Radiative electron capture in the first forbidden unique decay of ⁸¹Kr

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The photon spectrum accompanying the orbital K-electron capture in the first forbidden unique decay of 81 Kr was measured. The total radiation intensity for the photon energies larger than 50 keV was found to be $1.47(6) \times 10^{-4}$ per K-capture. Both the shape of the spectrum and its intensity relative to the ordinary, non-radiative capture rate, are compared to theoretical predictions. The best agreement is found for the recently developed model which employs the length gauge for the electromagnetic field.

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I. INTRODUCTION

Radiative electron capture (REC) is a process in which one of orbital electrons is captured by the atomic nucleus and in addition to an electron neutrino a photon is emitted [1]. Because of the three-body character of the process, the energy spectrum of these photons is continuous, reaching up to the maximum value of the decay energy Q_{EC} reduced by the binding energy of the captured electron in the daughter atom, B(nl). Such a radiative process occurs with the probability of the order of 10^{-4} with respect to the ordinary, radiation-less electron capture.

Early studies have shown, for numerous decays involving allowed nuclear transitions (nuclear spin changes by $\Delta J = 0, 1$ with no parity change), that the intensity and shape of the REC spectrum can be well understood by assuming that radiation is emitted by a captured electron (internal bremsstrahlung, IB) [1]. However, a measurement of REC in case of the forbidden nuclear transition in 41 Ca [2] revealed a strong disagreement with existing models. Moreover, the most advanced theoretical model of IB, developed by Zon and Rapoport for the arbitrary degree of forbiddenness [3, 4], showed the largest deviation from the experiment. The EC decay of ⁴¹Ca belongs to a category of first-forbidden unique (1u) transitions $(\Delta J = 2, \pi_i \pi_f = -1)$. Unique transitions are of special interest, because in the probability ratio of radiative to non-radiative decays, the nuclear matrix elements cancel out and the process should be governed solely by electromagnetic interactions. For ⁴¹Ca, the total probability of the REC process, per non-radiative decay, was found to be 6 times *larger* than the detailed prediction of the Zon and Rapoport model [2].

To explain this finding, a hypothesis of the so called detour transitions was considered by Kalinowski [5, 6] following an idea of Ref. [7]. According to it, a large part of radiation is emitted by a nucleus in addition to the IB mechanism. It was argued that such a nuclear contribution is particularly significant in case of 1u decays, where the combination of an allowed $GT \to C$ decay with a nuclear E1 gamma transition, may compete with the direct nuclear 1*u* transition accompanied by an IB photon. Apparently, for the case of ⁴¹Ca this hypothesis could fully account for the missing intensity found by the experiment [5, 6]. Surprisingly, in another case of the 1u REC — the decay of 204 Tl [8] — an opposite situation was observed. The intensity of the measured REC spectrum was found to be *smaller* by a factor of 4 than the predictions of the Zon and Rapoport model. Thus, no room for detour transitions was left, since their contribution is always positive in the model of Kalinowski.

To solve this conundrum, Pachucki et al. have undertaken a new approach to description of the REC process [9]. The main ingredient of the proposed model was the description of the electromagnetic field in the *length* gauge [10] in contrast to the Coulomb gauge used in all previous calculations. The key point is that although the predictions are gauge-invariant, the particular *length* gauge is strongly preferred for the actual calculation in this case. One reason is a suppression of the nuclear contributions, which allows to neglect the detour transitions. The new model was found to agree very well with experimental spectra of ⁴¹Ca and ²⁰⁴Tl [9].

In this paper, we report on the REC measurements for the third case of 1u transitions — the decay of ⁸¹Kr. This nucleus, with an intermediate mass and atomic number, located between ⁴¹Ca and ²⁰⁴Tl, suits very well as a testing case for the theory of radiative electron capture accompanying forbidden decays. In addition, this is the last case known where the radiation accompanying a firstforbidden unique (1*u*) decay can be measured for the pure ground-state-to-ground-state transition, which makes it

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experimentally feasible.

1. Decay properties of ⁸¹Kr

The $7/2^+$ ground state of ⁸¹Kr decays with the probability of almost 100 % to the $3/2^-$ ground state of ⁸¹Br with the half-life of 2.3×10^5 y [11]. This transition is of the first-forbidden unique type. The decay energy Q_{EC} amounts to (280.8 ± 0.5) keV [11], so that the decay can proceed only by an electron capture process. A small branch of the ⁸¹Kr decay feeds the first excited state of ⁸¹Br at 276 keV with the spin/parity of $5/2^{-}$. The branching for this first forbidden non-unique transition was measured by Axelsson et al. to be $3.0(2) \times 10^{-3}$ [12]. However, the decay energy for this weak branch equals only 4.7 keV and is much smaller than the K-electron binding energy in a bromine atom which amounts to $B_K = 13.47$ keV. Therefore, the transitions to the excited state in ⁸¹Br proceed only by electron captures from higher shells. In turn, the K-electron capture can occur only in a ground-state to ground-state transition. Thus, by a condition of coincidences with Br KX-rays one can select pure 1u K-capture transitions between ground states of ⁸¹Kr and ⁸¹Br. The maximum energy available for REC photons in the K-capture is given by $q_K = Q_{EC} - B_K$, which leads to $q_K = 267.3$ keV.

II. EXPERIMENTAL TECHNIQUE

A. ⁸¹Kr source

The activity of ⁸¹Kr was collected at CERN-ISOLDE [13], using a proton beam of 1 GeV energy that induced spallation reactions in a ⁹³Nb production target. The reaction products were ionized in a surface ion-source, accelerated to a kinetic energy of 50 keV and subsequently mass separated by means of the GPS separator tuned for the transmission of the particles with A = 81. The selected products were implanted in an aluminum catcher foil of 10 mg/cm² thickness. Krypton ions, although produced in the reaction, are not ionized and extracted from the source with any significant efficiency. However, rubidium is readily ionized with high yield and it was thus primarily ⁸¹Rb ions that were implanted, with some additional component of ⁸¹Sr. These implanted ions, having the half-life of 4.58 h and 22.2 m respectively, eventually decayed to ⁸¹Kr forming the source of interest.

To avoid sputtering of already implanted atoms by the impinging beam, the position of the foil was changed. After the irradiation, the foil was left in the chamber for a few days to let the short-lived precursor activity decay out. The final source material was collected on two foils in two irradiation sessions, in 1998 and in 2000, yielding six active spots, each having the area of a few mm squared. Two longer-lived contaminants were found in the source, stemming from incomplete suppression of neighbouring isobars in the mass separator: 82 Sr ($T_{1/2} = 22.5$ days) and 83 Rb ($T_{1/2} = 86.2$ days). Their gamma ray activity, in particular Kr- KX rays, makes the 81 Kr REC measurements very difficult. Instead of undertaking any chemical purification of the source, we decided to wait few years until these contaminants have decayed. The activity of 81 Kr and of the contaminants was monitored. The gamma and KX spectra of collected samples, measured in 2001 and 2003, are shown in Figs. 1 and 2.



FIG. 1: Singles gamma-ray spectra of the 81 Kr source taken in November 2001 (upper panel) and in March 2003 (lower panel). Gamma transitions following decays of contaminant 83 Rb atoms are seen as well as a 276 keV line from a decay of 81 Kr to the $5/2^{-}$ excited state in 81 Br.

Finally, in 2005 the influence of 83 Rb was found to be negligible. The active spots were carefully cut out from the catcher foils and placed on a perspex disk 1 mm thick and 2 cm in diameter and covered tightly with a thin mylar foil. The number of 81 Kr atoms in the final source was measured to be $4.1(2) \times 10^{15}$.

B. Photon spectrometry

Measurements were performed in the Institute of Experimental Physics at the Physics Department of Univer-



FIG. 2: Singles KX-ray spectra of the ⁸¹Kr source taken in November 2001 (upper panel) and in March 2003 (lower panel). In the upper spectrum contaminant Kr KX-rays originating from decays of ⁸³Rb can be seen.

sity of Warsaw. To select the K-capture component, the REC spectrum was measured in coincidence with the Br-KX rays. REC photons were recorded with an ORTEC GMX 45% Ge detector. The X-rays were measured by means of an ORTEC LOAX spectrometer. Both detectors had 0.5 mm thick beryllium windows.

The source was mounted between the two detectors in a close face-to-face geometry, with the surface covered with mylar foil directed towards the X-ray detector. The whole setup was placed inside a lead shielding of 10 cm thickness. The inside of the shielding was covered with 1.2 mm thick cadmium sheets and 8 mm thick copper plates.

A single DGF-4C CAMAC module [14] was used to process electronic signals from both detectors. The preamplifier outputs were directly connected to the two inputs of the DGF-4C module. Each input signal was digitized with a 40 MHz frequency followed by a real-time digital signal processing which included time stamping. Output data were read out by a PC computer and stored on a hard disk. In the off-line analysis, the coincidence relationships between signals could be restored with help of the time stamps. In addition, a 1 Hz pulser was connected to the test inputs of both preamplifiers and to a free running scaler. Comparison of the scaler readings with the number of counts recorded by the DGF-4C module allowed to estimate and to monitor the dead time of the acquisition system.

The long coincidence sessions (1-2 days) were alternated with short (ca. 10 min.) singles measurements necessary for the KX-rays intensity monitoring.

Two series of measurements were carried out. In the first one, in 2005, the total time of γ -KX coincidence runs amounted to 925.7 h (\simeq 38.6 days). The average rate of singles Br-KX rays was determined to be 107.7(5) counts/s. After all runs, the ⁸¹Kr source was removed and the background was measured in the same coincidence conditions for 113 h.

In the second series of measurements, in 2007/2008, the γ -KX coincidences were collected for total of 78 days. The average singles intensity of Br-KX rays over this period was found to be 95.1(5) counts/s. This time, however, for reasons explained in the following, the ⁸¹Kr runs were alternated with background γ -KX coincidences. The total time of background coincidence measurements was 71 days. The total time of the second experiment amounted to 190 days.

C. Calibrations and corrections

The calibrations of both Ge detectors, including the efficiency calibration of the GMX detector, were performed with help of standard sources: 57 Co, 60 Co, 137 Cs, 152 Eu, 203 Hg, and 241 Am. In order to avoid summing effects, the full-energy peak efficiency of the GMX detector was measured in a far geometry with all sources and in the close geometry, corresponding to the 81 Kr source position, only with 241 Am, 203 Hg, and 137 Cs, in which no summing occurs. Then, the far-geometry curve was scaled to the close-geometry position based on the efficiency ratios determined with the latter sources. The finally determined efficiency of the GMX detector at 100 keV amounted to about 14% for both runs of 81 Kr measurements.

The performance of the coincidence circuitry was tested with a ¹³³Ba source measured in a far geometry to avoid summing effects. From the decay scheme of ¹³³Ba the probabilities of γ -line emission per K-capture were determined for the lines at 81 keV, 276 keV, 302 keV, 356 keV, and 384 keV. These values were compared with the measured γ -KX coincidences yielding the coincidence efficiency consistent with 100% within error bars of a few percent for all tested energies.

Since the REC spectrum of photons is continuous, it has to be corrected for the Compton scattering. To estimate the magnitude and shape of the corresponding correction, the response of the GMX detector to monoenergetic radiation had to be determined. First, the response was measured with help of sources: ²⁴¹Am, ⁵⁷Co, ²⁰³Hg, and ¹³⁷Cs. Then, the spectra were compared with Monte-Carlo simulations performed with the GEANT package [15]. The effective dimensions of the detection set-up were adjusted to obtain a satisfactory agreement between the measured and simulated spectra. This allowed to calculate the detector response to the radiation of arbitrary energy within the range of interest.

D. Background

Selection of the K-component of the REC spectrum requires coincidences with KX-rays of the daughter atom. Usually, such condition practically removes any background contribution to the measured gamma spectrum. However, in case of ⁸¹Kr, the energy values of Br-KX rays, 11.9 keV and 13.3 keV, unfortunately happen to overlap with LX-ray energies of heavy elements, like protactinium and thorium. Since some radioactive isotopes of these elements belong to natural radioactivity chains, traces of them appear in the surroundings, mainly in lead bricks. Coincidence conditions result in a selection of those gamma rays from their decays which are coincident with transitions strongly converted on the L-shell. An example of background gamma spectrum, measured without the ⁸¹Kr source but with the same coincidence conditions, in particular with a gate on Br-KX energy range in the LOAX detector, is shown in Fig. 3.



FIG. 3: The background gamma spectrum gated by radiation in the Br-KX ray energy range detected in the LOAX detector. The total measurement time was 70 days.

The strongest γ lines present in Fig. 3 are identified as emitted by members of uranium-radium and actinium natural radioactivity chains, starting from ²³⁸U and ²³⁵U, respectively. For example, a strong line at 63.3 keV is emitted from ²³⁴Pa after the β^- decay of ²³⁴Th (itself a decay product of ²³⁸U). This line is seen because of coincidence with an E2 transition of 29.5 keV which has a large L-shell conversion coefficient in protactinium. Another L-converted transition of 20.0 keV (M1+E2) in ²⁴³Pa, is in coincidence with the 92.4 keV gamma line which can also be seen in the spectrum. Two other strong lines, at 143.8 keV and 185.7 keV, correspond to transitions between states in ²³¹Th fed by α decay of ²³⁵U. Both are coincident with a 19.6 keV M1+E2 transition which is strongly L-converted.

The contribution of natural background to the REC spectrum was identified after the first series of measurements when the coincidence run without the 81 Kr source was taken only for 113 h. Therefore, in the second series, the coincident background runs were alternated with REC runs, so that the background spectrum with much higher statistics was accumulated.

III. RESULTS

A. REC spectrum

The K-component of the REC spectrum of ⁸¹Kr was determined separately from both measurement series by applying the following procedure. First, the spectrum of time differences between γ -ray and KX-ray events was created. The peak in this spectrum allows to discriminate the true KX- γ coincident events from a flat background representing random coincidences. Then, by appropriate gating on the X-ray coordinate to select Br-KX rays and on the spectrum of time differences, the spectrum of γ rays coincident with Br-KX rays was constructed for all runs measured with the ⁸¹Kr source. Exactly the same gating procedure was used to determine the γ spectrum of background coincidences, discussed in the previous section. After normalization, the background spectrum was subtracted from the former one yielding the raw REC spectrum of ⁸¹Kr accompanying the K-shell electron capture. As an example, the spectra from the second experiment are shown in Fig. 4. The total number of counts in the REC spectrum is 16800, which corresponds to about 9 coincident events per hour.

In the next step, the correction for Compton scattering is introduced. Calculated contributions of each γ -energy bin are subtracted one by one, starting from the highenergy end (a peeling-off method). This correction is largest at low energy where contributions from all higherenergy bins add up, and amount to about 30%. The corrected REC spectrum is shown in Figure 4 by points with the error bars.

Finally, the absolute normalization of the REC spectrum is made. Each bin of the Compton-corrected REC spectrum is divided by the corresponding full-energypeak efficiency of the GMX detector and by the total number of Br-KX rays recorded by the LOAX detector during the all ⁸¹Kr coincidence runs. This number was determined from the singles KX-ray measurements and includes corrections for the dead-time in both the singles and coincidence runs. Such a procedure yields the probability distribution of the REC photon energy normalized to ordinary, non-radiative K-capture rate. This final REC energy spectrum can be directly compared to theoretical predictions. The resulting distributions from two series of measurements are presented in Figure 5.





FIG. 4: (Color online) Gamma energy spectrum coincident with Br-KX rays from the second series of measurements. The upper histogram was extracted from runs with ⁸¹Kr source. The contribution of background lines can be seen. The lower histogram shows the result of background subtraction. The circles filled in red illustrate the result of Compton correction and represent the raw K-component of the ⁸¹Kr REC spectrum.

B. Theoretical models

In a very general way, the probability of radiative electron capture from the 1S state (K-capture) in which a photon in the energy range (k, k + dk) is emitted, per ordinary, non-radiative K-capture can be written as [1]:

$$\frac{dw_K^{REC}(k)}{w_K} = \frac{\alpha}{\pi (m_e c^2)^2} \frac{k(q_K - k)^2}{q_K^2} R_K(k) \, dk, \qquad (1)$$

where α is the fine structure constant, $m_e c^2$ is the electron rest energy, q_K is the photon end-point energy, and the dimensionless function $R_K(k)$ is the shape factor of the spectrum. In the simplest Coulomb-free, non-relativistic approximation, valid for allowed nuclear transitions $R_K(k) \equiv 1$ [16].

In a more appropriate theoretical approach which takes into account both relativistic effects and the influence of the Coulomb field, and which is valid for first-forbidden unique nuclear transition, the shape factor is given by:

$$R_K^{1u}(k) = \left(1 - \frac{k}{q_K}\right)^2 R_K^{(1)}(k) + \left(\frac{k}{q_K}\right)^2 R_K^{(2)}(k), \quad (2)$$

where the functions $R_K^{(1)}$ and $R_K^{(2)}$ describe the effects of the Coulomb interaction between the nucleus and the radiating electron. Usually, these functions have to be calculated numerically for each specific case.

The first advanced description of REC in forbidden transitions was provided by Zon and Rapoport [3, 4], who

FIG. 5: (Color online) The intensity of the K-component of the REC spectrum of ⁸¹Kr per ordinary, non-radiative K-capture and per 1 keV. Open and closed circles represent the first and the second series of measurement, respectively. The solid line shows the prediction of the model by Pachucki et al. [9], while its Coulomb-free limit is represented by the dashed line. The dash-dotted line illustrates the prediction of the Zon and Rapoport model [3, 4].

extended a framework for allowed decays, developed previously by Glauber and Martin [17, 18], and generalized it to nuclear transitions of any order of forbiddenness. In this model both functions, $R_K^{(1)}$ and $R_K^{(2)}$, approach unity in the Coulomb-free limit $(Z \to 0)$. The model of Zon and Rapoport was found to fully reproduce the earlier results of Glauber and Martin for allowed decays which were quite well confirmed experimentally [1]. However, both the full and the Coulomb-free versions of this model failed to describe the REC spectrum measured in case of 1u decays in 41 Ca and 204 Tl [2, 8].

This observation motivated Pachucki et al. to reconsider the problem of REC in forbidden decays following a different approach [9]. The important conclusion from that work was that, although final results must not depend on a particular choice of gauge for the electromagnetic field, the Coulomb gauge applied by previous authors has rather unfortunate consequences when, as exemplified in the work of Zon and Rapoport, the approximations may later result in diverging terms. In addition, in the Coulomb gauge one has to include contributions from nuclear degrees of freedom. In turned out that the adoption of a different gauge for the electromagnetic field, so called *length* gauge, simplifies calculations considerably and avoids, in fact, the difficulties mentioned before. In particular, it was demonstrated that in the *length* gauge the contribution form the nuclear (detour) transitions can be neglected [9].

The results of Pachucki et al. confirmed all previous predictions for allowed transitions as well as the detailed form of the function $R_K^{(1)}$, appearing in eq. (2). However, a different form was derived for the second function — $R_K^{(2)}$, which is the one affected by a diverging term in the model of Zon and Rapoport. In particular, the Coulombfree limit for this function was found to be [9]:

$$\lim_{Z \to 0} R_K^{(2)}(k) = 1 + \frac{m_e c^2}{k} + 2\left(\frac{m_e c^2}{k}\right)^2, \qquad (3)$$

which obviously differs from 1.

The model of Pachucki et al. was found to be in almost perfect agreement with measured spectra in ⁴¹Ca and in ²⁰⁴Tl [9]. Its prediction for the present case of ⁸¹Kr is plotted in Figure 5 with the solid line. The result of the Coulomb-free limit and of the Zon and Rapoport model are also shown.

C. 1S shape factor

The shape of the REC spectrum, contributions of different terms, and comparison with data can be shown in detail by plotting the dimensionless shape factor of the spectrum, $R_K(k)$. The experimental shape factors were extracted from the determined REC spectra with help of Eq. (1). The values obtained are shown in Figure 6 which contains also theoretical predictions.

As seen in Figs. 5 and 6, there is a systematical difference between results obtained from two experiments. The collected statistics in the first series was smaller which is partly reflected in larger errors. In addition, the correction for background was less accurate in the first experiment, which could result in a systematical shift. Since the investigated effect is small and the rate of coincidences was very low, a proper control of possible systematical effects was difficult. We note however, that each spectrum taken separately would lead to the same conclusion. The prediction of Pachucki et al. reproduces quite well both experimental spectra, in contrast to the model by Zon and Rapoport.

D. Total intensity

The total probability of the K-component of the REC spectrum per non-radiative K-capture, calculated by integrating the spectrum from 50 keV to the end-point, is given in Table 1 together with values obtained from theoretical models. The weighted average of the total intensity from two measurements equals to $1.47(6) \cdot 10^{-4}$ per K-capture.

The value of total intensity from both measurements, as well as their average, are closest to the prediction of Pachucki et al. It is interesting to note the large difference between the Coulomb-free limits of both theories. In case of Pachucki et al., this approximation overestimates the spectrum intensity. The inclusion of Coulomb and relativistic effects reduces the intensity to the value



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FIG. 6: (Color online) The shape factor for the K-component of the first forbidden unique REC in ⁸¹Kr. Open and closed circles represent the first and the second series of measurement, respectively. The meaning of theoretical lines is the same as in Fig. 5, while the additional dotted line shows the Coulomb-free version of the Zon and Rapoport model.

150

Energy [keV]

200

250

100

50

TABLE I: The total probability for the K-component of REC per K-capture in 81 Kr, integrated from 50 keV to the spectrum end-point, in units of 10^{-4} obtained from the first (a) and the second (b) experiment. The results for a full calculation, as well as the Coulomb-free version (CF), are given for both theoretical models.

Experiment	Pachucki et al.		Zon and Rapoport	
	Full	\mathbf{CF}	Full	\mathbf{CF}
a) 1.2(1)	1.37	4.00	0.60	0.26
b) 1.64(8)				

which almost perfectly agrees with the experiment. The model of Zon and Rapoport behaves in the opposite way. The Coulomb-free limit lies below the full calculation, which in turn is still much lower than the experimental values.

IV. CONCLUSIONS

We have measured the spectrum of photons accompanying the first-forbidden unique electron capture decay of ⁸¹Kr. By detecting γ -rays in coincidence with Br-KX rays we have selected the pure K-component of radiative electron capture (REC) taking place between ground states of ⁸¹Kr and ⁸¹Br. From the measured spectra we have extracted the absolute probability of REC emission per normal, non-radiative K-electron capture as a function of photon energy, as well as the shape factor of the photon spectrum. Both the shape and intensity of the REC spectrum are found to be well reproduced by the recent theoretical model of Pachucki et al. [9], while the strong disagreement is found with predictions of the old model by Zon and Rapoport [3, 4]. Thus, ⁸¹Kr is the third case displaying REC in a 1u nuclear transition, in addition to ⁴¹Ca and ²⁰⁴Tl, which strongly supports the approach taken by Pachucki et al. In contrast, the model of Zon and Rapoport fails to describe correctly the probability and shape of the REC spectrum in all these three cases.

The important feature of the first-forbidden unique decays is that they are governed predominantly by a single nuclear matrix element. In the probability ratio of the radiative capture to the ordinary, non-radiative one, which is determined in experiment, this nuclear matrix element cancels out. Thus, the result should not depend on detailed knowledge of nuclear wave functions and should be fully determined by the electro-weak sector alone. An important consequence is that the theoretical models discussed do not contain any adjustable parameters. The predictions are fully determined by the mass and atomic numbers of the decaying nucleus and by the maximal photon energy, related to the decay energy Q_{EC} .

The main difference between the two theoretical models discussed is the selected gauge of electromagnetic field. While the Zon and Rapoport use the *Coulomb* gauge, Pachucki et al. apply the *length* gauge. Our conclusion is that the latter is preferred. The reason is that calculations in the *length* gauge are technically simpler and do not require additional approximations. Also, Zon and Rapoport introduced approximations which were not valid and led to a diverging term which was found to be

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responsible for the final failure in comparison with the experiment. Moreover, in the *length* gauge the contribution from nuclear degrees of freedom to the emitted radiation can be shown to be negligible, while this is not true in case of the Coulomb gauge.

The 1*u* decays of 41 Ca, 81 Kr, and 204 Tl are the only known where the pure transition between ground states can be selected. Thus, only in such cases, the weak radiative branch can be determined with sufficient accuracy. Additional tests of this alternative description of REC are in principle possible with transitions of second-forbidden non-unique type (2nu: $\Delta J = 2, \pi_i \pi_f = +1$). To this class belong ⁵⁹Ni and ¹³⁷La in which K-component of REC spectrum was measured [19, 20]. The results obtained for ⁵⁹Ni disagree with Zon and Rapoport predictions, while the spectrum for ¹³⁷La could not be shown to contradict them. However, in case of 2nu decays, the REC spectrum depends on one additional parameter which represents a ratio of nuclear matrix elements. Since these matrix elements are not known, and are difficult to estimate, this parameter adds an extra degree of freedom. In consequence, the comparison between an experiment and the theory is not as unambiguous as in case of 1u transitions. Nevertheless, in the future we plan to compare the measured spectra of ⁵⁹Ni and ¹³⁷La with predictions of the new model of Pachucki et al.

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