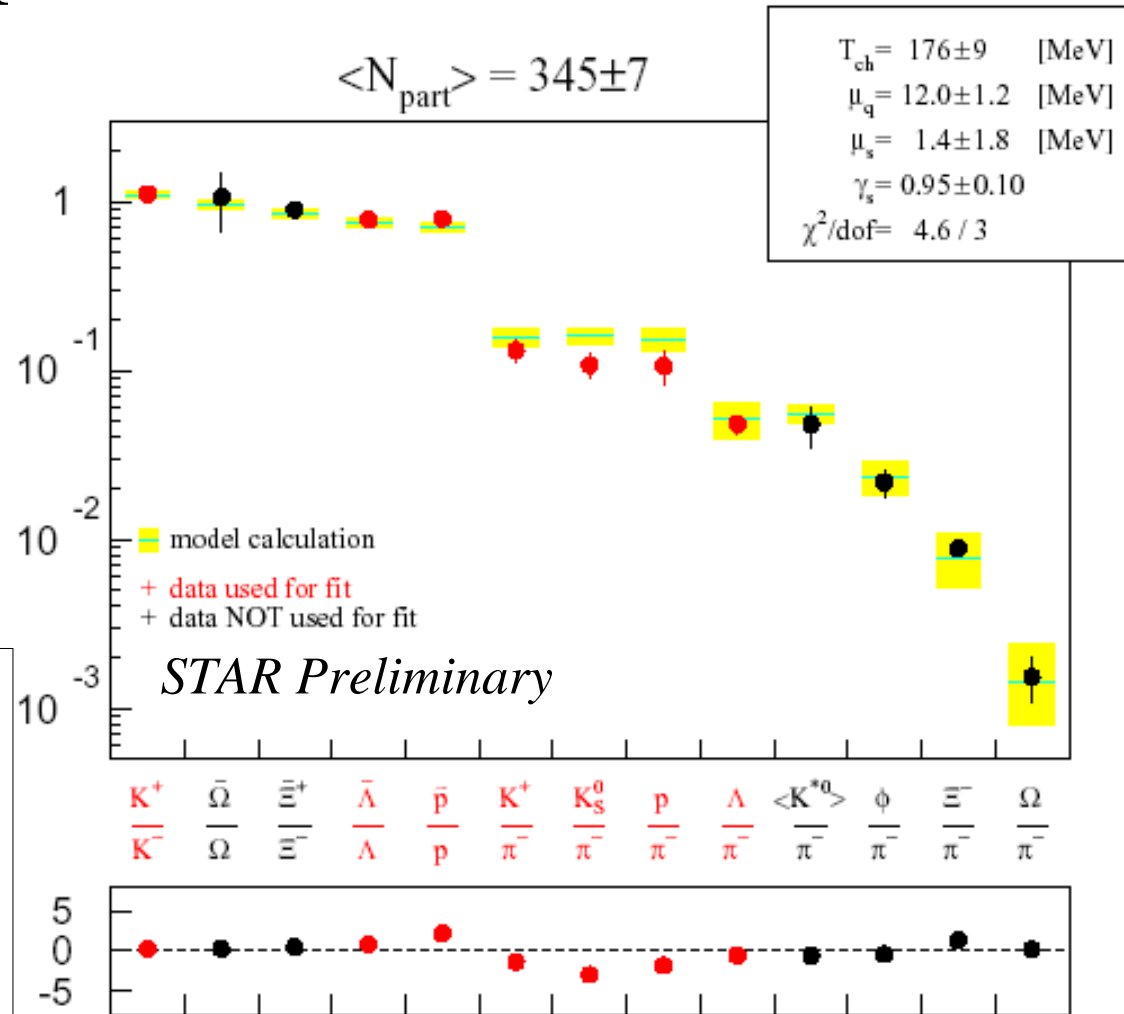
Ratios, experiment vs. a model

Central
130 GeV Au+Au
Preliminary Data

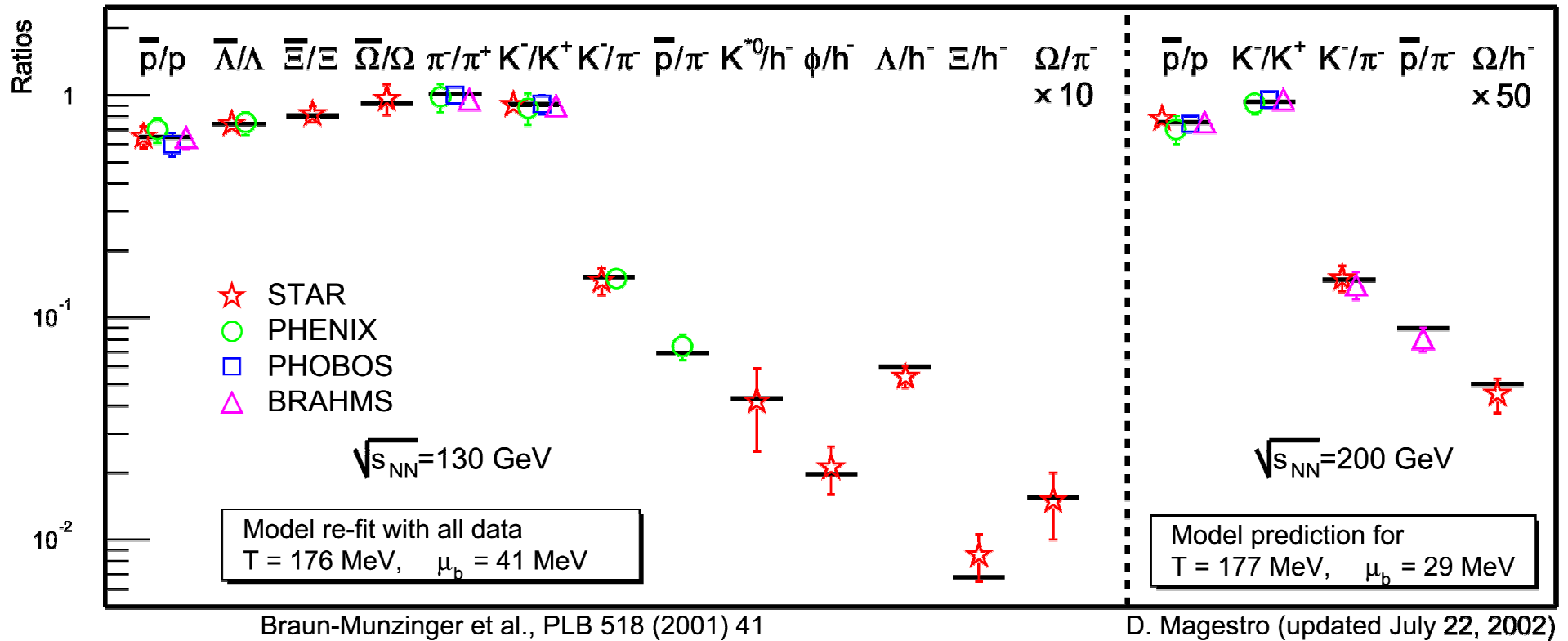
Agreement between
model and data is
very good!

Thermal fits to mid-rapidity
spectra have caveats regarding
non- 4π measurements (local
vs. global equilibrium, boost-
invariance).

M. Kaneta and N. Xu,
J. Phys. G27 (2001) 589



Statistical Model: First Look at AuAu @ 200 GeV



- **All 200 GeV data taken from QM talks:**
- F. Wang (STAR)/G. Van Buren (STAR)/
- T. Chujo (PHENIX)/Ouerdane (BRAHMS)
- J. Lee (BRAHMS)/B. Wosiek (PHOBOS)
- **New 130 GeV data are:**
- C. Suire (STAR)/J. Castillo (STAR)

Predictions:

phenomenologically:

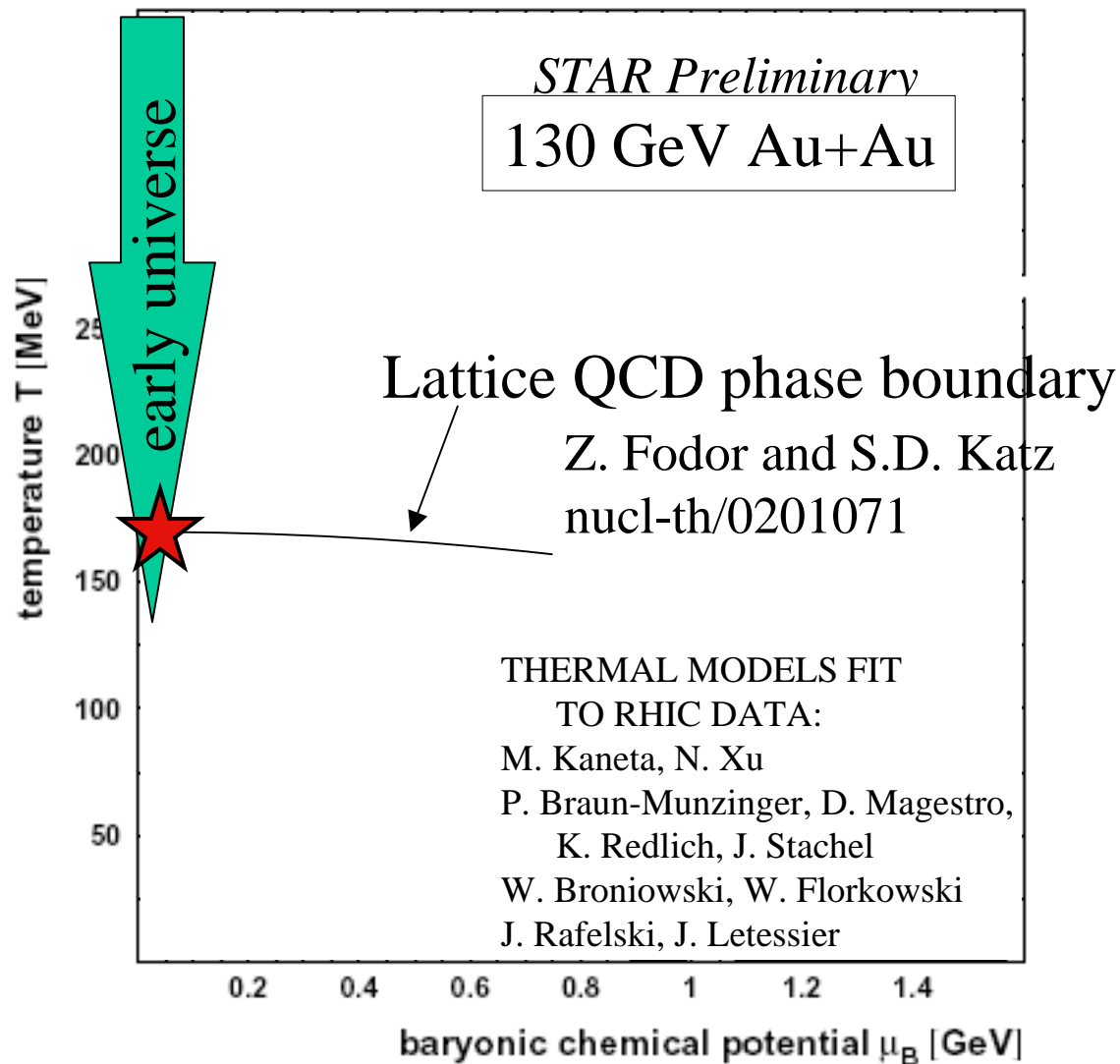
$$\mu_B \sim 1.3 \text{ GeV} (1 + \sqrt{s}/4.5 \text{ GeV})^{-1}$$

assume unified freeze-out condition:

$$\langle E \rangle / \langle N \rangle \sim 1.1 \text{ GeV} \Rightarrow T$$

Where are we?...Are we there (yet)?

- Chemical, thermal equilibrium description gives high T , low μ_B :
 - Consistent with region of phase boundary from lattice QCD!
 - Similar to early universe!

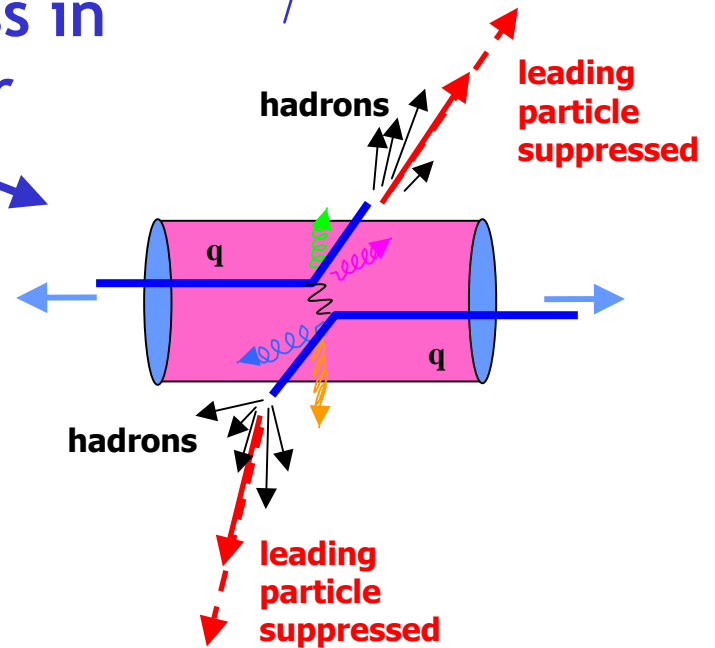
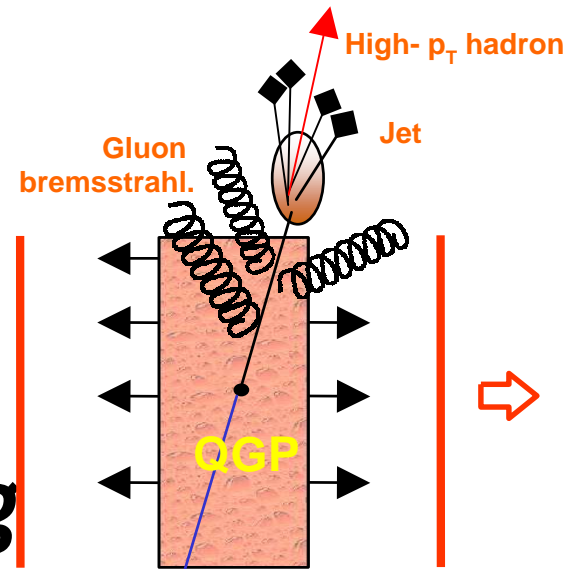
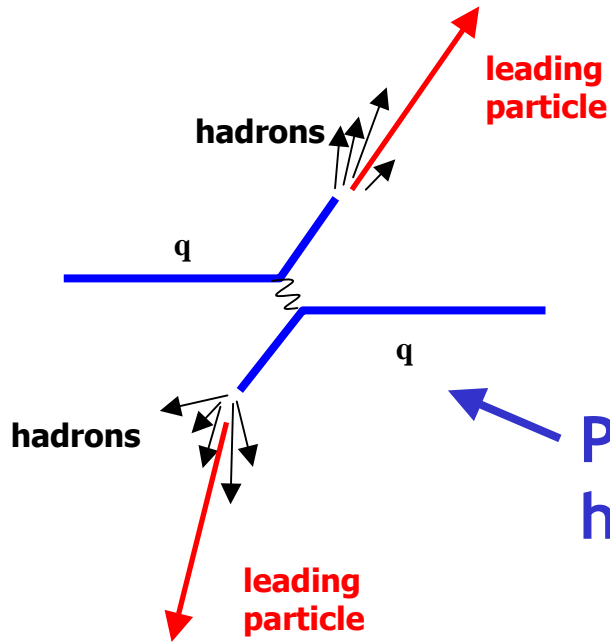


Leading Particle Suppression

Jets modified in heavy ion collisions?

Jet quenching

Partonic Energy loss in high density matter



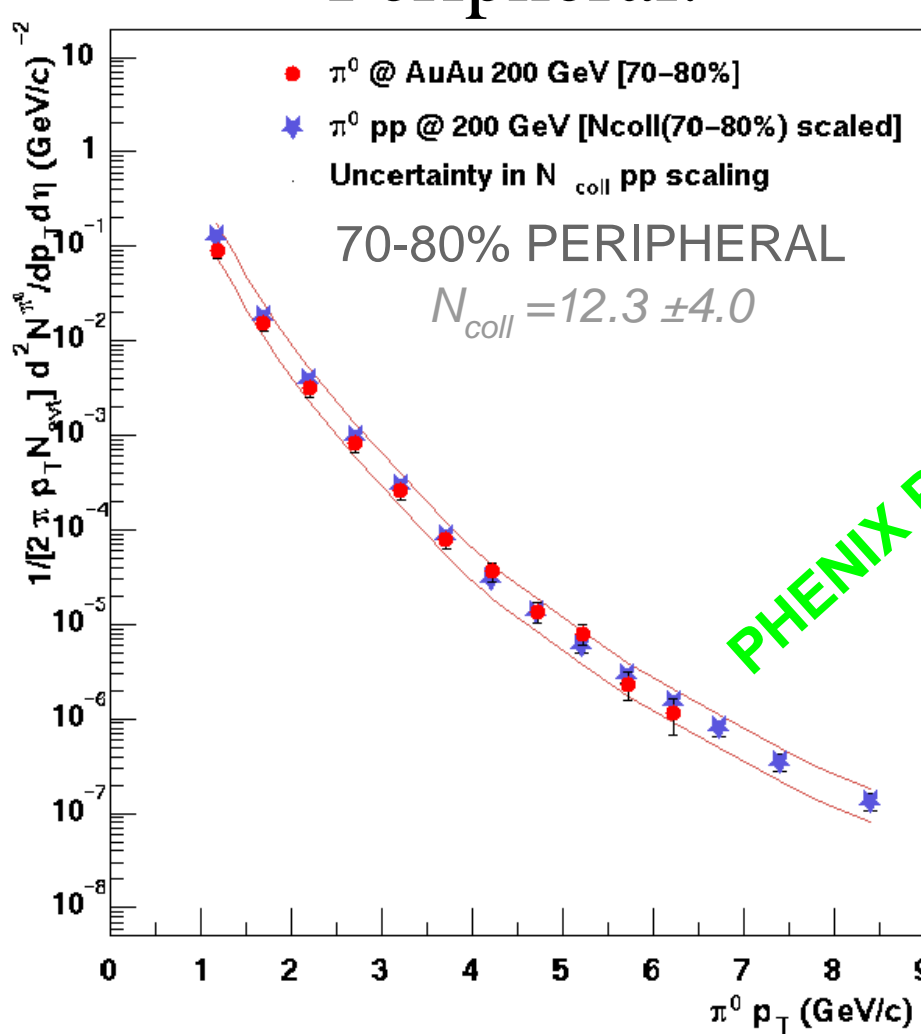
Nuclear Modification Factor:

$$R_{AA}(p_T) = \frac{d^2 N^{AA} / dp_T d\eta}{T_{AA} d^2 \sigma^{NN} / dp_T d\eta}$$

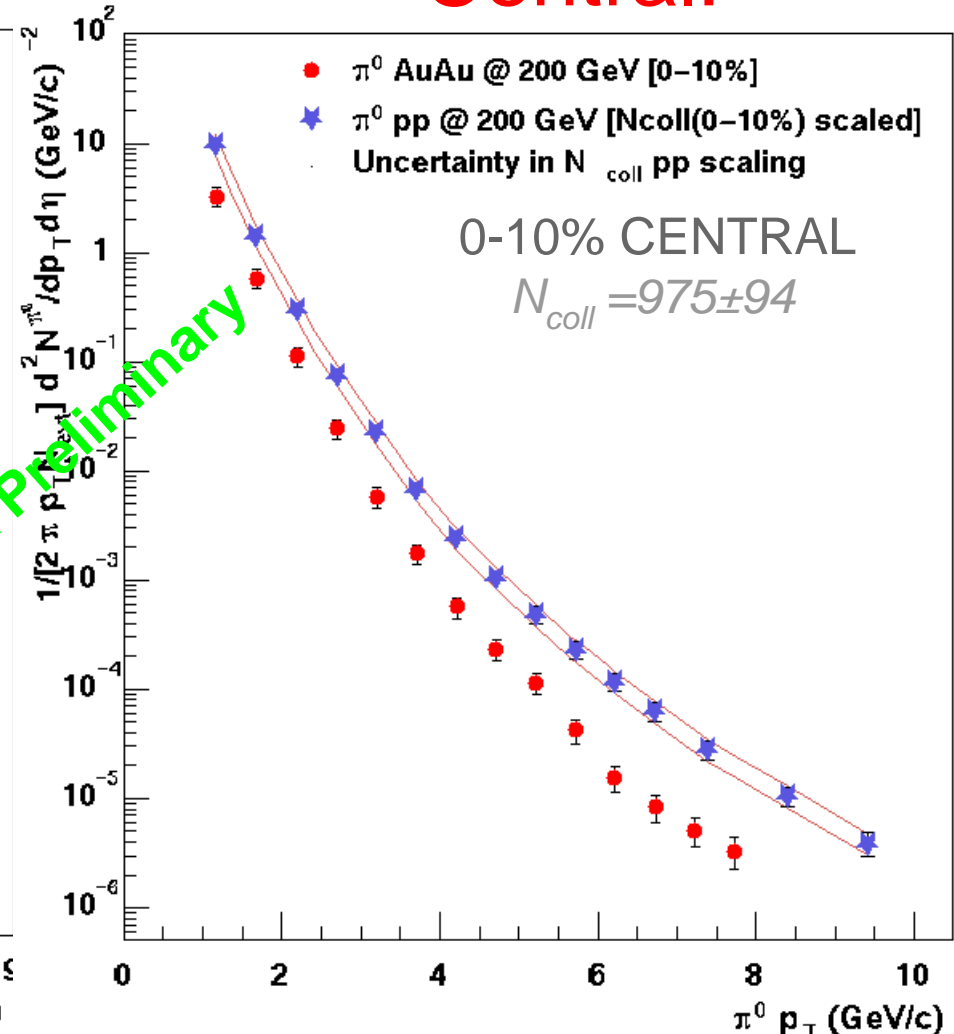
Glauber Model

The results : pp vs AuAu (peripheral & central)

• Peripheral:



• Central:

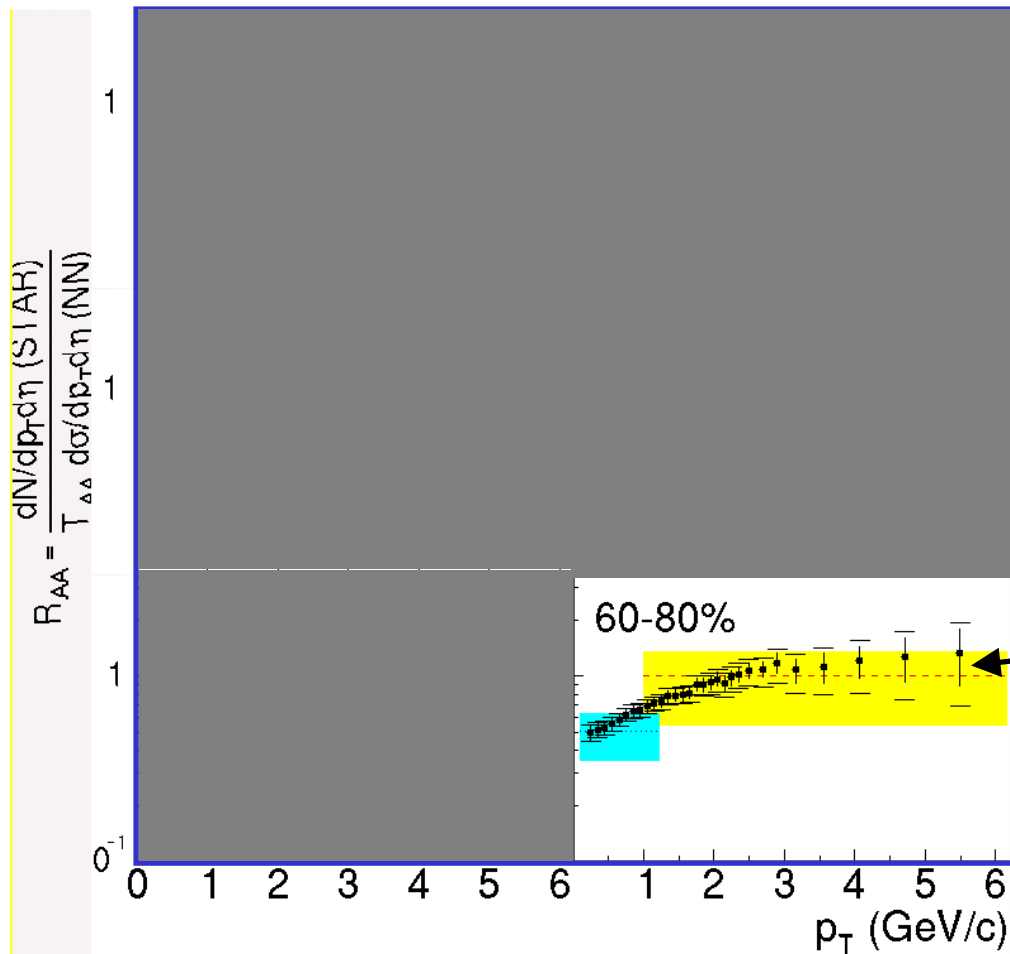


PHENIX Preliminary

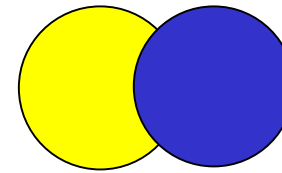
• Large suppression (increasing with p_T) above $p_T \sim 1$ GeV/c compared to scaled pp.

R_{AA} Comparison to $p_T = 6 \text{ GeV}/c$

130 GeV nucl-ex/0206011



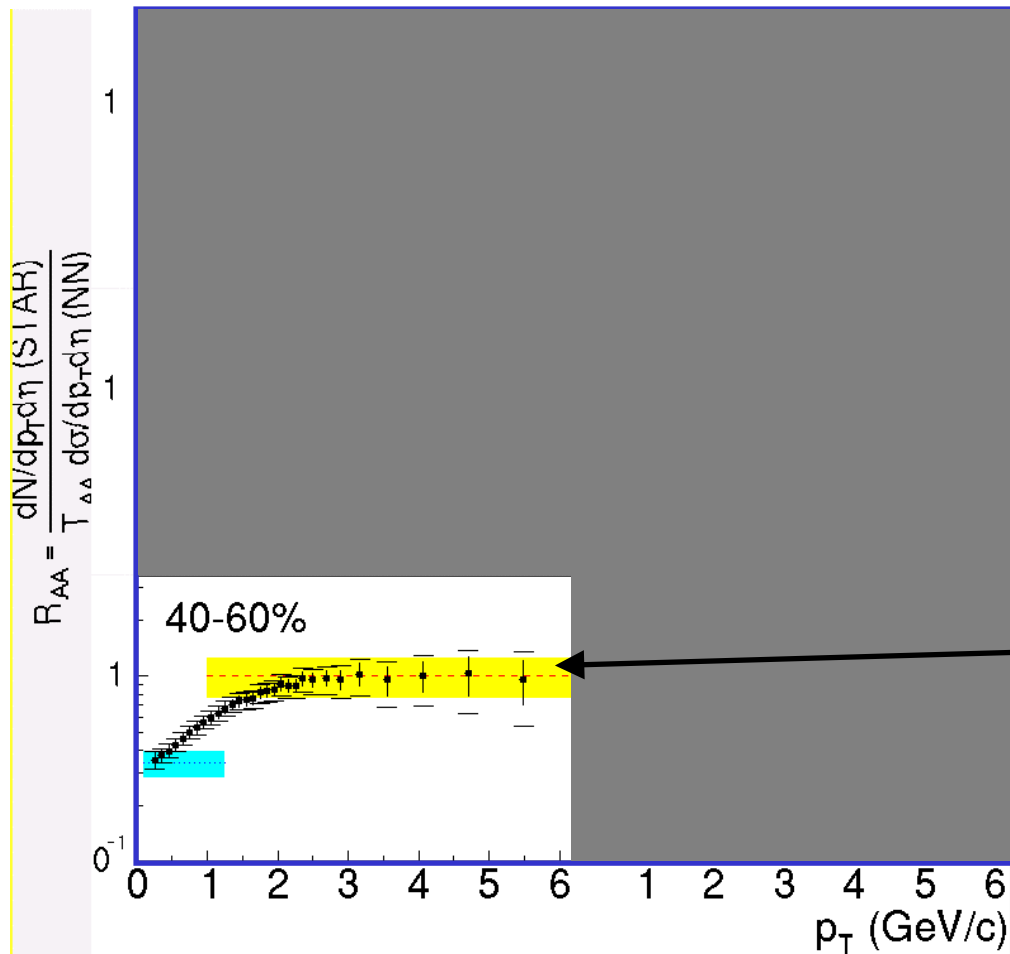
Centrality: 60-80%



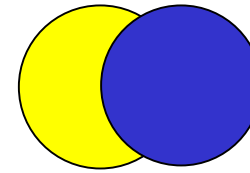
Peripheral R_{AA} scales as N_{binary} above $p_T = 2 \text{ GeV}$

R_{AA} Comparison to $p_T = 6 \text{ GeV}/c$

130 GeV nucl-ex/0206011



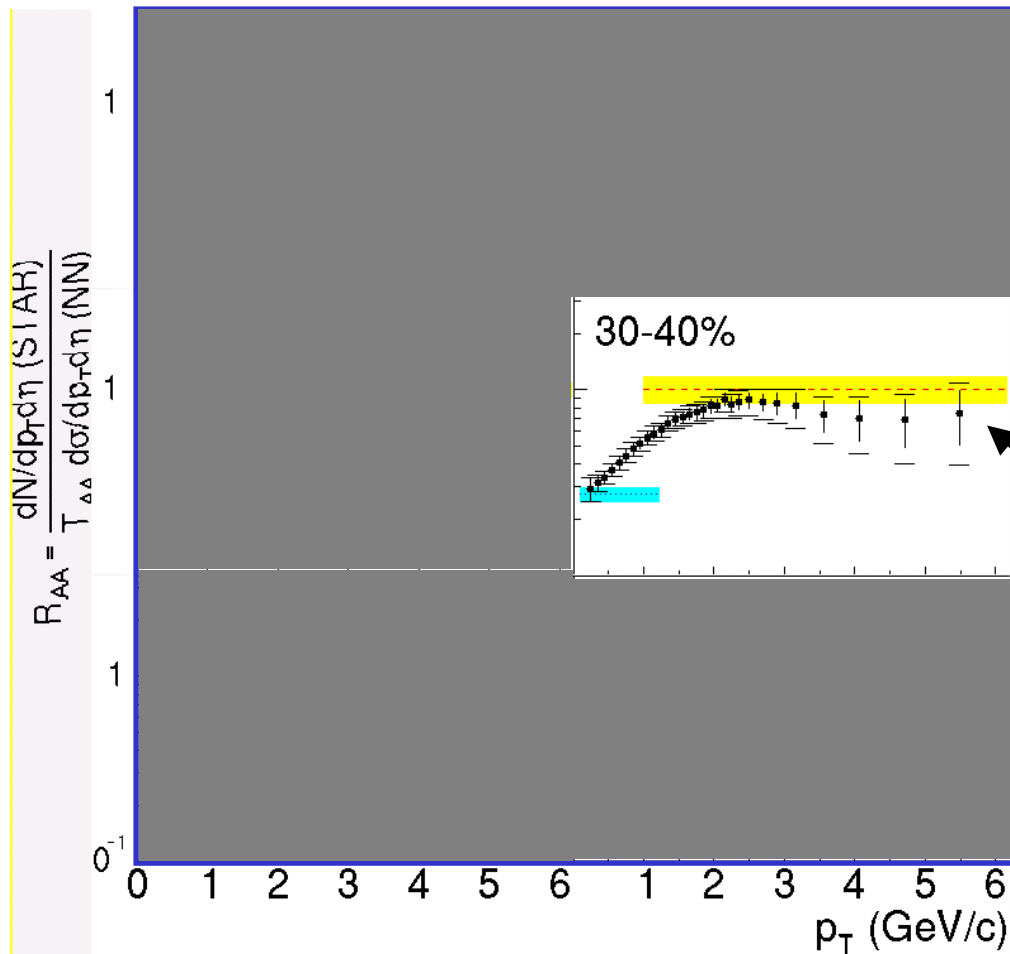
Centrality: 40-60%



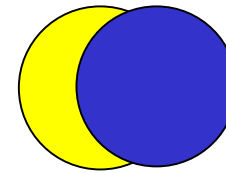
Peripheral R_{AA} scales as N_{binary} above $p_T = 2 \text{ GeV}$

R_{AA} Comparison to $p_T = 6 \text{ GeV}/c$

130 GeV nucl-ex/0206011



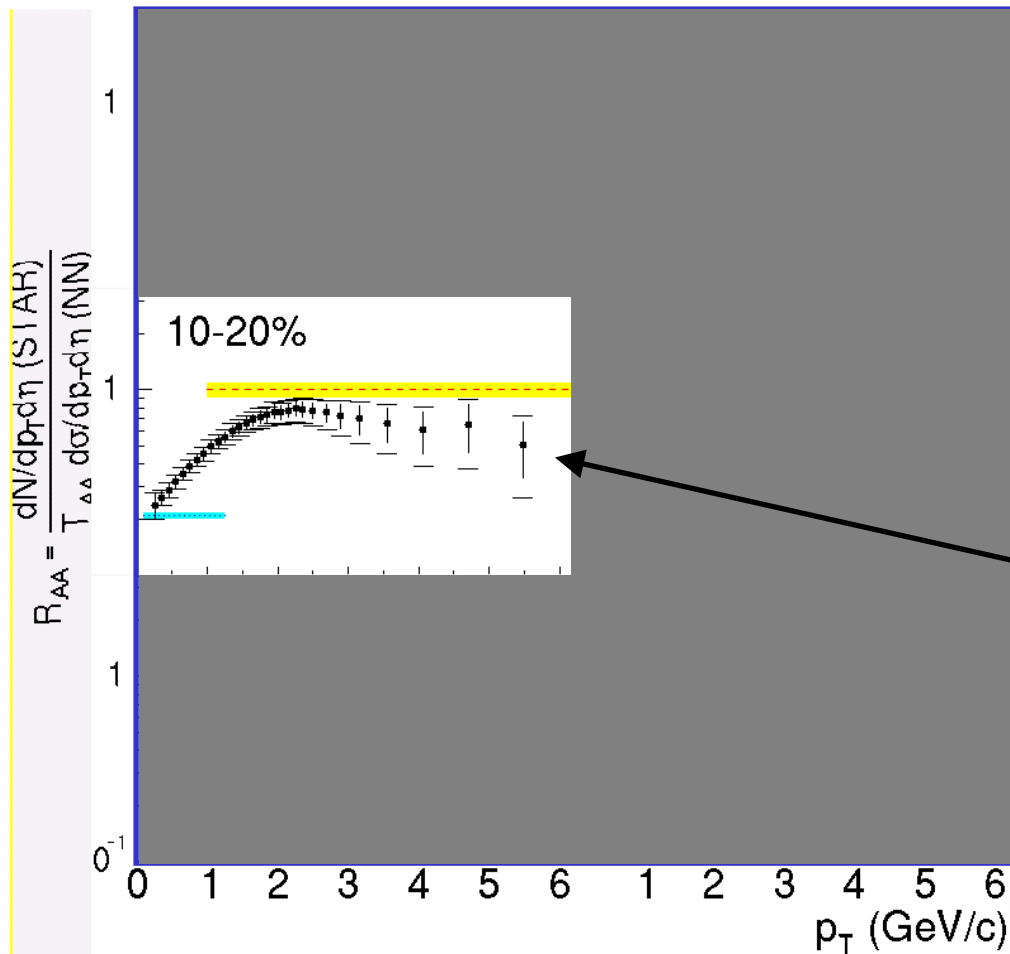
Centrality: 30-40%



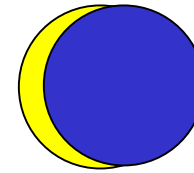
As centrality decreases,
suppression sets in
above $p_T = 2 \text{ GeV}$

R_{AA} Comparison to $p_T = 6 \text{ GeV}/c$

130 GeV nucl-ex/0206011



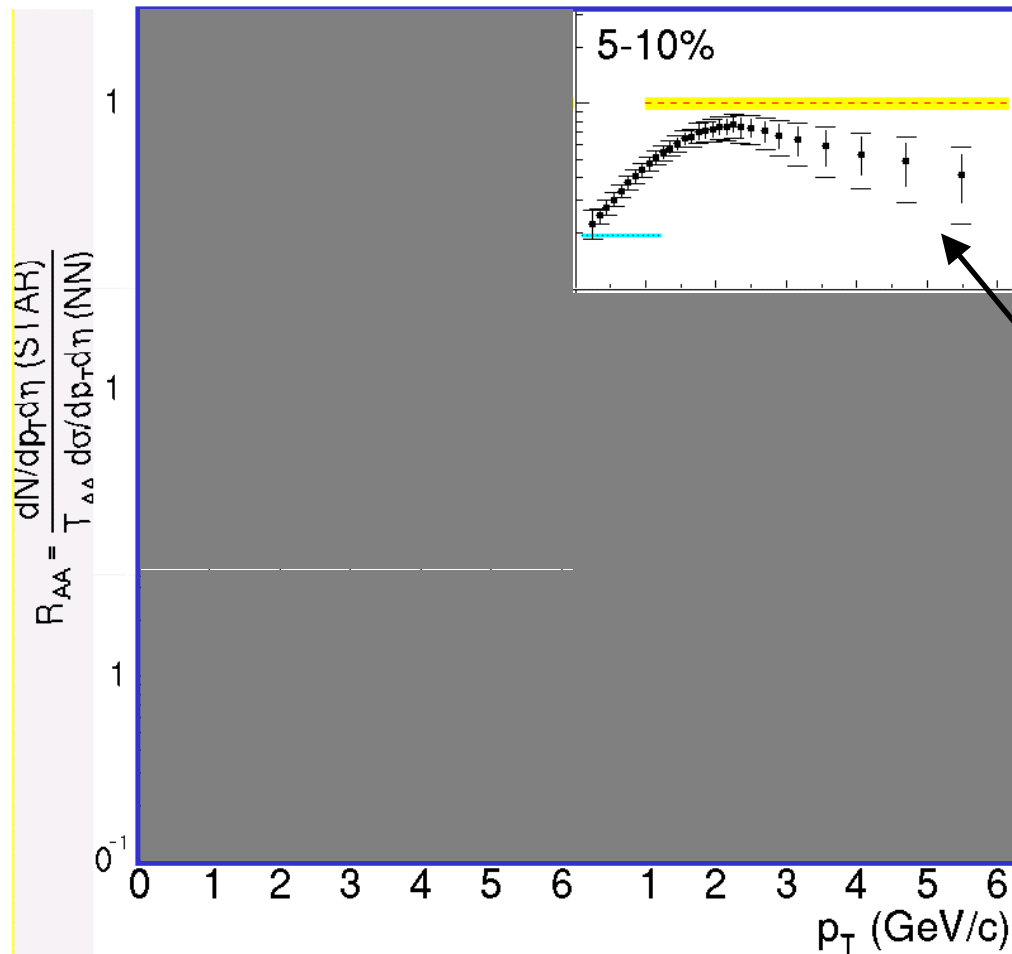
Centrality: 10-20%



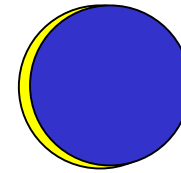
As centrality decreases,
suppression sets in
above $p_T = 2 \text{ GeV}$

R_{AA} Comparison to $p_T = 6 \text{ GeV}/c$

130 GeV nucl-ex/0206011



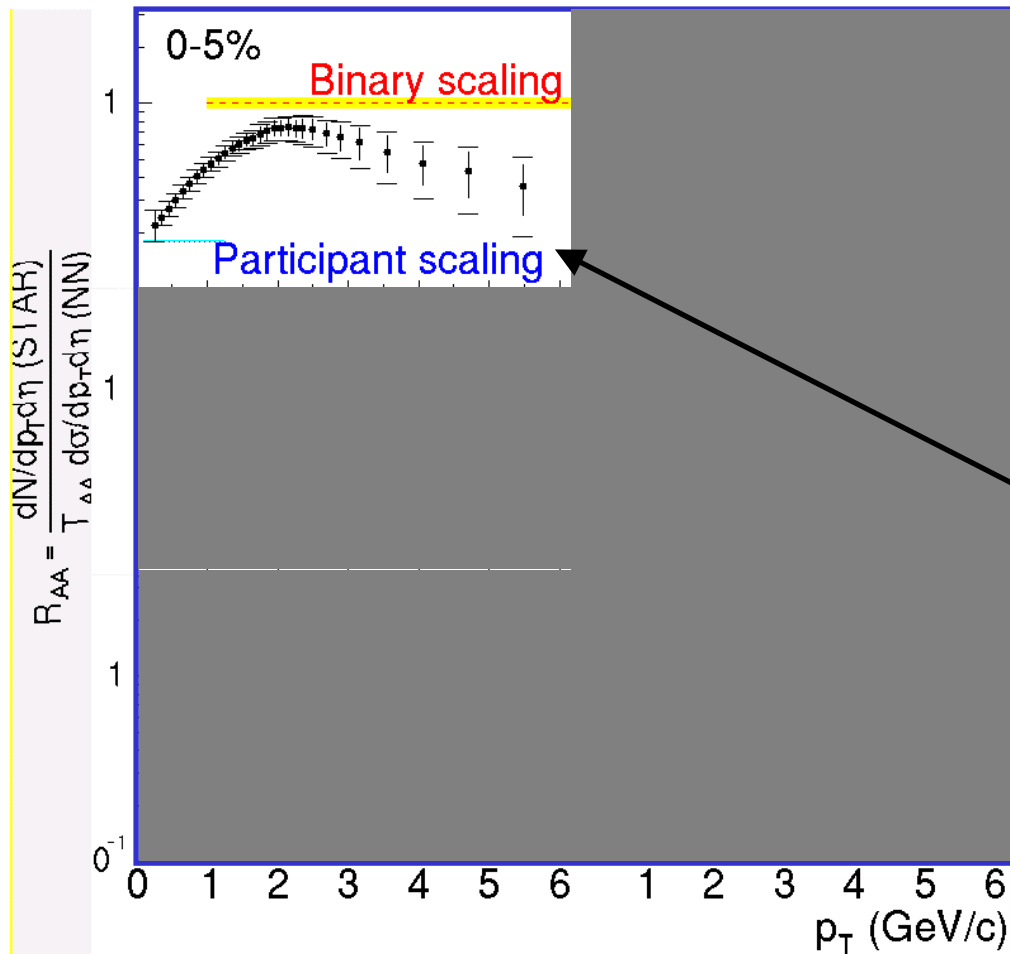
Centrality: 5-10%



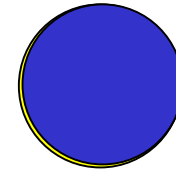
In the most central collisions, suppression is clear

R_{AA} Comparison to $p_T = 6 \text{ GeV}/c$

130 GeV nucl-ex/0206011

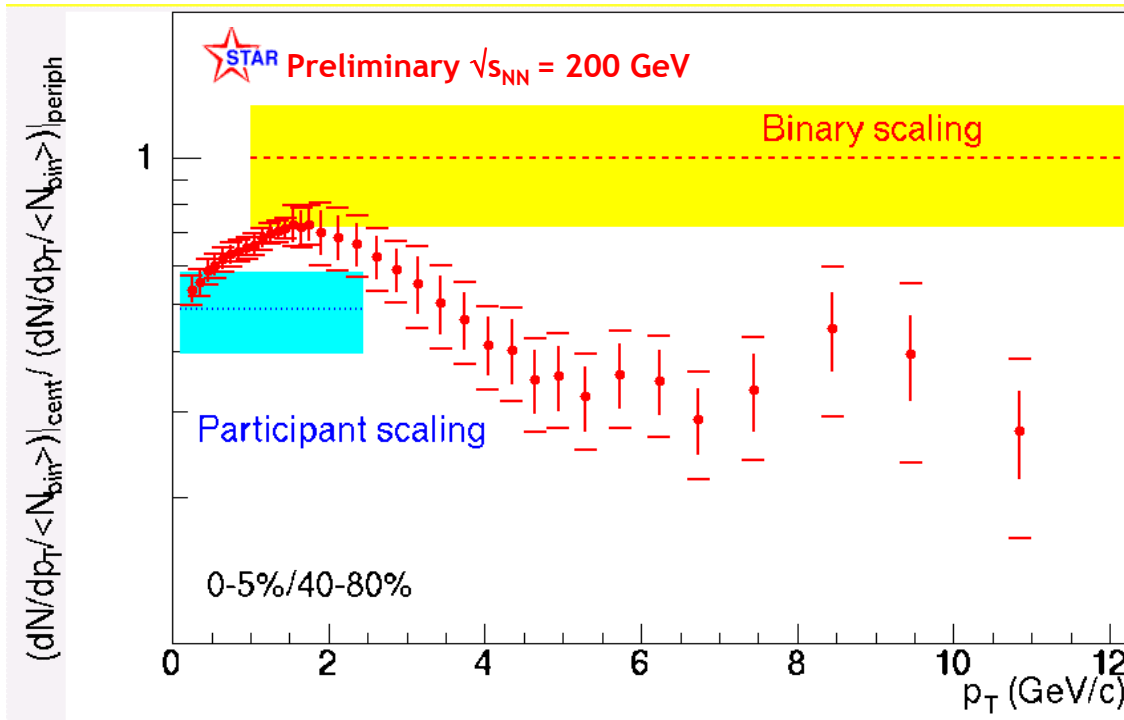


Centrality: 0-5%



In the most central collisions, suppression is strongest

Summary & Outlook

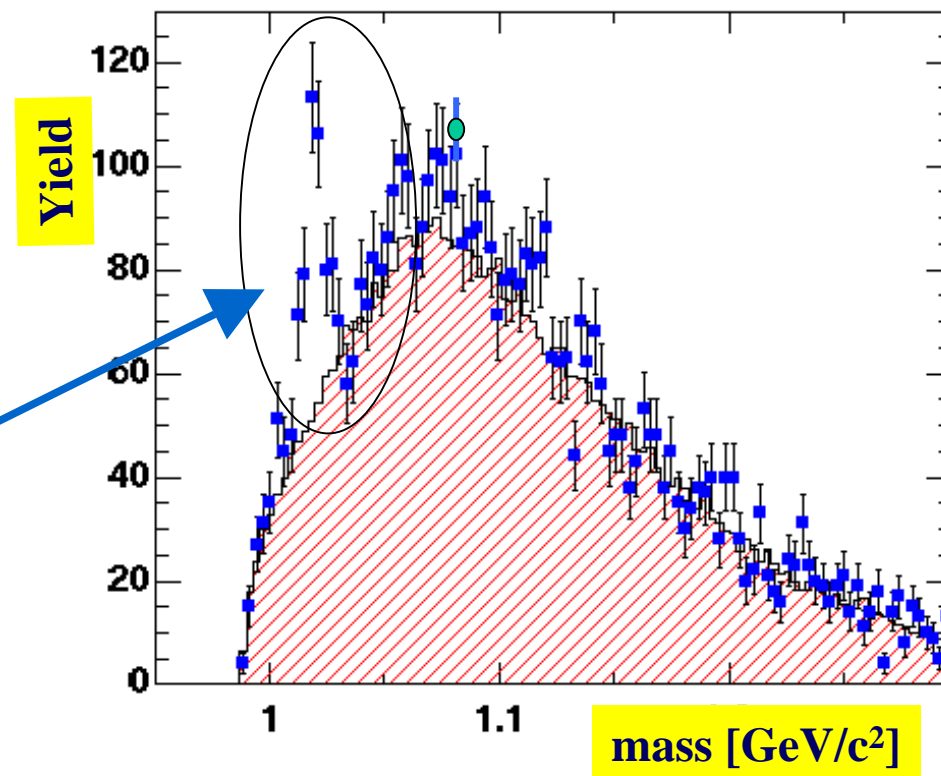
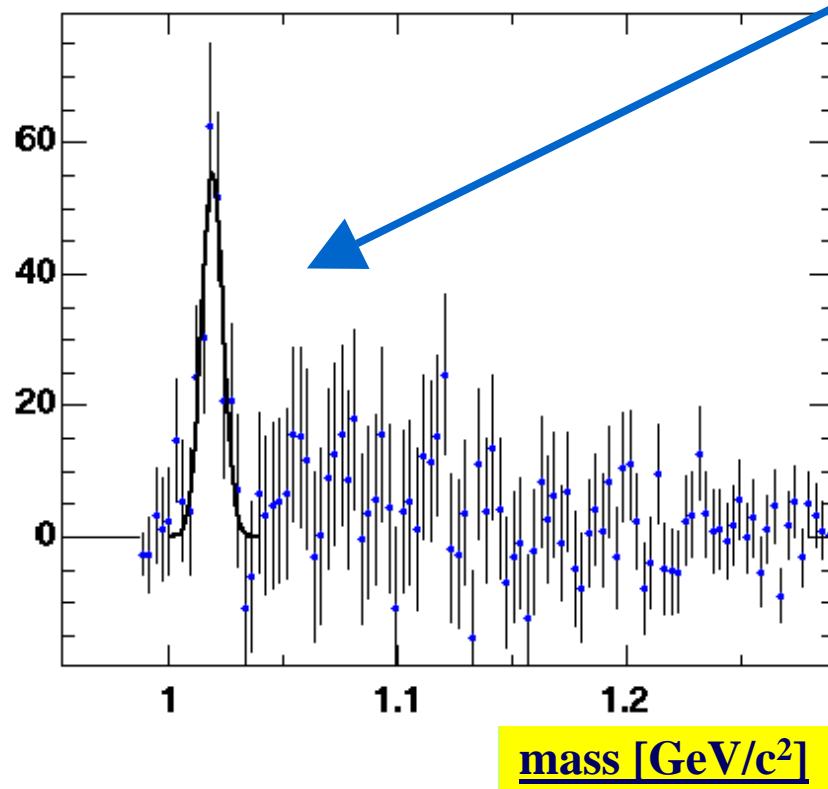


- High p_T hadron suppression saturates above $p_T = 6$ GeV/c

$\Phi \rightarrow K^+K^-$ Invariant Mass

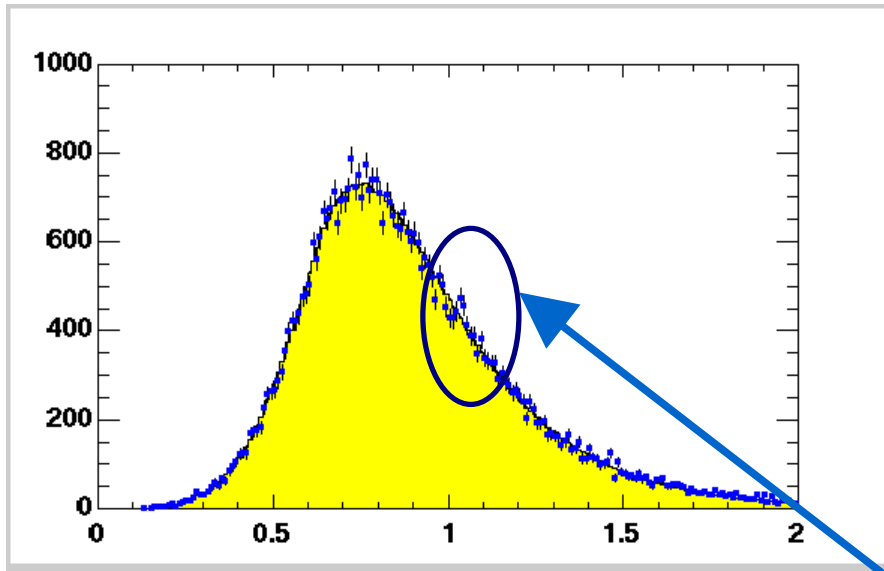
Au + Au $\sqrt{s_{NN}} = 200$ GeV

Centrality 40 –80 %



- $S = 218 \pm 24$
- $S/B = 1/2$

$\Phi \rightarrow e^+e^-$ Mass Spectrum



mass [GeV/c²]

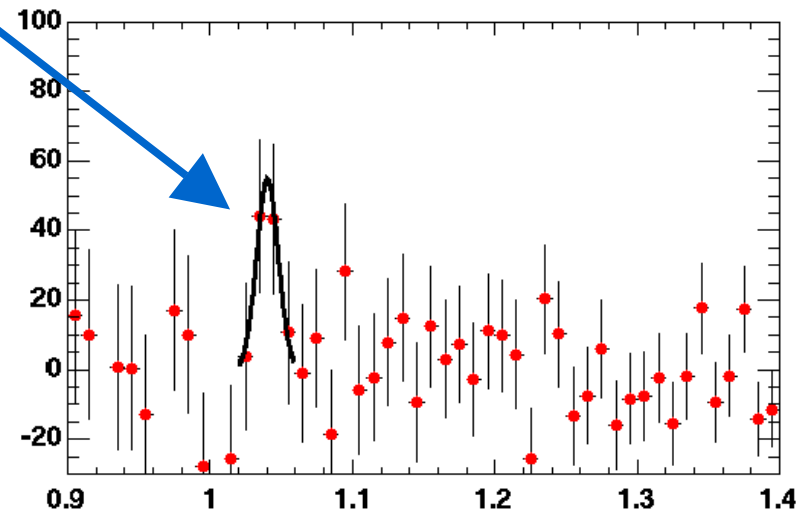
$$S = 101 \pm 47 \text{ (stat)}_{-20}^{+56} \text{ (syst)}$$

$$S/B = 1/20$$

Preliminary
PHENIX

Au + Au $\sqrt{s_{NN}} = 200$ GeV

Minimum bias



mass [GeV/c²]

Minimum bias dN/dy

- $\Phi \rightarrow K^+K^-$ (TOF)

$$\frac{dN}{dy} = 2.01 \pm 0.22^{+1.01}_{-0.52}$$

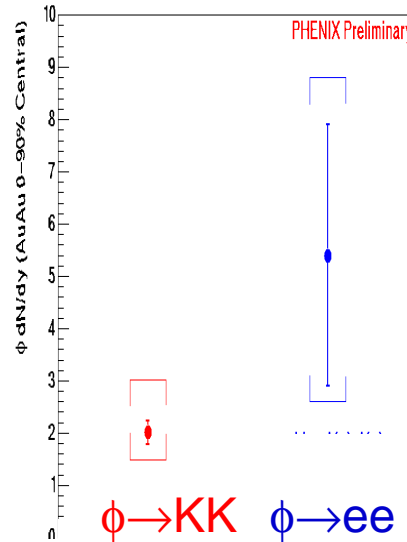
stat syst

- **only** probe at RHIC for chiral symmetry restoration (until PHENIX upgrade)
- STAR & PHENIX can (in principle) measure both channels
- **requires high statistics, high precision measurement**

- $\Phi \rightarrow e^+e^-$

$$\frac{dN}{dy} = 5.4 \pm 2.5^{+3.4}_{-2.8}$$

stat syst

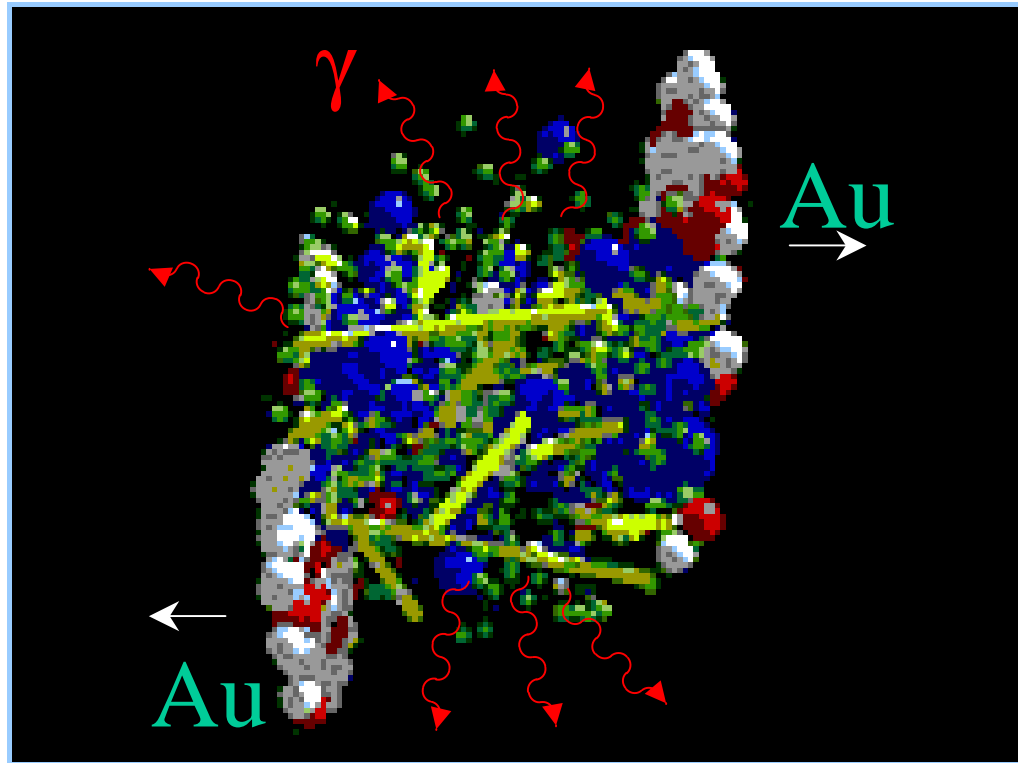


Data are consistent with free vacuum PDG branching fraction values within 1σ statistical errors.

Preliminary



Why are photons interesting?



- To Study Hot Dense Matter

Long mean free path in dense matter – much larger than the transverse size of matter created in Heavy Ion Collisions

- Probe the Partonic Stages of Heavy Ion Collisions

Partonic Interactions - Compton ($qg \rightarrow q\gamma$) and Annihilation ($q\bar{q} \rightarrow g\gamma$, $q\bar{q} \rightarrow \gamma\gamma$)

Electromagnetic Bremsstrahlung ($q \rightarrow q\gamma$)

How do we measure photons?

By reconstructing photon conversions

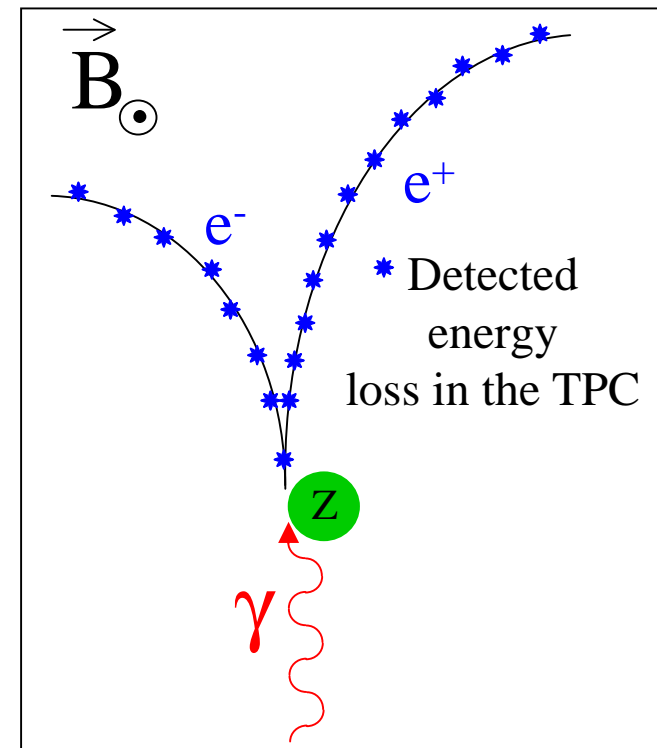
$$-\ \gamma Z \rightarrow e^+ e^- Z$$

TPC was used as a Pair Spectrometer

- low efficiency (1%)
- but large acceptance
- excellent energy resolution

$$\Delta p_t/p_t \sim 2\% \text{ at } p_t=0.5\text{GeV}/c$$

$$\Delta p_t/p_t \sim 4\% \text{ at } p_t=3.0\text{GeV}/c$$



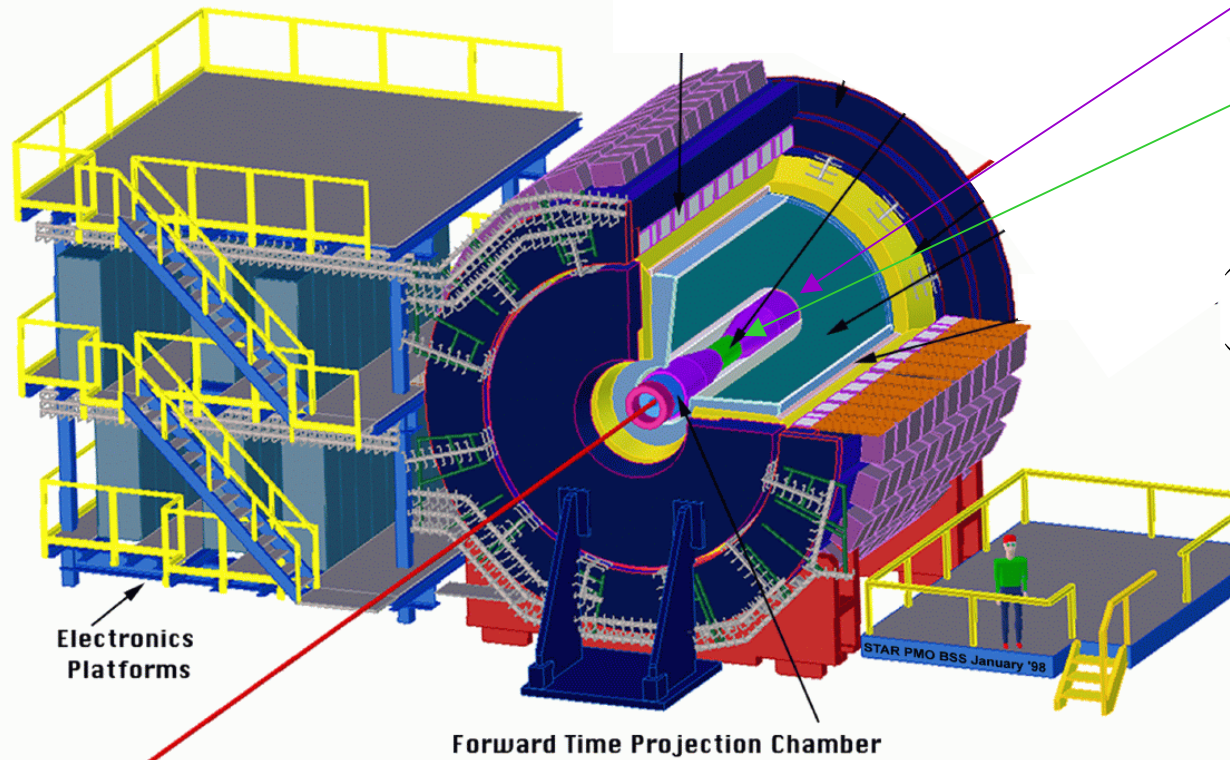
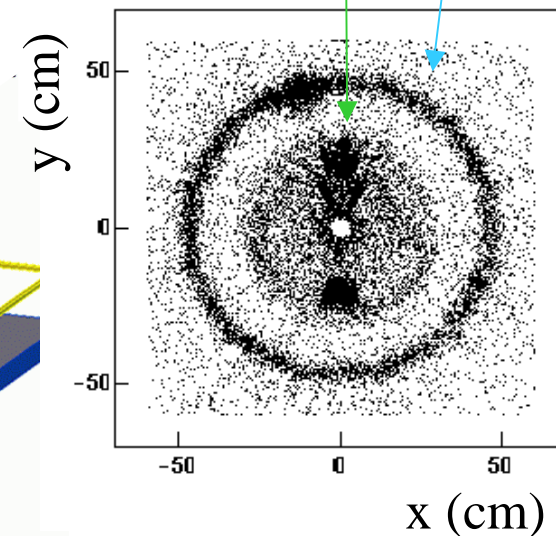
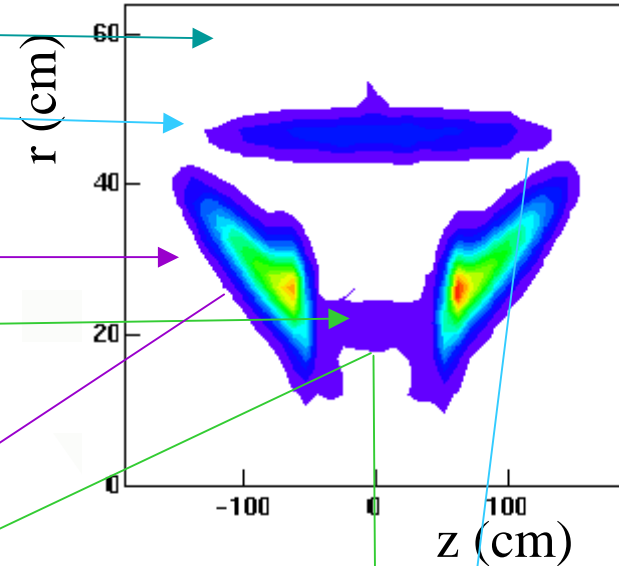
Converters in STAR

TPC

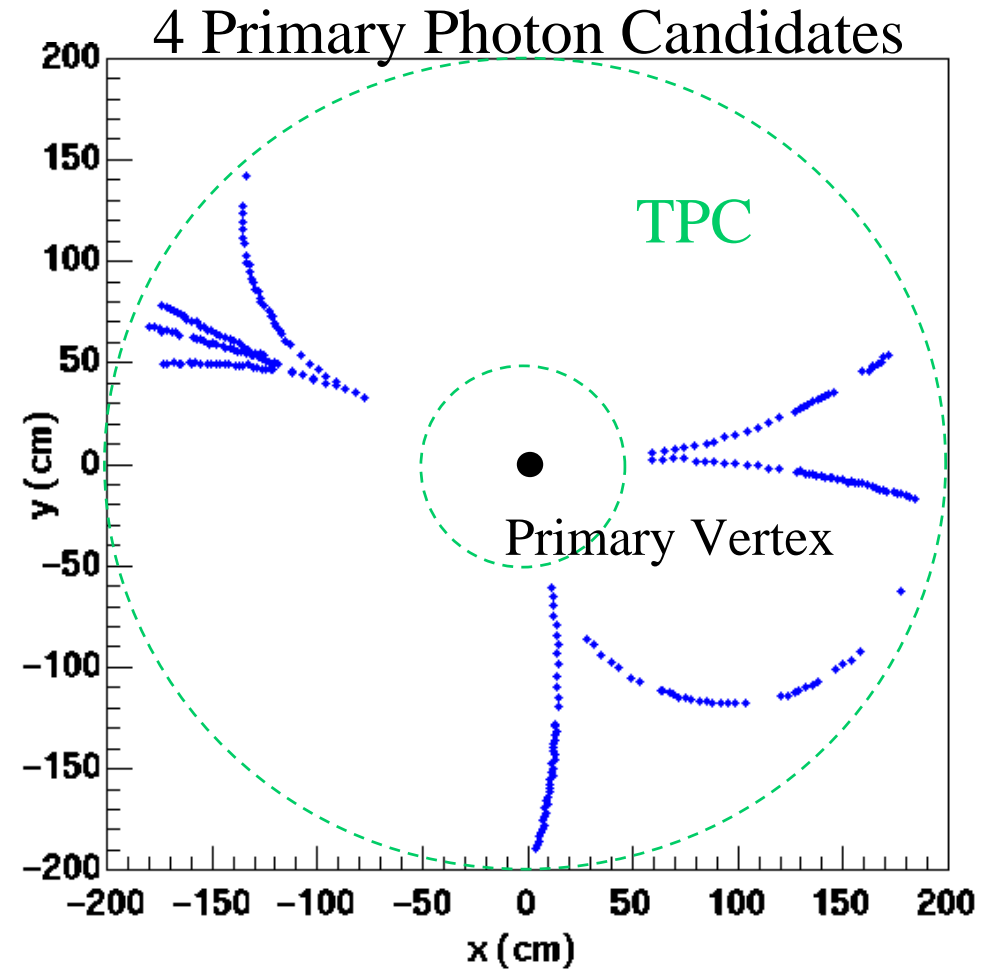
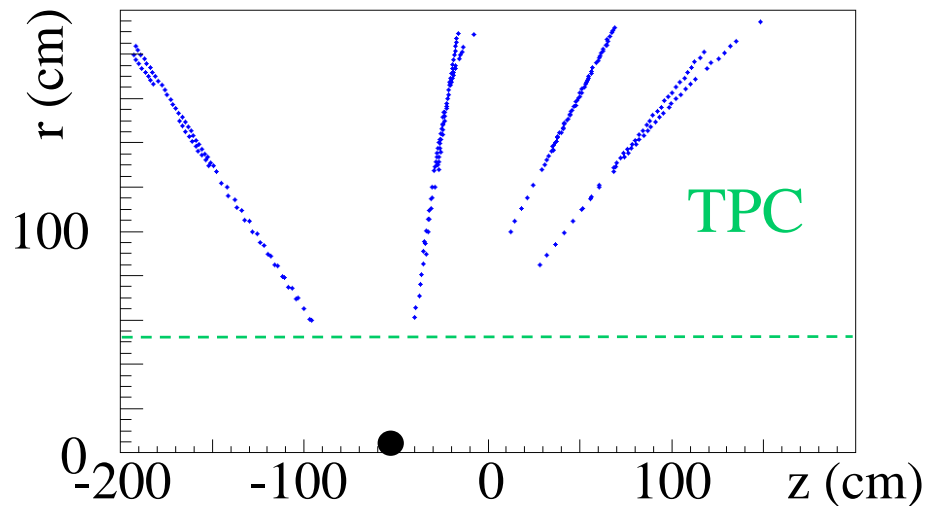
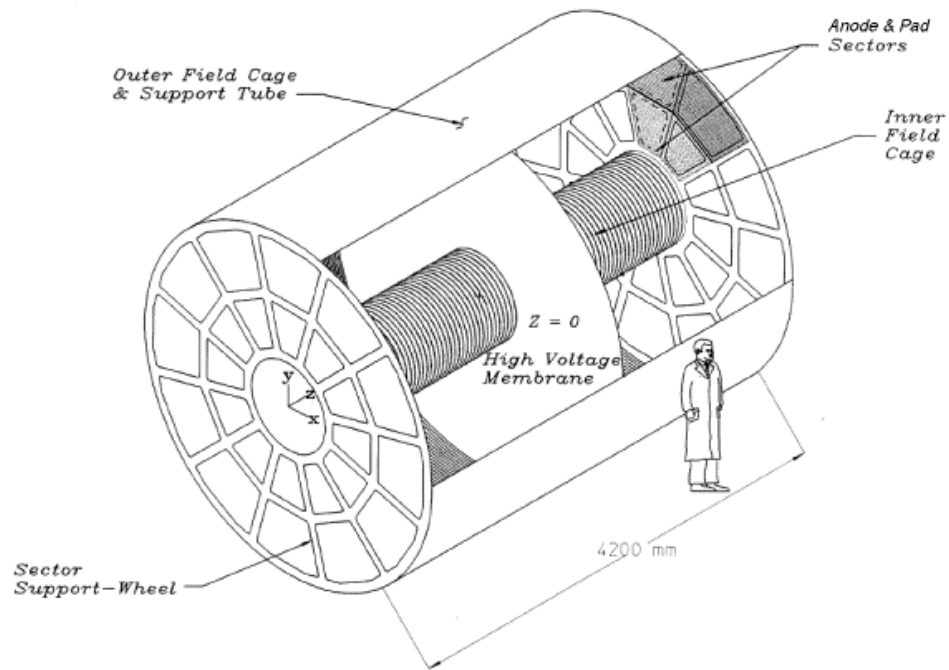
- gas
- inner field cage

Silicon Vertex Tracker (SVT)

- support cones
- support rods and silicon ladder

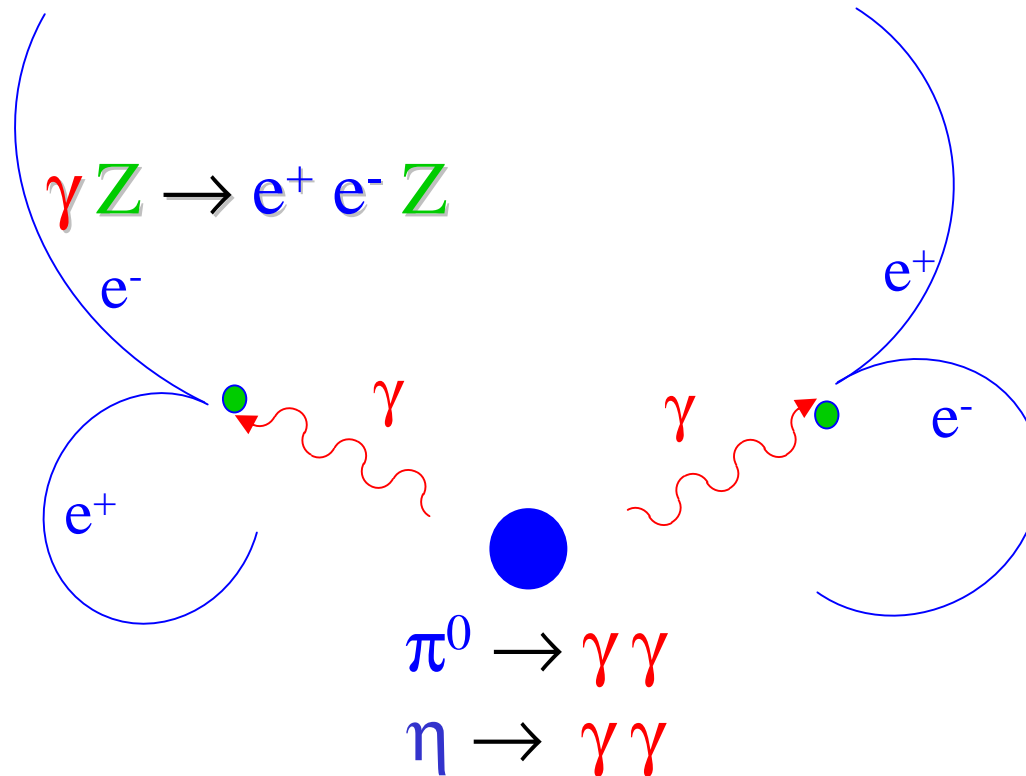


A Typical STAR Event



Note: Most tracks are not shown

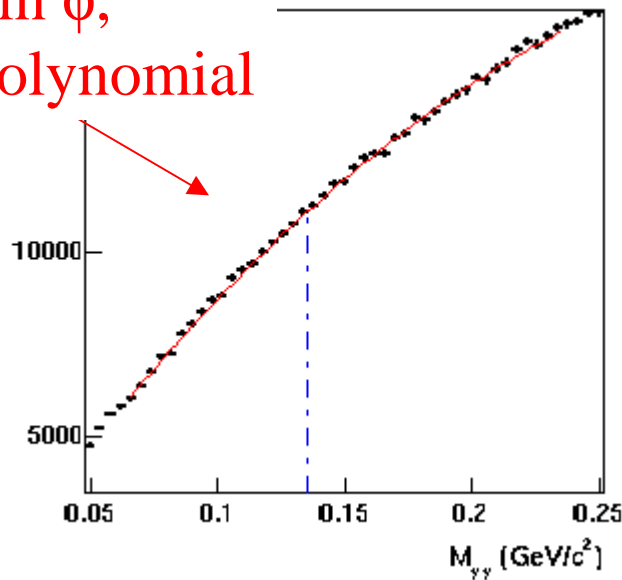
Two Photon Decays



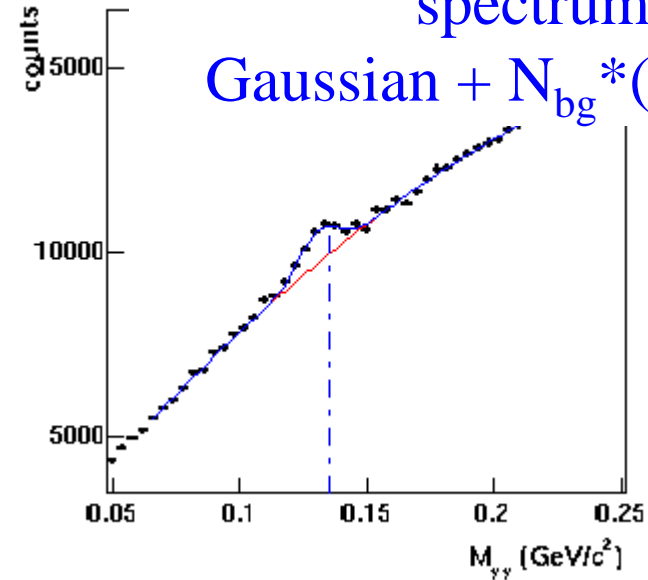
- $\pi^0 \rightarrow \gamma\gamma$ Branching Ratio 98.80 %
- $\gamma Z \rightarrow e^+ e^- Z$ Conversion Probability $\sim 1\%$
- e^+ and e^- Tracking Efficiency 60 - 90%
- Overall π^0 Reconstruction probability $\sim 10^{-4}$

Extracting $\pi^0 \rightarrow \gamma\gamma$ Yields

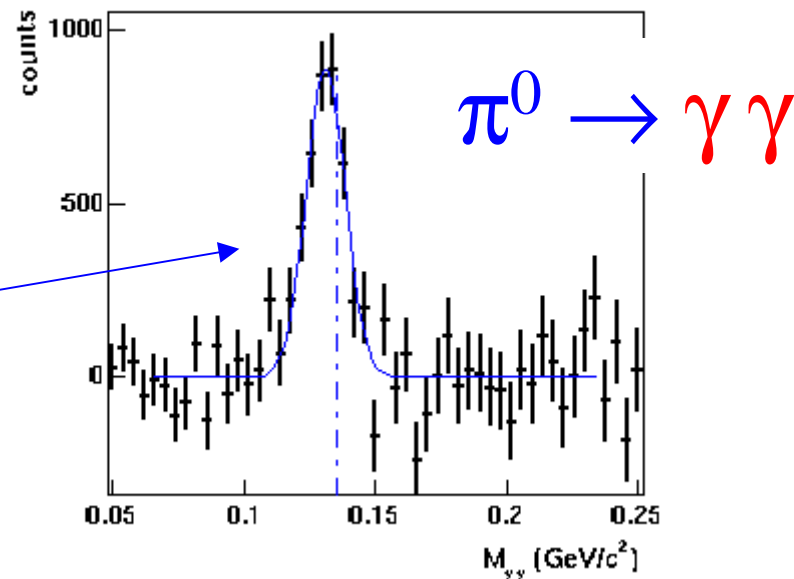
One photon rotated
by π in ϕ ,
2nd order polynomial



Two photon invariant mass
spectrum,
Gaussian + $N_{bg} * (2^{nd} \text{ poly})$

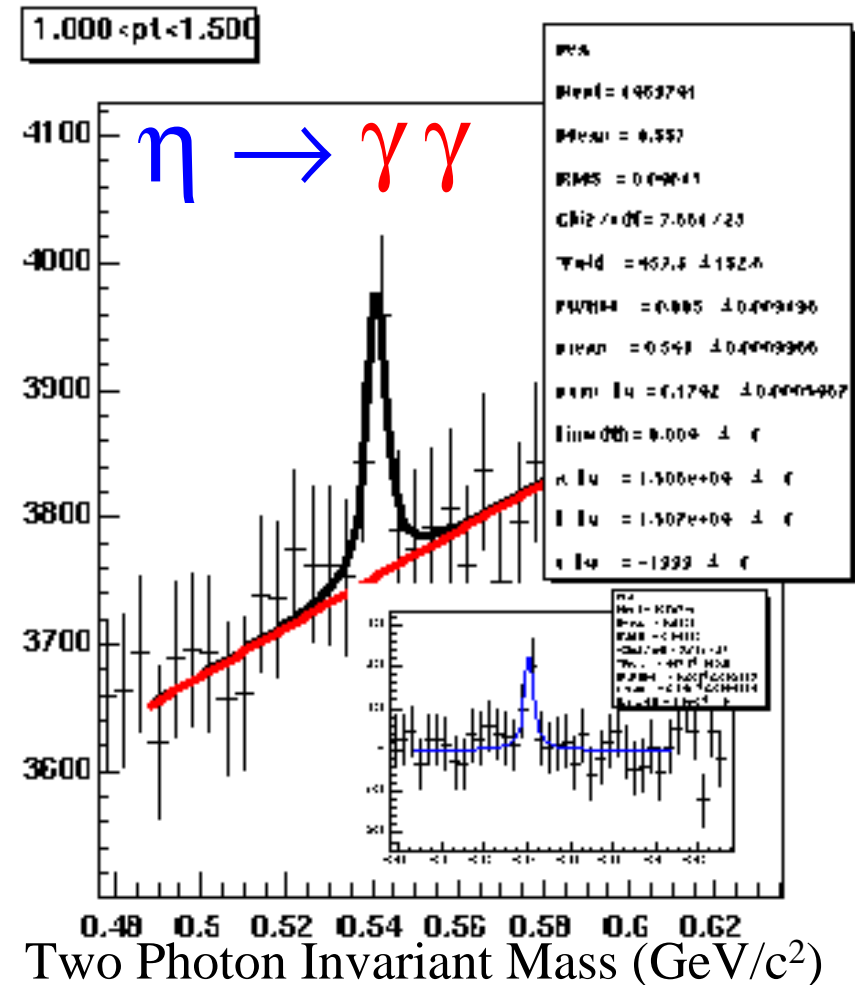


After background subtraction,
Gaussian ($\sigma \sim 10 \text{ MeV}/c^2$)



Outlook

- Measure the η Cross Section
 - $\eta \rightarrow \gamma\gamma$ was observed! (2000)
- With photon, π^0 and η measurements
 - extract a cross section for direct photon production at $\sqrt{s_{NN}} = 200$ GeV



Conclusions



We are excited
about the first
results.

We are working
on systematic
errors.

